

QUANTUM MECHANICS

(1st quantisation, classical)

Scalar
(or sometimes vector)

$\Psi(x, t)$

|4> Dirac notation

- * Fundamentals * State of a particle is determined by its wavefunction
- * Probability of finding particle in $d^3x = |\Psi|^2 d^3x$
- * Flux (# particles crossing unit area/unit time) = $-\frac{ie}{2m} (\Psi^* \nabla \Psi - \Psi \nabla \Psi^*)$
- * Classical observables \leftrightarrow Hermitian operators in QM.

Position $\rightarrow \hat{x} = x$	Momentum $\rightarrow \hat{p} = -ie\nabla$
Energy $\rightarrow \hat{E} = ie\frac{\partial}{\partial t}$	Angular momentum $\rightarrow \hat{L}_z = ie(x\frac{\partial}{\partial y} - y\frac{\partial}{\partial x})$ (z component)
classical: $L = \sum p_i x_i$	$= ie\frac{\partial}{\partial \theta} \leftarrow$ cartesian spherical polar.
- * Most operators don't commute.
c.f. conjugate variables classically.
e.g. $[\hat{x}, \hat{p}_x] = ie$ * Possible outcomes of an observation are the (recall) Eigenvalues of the operator - correspond to specific eigenstates of the operator. $\hat{A}|\Psi\rangle = a_n|\Psi_n\rangle$
- * If two operators commute \rightarrow can have common eigenstates \Rightarrow measured simultaneously \rightarrow uncertainty principle for operators \hat{A}, \hat{B}

$$\Delta A \Delta B \geq \frac{1}{2} | \langle i [\hat{A}, \hat{B}] \rangle | \quad \Delta A = \sqrt{\langle \hat{A}^2 \rangle - \langle \hat{A} \rangle^2}$$
- * Any state Ψ can be expressed as a superposition of all the eigenstates of any operator i.e. if $\hat{A}|\Psi_n\rangle = a_n|\Psi_n\rangle \quad \Psi = \sum_n b_n |\Psi_n\rangle$.
- * If a system is described by state Ψ , measure using operator \hat{A}
 - \rightarrow probability of measuring eigenstate a_n is $|b_n|^2$ or $|\langle \Psi_n | \Psi \rangle|^2$. ONLY IF eigenstates Ψ_n are orthonormal. $\int \Psi_i^* \Psi_j dx \equiv \langle \Psi_i | \Psi_j \rangle = \delta_{ij}$
 - \rightarrow After measurement system adopts state Ψ_n if a_n is measured.
- * Find sets of wavefunctions by $\hat{A}|\Psi_n\rangle = a_n|\Psi_n\rangle$ equation. If $\{E_n\}$ are energy eigenvalues $\{|\Psi_n\rangle\}$ eigenstates are found via Schrödinger Equation. $\hat{H}|\Psi_n\rangle = E_n|\Psi_n\rangle$
E = total energy. Hamiltonian \hat{H} is formed from classical statement of energy.
classical: $E = T + V$ θ M: $\hat{H} = \frac{\hat{p}^2}{2m} + \hat{V}$. Now Ψ_n is not usually time independent, so using Hermitian operators above: so analytical/numerical/approximate
- \rightarrow Time dependent S.E. $-\frac{\hbar^2}{2m} \nabla^2 \Psi + V\Psi = ie\frac{\partial \Psi}{\partial t} \rightarrow$ solution for Ψ possible if V is known.
g for multiple particle systems.
- * usually eigenstates Ψ of energy (well of \hat{H}) are time separable. i.e. $\Psi(x, t) = \phi(x)T(t)$
 \Rightarrow can utilise fact that energy eigenstates have well defined energy eigenvalues E_n to form time independent S.E. $\hat{H}|\phi_n(x)\rangle = E_n|\phi_n(x)\rangle$ (cancel $T(t)$). $\Psi(x, t)$ must be separable in time?
cancelling $\phi_n(x)$ we use time dependence of $E_n|\Psi_n\rangle$ to find $T(t)$.
- * $i\hbar \frac{d}{dt} E_n = E_n \dot{T}_n \Rightarrow T_n(t) = e^{-iE_n t/\hbar}$ so $\Psi_n(x, t) = \phi_n(x) \exp(-iE_n t/\hbar)$. \therefore only need to solve time independent S.E. to find $E_n, \phi_n(x)$ to know $\Psi_n(x, t)$. Note $|\Psi(x, t)|^2 = |\phi(x)|^2$ "stationary states"
- * observables have following time dependence $\frac{d}{dt} \langle \hat{A} \rangle = i/\hbar \langle [\hat{H}, \hat{A}] \rangle$
For any stationary state $\langle \hat{A} \rangle = \text{const.} \Rightarrow$ If $[\hat{H}, \hat{A}] = 0 \Rightarrow \langle \hat{A} \rangle$ is a constant of the system, conserved.
- * Symmetry of $\hat{H} \Rightarrow$ conservation laws.
- * Note can use superposition principle and time dependence above to calculate time dependence of any state Ψ . (May not be an eigenstate Ψ etc....)

Invariance under \leftrightarrow

Spatial Translations \leftrightarrow

Temporal " \leftrightarrow

Rotations \leftrightarrow

Parity ($+ \rightarrow -$) \leftrightarrow

Conservation of

Momentum

Energy

Angular Momentum

Parity

2) Angular momentum and spin * S.E. for spherically symmetric potential $V(r)$
 → Solutions of form $\psi(r, \theta, \phi) = R_{nl}(r) Y_{lm}(\theta, \phi)$. i.e., states characterised by integer quantum numbers n, l, m . Total energy only depends on n . → For hydrogen atom $\psi(r) \propto \frac{1}{r}$, $E_n \propto -\frac{1}{n^2}$. Now given n , permitted $l = 0 \dots n-1 \rightarrow$ given l , permitted $\epsilon \in \mathbb{Z}^+ \neq 0$ $m = -l, -l+1 \dots l$ comes from solving S.E.

* Note form of $V(r)$ only determines $R_{nl}(r)$ NOT $Y_{lm}(\theta, \phi)$ 2l+1 possible values

* Important properties of angular momentum in QM can be derived from commutation relationships. $[\hat{l}_x, \hat{l}_y] = i\hbar \epsilon_{xyz} \hat{l}_z$ * All components of \hat{l} commute with $\hat{l}^2 \equiv \hat{l}_x^2 + \hat{l}_y^2 + \hat{l}_z^2 \Rightarrow$ can measure \hat{l}^2 and at most one component of \hat{l} . → convention \hat{l}_z .

* From S.E. → eigenstates of \hat{l}_z, \hat{l}^2 are: $\hat{l}_z^2 |Y_{lm}\rangle = l(l+1) \hbar^2 |Y_{lm}\rangle$ $\hat{l}_z |Y_{lm}\rangle = m\hbar |Y_{lm}\rangle$

* Ladder operators transform between eigenstates of \hat{l}_z $\hat{l}_{\pm} = \hat{l}_x \pm i\hat{l}_y$ \leftarrow Not observables
 If $\hat{l}_z |\phi_m\rangle = m\hbar |\phi_m\rangle \rightarrow$ can show using commutation relations $\hat{l}_z (\hat{l}_{\pm} |\phi_m\rangle) = (m \pm 1) \hbar (\hat{l}_{\pm} |\phi_m\rangle)$
 $\Rightarrow \hat{l}_{\pm} |\phi_m\rangle = \pm \hbar [l(l+1) - m(m \pm 1)]^{1/2} |\phi_{m \pm 1}\rangle$ (correctly normalised).

* Dirac equation (S.E. + relativistic considerations) predicts 'intrinsic angular momentum or SPIN' in addition to 'orbital' angular momentum L . Many particles have $\frac{1}{2}$ integer spin e.g., electron $s = \frac{1}{2}$. → Modify wavefunction for electron in $V(r)$ potential. $\psi(r, \theta, \phi) = R_{nl}(r) Y_{lm}(\theta, \phi) \chi_{sm}$ in non-relativistic limit. (i.e. "patch" to S.E.)

* S is fixed by particle type but find χ_{sm} in some way for L .

If $[\hat{s}_x, \hat{s}_y] = i\hbar \epsilon_{xyz} \hat{s}_z$; $\hat{s}_{\pm} = \hat{s}_x \pm i\hat{s}_y$; $\hat{s}_{\pm} |\chi_{sm}\rangle = \pm \hbar [s(s+1) - m_s(m_s \pm 1)]^{1/2} |\chi_{sm \pm 1}\rangle$
 $\hat{s}_z^2 |\chi_{sm}\rangle = \hbar^2 s(s+1) |\chi_{sm}\rangle$; $\hat{s}_z |\chi_{sm}\rangle = \pm m_s |\chi_{sm}\rangle$; GIVEN s , $m_s = -s, -s+1 \dots s$ $\uparrow 2s+1$ values

* Spin algebra has handy isomorphism with Pauli matrices. (In fact many formulation of QM is sometimes very useful). Only for $s = \frac{1}{2}$ though....

→ If $s = \frac{1}{2}$; $\hat{s}_x = \pm \frac{1}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$; $\hat{s}_y = \pm \frac{1}{2} \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$; $\hat{s}_z = \pm \frac{1}{2} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$ → Eigenstates are $\chi_+ = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$
 $\chi_- = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$

3) Combination of angular momentum \leftarrow Non trivial because L, S are vectors with QM restrictions on their orientation.

* Define total angular momentum operator

$\hat{\Sigma} = \hat{L} + \hat{S}$. Quantum numbers j, m_j $m_j = -j, -j+1 \dots j$

Properties: $[\hat{s}_x, \hat{s}_y] = i\hbar \epsilon_{xyz} \hat{s}_z$; $\hat{s}_z^2 |\phi_{jm_j}\rangle = j(j+1) \hbar^2 |\phi_{jm_j}\rangle$; $\hat{s}_z |\phi_{jm_j}\rangle = \pm m_j \hbar |\phi_{jm_j}\rangle$ $\uparrow 2j+1$ values

$\hat{s}_{\pm} = \hat{s}_x \pm i\hat{s}_y$ ($= \hat{l}_{\pm} + \hat{s}_{\pm}$ e.g.); $\hat{s}_{\pm} |\phi_{jm_j}\rangle = \pm \hbar [j(j+1) - m_j(m_j \pm 1)]^{1/2} |\phi_{jm_j}\rangle$ as for L, S

* Note $j = l+s, l+s-1 \dots |l-s|$ for $\hat{\Sigma} = \hat{L} + \hat{S}$

* $\hat{l}^2, \hat{s}^2, \hat{l}_z, \hat{s}_z$ mutually commute - so do $\hat{\Sigma}^2, \hat{l}^2, \hat{s}^2, \hat{l}_z, \hat{s}_z$ BUT NOT TOGETHER. can! characterise angular momentum eigenstates by $|\phi_{jm_j}\rangle$ or combinations of $|Y_{lm}\rangle, |\chi_{sm}\rangle$

→ LS, jj coupling see later!

* For $\hat{\Sigma} = \hat{L} + \hat{S}$ combinations use CLEBSCH-GORDON algorithm/method to find wavefunctions
 (i) start with largest $m_j = l+s$, $j = l+s$ state (ii) apply \hat{s}_- to find other $j = l+s$ states
 (iii) use orthogonality to find the state $j = l+s-1$ $m_j = l+s-1$ (iv) apply \hat{s}_- on this state to find other $j = l+s-1$ etc repeat until all states are found.

4) Identical particles * Symmetry depends on spin of particle → Integer spin - 'BOSONS'

w.t. symmetric under particle interchange $\rightarrow \frac{1}{2}$ integer spin - 'FERMIOS' w.t. anti-sym " " "
 BOSONS can have same quantum state Fermions can't. (w.t. $\rightarrow 0$) PAULI EXCLUSION PRINCIPLE // QM ②

Approximation methods in QM - S.E. too complicated for analytical solution in most cases.

Variational method (i) Need to solve $\hat{H}\psi_n = E_n \psi_n$ $E_0 \leq E_1 \leq E_2 \dots$ (ii) consider arbitrary function ϕ . (iii) Expand in terms of ψ_n $\phi = \sum c_n \psi_n$.
vi) consider expectation value of \hat{H} in ϕ : $\langle \phi | \hat{H} | \phi \rangle = \sum_m \sum_n c_m^* c_n \langle \psi_m | \hat{H} | \psi_n \rangle = \sum_m \sum_n c_m^* c_n E_n \delta_{mn}$
v) Define ψ_{n3} to be orthogonal s.t. $\langle \psi_m | \psi_{n3} \rangle = \delta_{mn}$ $= \sum_n |c_n|^2 E_n$. (vi) Also consider $\langle \phi | \hat{H} | \phi \rangle = \sum_m \sum_n c_m^* c_n \langle \psi_m | \psi_n \rangle = \sum_n |c_n|^2$. Now clearly $\sum_n |c_n|^2 E_n \geq \sum_n |c_n|^2 E_0$
from (ii). Now $\sum_n |c_n|^2 E_0 = E_0 \langle \phi | \phi \rangle$ so since $\sum_n |c_n|^2 E_n = \langle \phi | \hat{H} | \phi \rangle$
 $\Rightarrow \frac{\langle \phi | \hat{H} | \phi \rangle}{\langle \phi | \phi \rangle} \geq E_0$ VARIATIONAL THEOREM So use TRIAL function ϕ to guess upper bound to E_0 .

Raleigh-Ritz method (i) Guess ϕ from available insight/prior knowledge. Write in terms of adjustable parameters $\alpha_1, \alpha_2, \alpha_3 \dots$ i.e. $\phi(r, \alpha_1, \alpha_2, \dots)$
(ii) minimise $\frac{\langle \phi | \hat{H} | \phi \rangle}{\langle \phi | \phi \rangle}$ w.r.t. $\alpha_1, \alpha_2 \dots$ i.e. solve $\nabla(\dots) = 0$ using Cartesian product space of $\alpha_1, \alpha_2, \alpha_3 \dots$ substitute these values into (...) to compute 'best guess' to E_0 . ✓ Exited states
+ once ϕ_0 is found which minimises $\frac{\langle \phi | \hat{H} | \phi \rangle}{\langle \phi | \phi \rangle} \rightarrow$ construct orthogonal state ϕ' i.e. of form $\phi' = \phi - \langle \phi_0 | \phi \rangle \phi_0 \rightarrow$ re apply Raleigh Ritz. Note this calculation will not guarantee to $\geq E_1$ (only $\geq E_0$ guess). only works for certain if $c_0 = 0$ from symmetry, ϕ_0 is known to be exact.

Time independent, non-degenerate, perturbation theory - write $\hat{H} = \hat{H}_0 + \lambda \hat{H}'$
(λ is a bookkeeping device - usually set = 1). \hat{H}_0 = solvable Hamiltonian, \hat{H}' = small perturbation (e.g. effect of small E field). KNOWN solution to $\hat{H}_0 \psi_n = E_n \psi_n$. look for ϕ_n s.t.
 $(\hat{H}_0 + \lambda \hat{H}') \phi_n = E_n \phi_n$ (i) Expand ϕ_n, E_n in powers of λ : $\phi_n = A(\psi_n + \lambda \sum_{j \neq n} c_j^{(1)} \psi_j + \lambda^2 \sum_{j \neq n} c_j^{(2)} \psi_j + O(\lambda^3))$
 $E_n = E_n^0 + \lambda E_n^{(1)} + \lambda^2 E_n^{(2)} + O(\lambda^3)$. subst. into (i):
 $O(\lambda)$ (1st order) $\Rightarrow \hat{H}' \psi_n = \sum_{j \neq n} c_j^{(1)} (E_n^0 - E_j^0) \psi_j + E_n^{(1)} \psi_n \quad \langle \psi_n | \Rightarrow E_n^{(1)} = \langle \psi_n | \hat{H}' | \psi_n \rangle$
 $\langle \psi_k |$ (2nd) $\Rightarrow \langle \psi_k | \hat{H}' | \psi_n \rangle = c_k^{(1)} (E_n^0 - E_k^0) \Rightarrow c_k^{(1)} = \frac{\langle \psi_k | \hat{H}' | \psi_n \rangle}{E_n^0 - E_k^0}$
Note non-degenerate: $E_n^0 \neq E_k^0$.

For $O(\lambda^2)$ terms use above technique to find: $E_n^{(2)} = \sum_{j \neq n} \frac{|\langle \psi_n | \hat{H}' | \psi_j \rangle|^2}{E_n^0 - E_j^0}$ (for example).

Time independent, degenerate perturbation theory Suppose we have a set of n states which are degenerate w.r.t. \hat{H}_0 , i.e. $\hat{H}_0 \psi_i = E_0 \psi_i$ $i = 1, 2, \dots, n$. Any linear combination also has energy E_0 . i.e. state $\chi_K = \sum_{i=1}^n d_i \psi_i$ (define weight coefficient d_i). Wtchr $\langle \chi_i | \hat{H} | \chi_j \rangle = \delta_{ij}$ so $\langle \chi_i | \hat{H} | \chi_j \rangle \neq 0$ ($= 0$ invalid) \rightarrow Need to DIAGONALISE \hat{H}' matrix. Procedure: $(\hat{H}_0 + \hat{H}') \chi_K = (E_0 + E_K^{(1)}) \chi_K$ from 1st order perturbation ($E_K^{(1)} = \langle \chi_K | \hat{H}' | \chi_K \rangle$). $\Rightarrow (\langle \psi_i |) \sum_i d_i \langle \psi_i | \hat{H}' | \psi_i \rangle = E_K^{(1)} d_i$
 $\Rightarrow \sum_i d_i h_{ji} = d_i E_K^{(1)}$ MATRIX EIGENVALUE PROBLEM ($\hat{H}' \underline{d} = E_K^{(1)} \underline{d}$) \rightarrow solve $\begin{vmatrix} \underline{H}' & \underline{d} \\ \underline{d} & E_K^{(1)} \end{vmatrix} = 0$ to find 1st order shift. Apply same-ish technique for higher order terms. EXAMPLE - STARK EFFECT. Note degenerate states of partly drop up - i.e. H abm (partly $\sim 1^4$).

6) Transitions - Bound States (solutions to $\hat{H}_0 \psi = E\psi$) have time dependence $e^{-iEt/\hbar}$
 $\Rightarrow E\psi$ time independent. Apply time dependent perturbation $\hat{H}'(t)$ - in general likely
 to get transitions between states. Examples: (1) Atomic/molecular transitions (2) scattering (ΔE)

* 2 State System ψ_1, ψ_2 $\hat{H}_0\psi_1 = E_1\psi_1$; $\hat{H}_0\psi_2 = E_2\psi_2$ let $E_1 = \hbar\omega_1$, $E_2 = \hbar\omega_2$; Apply perturbation
 $\hat{H}'(t)$. Write $\psi(t) = c_1(t)e^{i\omega_1 t}\psi_1 + c_2(t)e^{i\omega_2 t}\psi_2$. Substitute into $\hat{H}\psi = i\hbar\frac{d\psi}{dt}$
 where $\hat{H} = \hat{H}_0 + \hat{H}'(t)$. Now if $\langle \psi_1 | \hat{H}' | \psi_2 \rangle = \langle \psi_2 | \hat{H}' | \psi_1 \rangle = 0$ and $\langle \psi_1 | \hat{H}' | \psi_2 \rangle = i\hbar\omega_1$

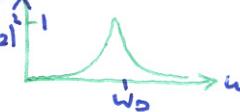
$$\langle \psi_1 | \hat{H}' | \psi_2 \rangle = i\hbar\omega_1 c_1 e^{i\omega_1 t} \psi_1 + i\hbar\omega_2 c_2 e^{i\omega_2 t} \psi_2 = i\hbar\omega_1 c_1 e^{i\omega_1 t} \psi_1$$

Now write $\omega_3 \equiv \omega_2 - \omega_1$ (Transition frequency), expand as $\cos\omega t = \frac{1}{2}(e^{i\omega t} + e^{-i\omega t})$

Neglect $e^{\pm i(\omega+\omega_3)t}$ terms (just oscillations) {i.e. save near resonance ($\omega \approx \omega_0 \gg |\omega - \omega_0|$)}

Eliminating $c_1 \Rightarrow i\dot{c}_2 + i(\omega - \omega_3)c_2 + \frac{1}{2}\omega^2 c_2 = 0$. $\{ \dot{c}_2 = -i\frac{1}{2}\omega c_2 e^{i(\omega-\omega_3)t} \}$. DAMPED SHM (ish).

$$\Rightarrow |c_2|^2 = \frac{\omega^2}{\omega^2 + (\omega - \omega_0)^2} \sin^2 \omega t / 2, \omega_0^2 = \omega_0^2 - \omega^2 + \omega^2 \cdot |c_2|^2$$



BASIS OF MAGNETIC RESONANCE.

* Time-Dependent Perturbation theory (General case g above!)

- Time indep. unperturbed \hat{H}_0
 $\hat{H}_0\psi_n = E_n\psi_n$ (set $E_n = \hbar\omega_n$) \rightarrow Apply perturbation $\hat{H}'(t)$ for $t > 0$ \rightarrow write $\psi(t) = \sum c_i(t)\psi_i e^{i\omega_i t}$

\rightarrow initial conditions $t=0$, $c_i = 1$, $c_j = 0$ for $j \neq i$ \rightarrow Sub. into $\hat{H}\psi = i\hbar\frac{d\psi}{dt}$ ($\hat{H} = \hat{H}_0 + \hat{H}'(t)$).

$\rightarrow \langle \psi_{ikl} | \text{ and use matrix relation } i\langle \psi_{kl} | \hat{H}' | \psi_j \rangle = H_{kj} \Rightarrow i_{kl} = \frac{1}{i\hbar} \sum_j c_j e^{i(\omega_k - \omega_l)t} H_{kj}(t)$

DIFFICULT TO SOLVE! Now IF $\hat{H}'(t)$ is WEAK $\Rightarrow c_0 \approx 1$; $|c_k| \ll 1 \forall k \neq 0$ (High prob. in init state).

\rightarrow most likely transition from g state. $\Rightarrow i_{kl} \propto \frac{1}{i\hbar} e^{i(\omega_k - \omega_l)t} H_{kg}(t)$ $t \gg \hbar$

\rightarrow let $\hat{H}'(t) = \hat{H}(t)e^{i\omega t}$, $t > 0 \Rightarrow c_{kl}(t) = H_{kg} \frac{1}{i\hbar} \int_0^t e^{i(\omega_k - \omega_g - \omega)t} dt$. write $\delta\omega = \omega_k - \omega_g - \omega$

$$\Rightarrow |c_{kl}(t)|^2 = 4|H_{kg}|^2 \frac{1}{\pi^2} \frac{\sin^2 \delta\omega t}{\delta\omega^2}$$



Now for probability of transition to a group of closely spaced states

\rightarrow Integrate over states. Introduce "density of states" $g(E_k)$, $E_k = \text{final state energy}$

$$\# \text{ states } E_k \rightarrow E_k + dE_k = g(E_k) dE_k = \frac{1}{\hbar} g(E_k) d\omega = \frac{1}{\hbar} g(E_k) d(\delta\omega)$$

\rightarrow Assume $g(E_k)$ invariant across sharp $\sin^2(\delta\omega t/2)$ peaks and $|H_{kg}|$ same for all states near ω \Rightarrow Transition probability = $4|H_{kg}|^2 \frac{1}{\pi^2} \int \left(\frac{\sin^2 \delta\omega t/2}{\delta\omega} g(E_k) \right) d(\delta\omega)$, let $\alpha = \delta\omega t/2$

$$\Rightarrow \Gamma = 4|H_{kg}|^2 g(E_k) \frac{1}{\alpha} \int_{-\infty}^{\infty} \left(\frac{\sin x}{x} \right)^2 dx. \text{ Now } \Rightarrow \Gamma = \text{const. } t. \therefore \text{Transition rate } \frac{d\Gamma}{dt}$$

$$\equiv \Gamma_{0 \rightarrow k} \Rightarrow \Gamma_{0 \rightarrow k} = \frac{2\pi}{\hbar} |H_{kg}|^2 g(E_k) \quad \text{FERMI'S GOLDEN RULE}$$

* Radiative Transitions in atoms and molecules - Apply above to transitions between atomic energy levels - induced by a perturbing oscillating EM field (photon).

- Photon in \hat{z} direction, polarized by \hat{z} in x direction. Photon: $\lambda \sim 500 \text{ nm} \gg$ atomic dimension

$\Rightarrow E_{\text{ex}} = E_0 e^{i(kz - \omega t)} \Rightarrow \hat{H}' = -E_0 q z e^{i(kz - \omega t)}$, $q = \text{electron charge}$. Photon: $\lambda \sim 500 \text{ nm} \gg$ atomic dimension

: kz has a value $\ll 1$ ($\sim 10^{-3}$ typically) across atom $\Rightarrow e^{i k z} = 1 + i k z + \dots \approx 1$. {where $\int H_{kg}$ is performed}

$\Rightarrow \hat{H}' = -E_0 q z e^{-i\omega t}$ ELECTRIC DIPOLE APPROXIMATION ($qz = \text{electric dipole operator } \hat{d}$)

$\rightarrow \hat{H}'$ into Fermi's golden rule \rightarrow Stimulated transitions - Also spontaneous transitions.

↓ consider gas of atoms in thermo eqn.
Atomic transition rate (2 levels $E_k > E_j$) $\frac{4\pi}{\hbar^3} \frac{1}{\text{atom}}$

Stimulated $j \rightarrow k = B_{jk} u(w)$ \rightarrow Balance rates: $n_k [A + B_{jk} u(w)] = n_j B_{jk} u(w)$ $n = \text{atoms/level}$

Stimulated $k \rightarrow j = B_{kj} u(w)$ \rightarrow terms eq: $n_j = \frac{e^{-E_j/k_B T}}{e^{E_k/k_B T}} = e^{-\hbar\omega/k_B T}$ Assume degeneracy = nil.

Spontaneous $k \rightarrow j = A$ $\Rightarrow A = [B_{kj} e^{-\hbar\omega/k_B T} - B_{jk}] u(w)$ Now $u(w) = \frac{\pi w^3}{\pi^2 c^3} \frac{1}{e^{\hbar\omega/k_B T} - 1}$

where $\hbar\omega = E_k - E_j$ (PLATTCK'S LAW) - But $A \neq A(T) \Rightarrow B_{jk} = B_{kj}$ Hence $A = B_{kj} \frac{\pi w^3}{\pi^2 c^3}$

$u(w) = \text{energy density/unit w}$ $\omega_{\text{rot}} = \frac{1}{2}(e^{i\omega t} + e^{-i\omega t}) \rightarrow$ Angular frequency $H' = -q E_0 x \cos \omega t$; $\text{Extensivity} = \frac{1}{2} E_0 E_0^2 \Rightarrow u(w) = \frac{1}{2} E_0 E_0^2 \text{tg}(E)$

Selection rules - Transition rates between atomic energy levels depend on $\langle k_l | l_j \rangle$ Electric dipole appear.

Note $|\langle k_l | l_j \rangle|^2 = |\langle H_{\text{dip}}^1 \rangle|^2 = \frac{1}{3} |\langle k_l | d_1 l_j \rangle|^2$ where $d = \pm \frac{1}{2}$. Some elements of $\langle k_l | l_j \rangle$ must be zero.

I- Electronic energy levels are eigenstates of parity. It has odd parity.
 $\Rightarrow \langle k_l | l_j \rangle = 0$ If $|k_l | l_j \rangle$ have same parity \Rightarrow parity must change in E.dipole transitions.

II- let $|k_l\rangle = \psi_k(r) X_{l_z}$, $|l_j\rangle = \psi_j(r) X_j$; X = spin w.f. $\therefore \langle k_l | l_j \rangle = \langle X_{k_z} | X_j \rangle \int \psi_k^* \psi_j d^3 r$
 \Rightarrow unless X_{k_z}, X_j identical. $\Rightarrow \Delta S = 0$

III- consider only one electron changing state in transition. consider $\langle k_l | z | l_j \rangle$. $z = r \cos \theta$
 $\psi_{k_z}(r)$ is of form $R(r) Y_{lm}(\theta, \phi)$. ϕ integral: $\int_0^\pi d\phi e^{-im_l \phi} e^{im_j \phi} = 2\pi \delta_{mj} \delta_{lk}$
 consider $\langle k_l | d_1 | l_j \rangle$ $d = r \sin \theta \cos \phi \Rightarrow \phi$ integral: $\int_0^\pi d\phi e^{-im_l \phi} \frac{1}{2} (e^{i\phi} + e^{-i\phi}) e^{im_j \phi} = 0$
 unless $m_j - m_l = \pm 1$ (similarly for y). $\Rightarrow \Delta M = \pm 1, 0 \Rightarrow \Delta l = \pm 1$. ($\Delta l = 0$
 not allowed since parity = $(-1)^l$ and must change). Note $E_z \leftrightarrow \Delta M = 0$ $E_{x,y} \leftrightarrow \Delta l = \pm 1$
 (c.f. Zeeman effect).

IV- Atoms not necessarily eigenstates of \hat{l}_z (spin orbit coupling) but are of $\hat{l}_z^2, \hat{\ell}_z^2$.
 \Rightarrow (See Atomic/nuclear notes) $\Delta S = \pm 1, 0$. Similar argument to above?

Note: if we include $i\hbar \omega$ term in \hat{H}' i.e. $\hat{H}' \approx -E_0 \sigma_x (1 + i\hbar \omega)$, $\langle k_l | x_z | l_j \rangle$ matrix elements appear. \Rightarrow even parity \Rightarrow No parity change. \Rightarrow "Electric Quadrupole" transitions. Magnetic transitions also exist/occur.

* Selection rules (2) - charged particle in 1D Harmonic oscillator w.f. $|n\rangle$ $E_n = (n + \frac{1}{2})\hbar\omega$
 Transitions dominated by $\langle n' | x | n \rangle$. Note ladder operators $\hat{a}^\pm = \sqrt{\frac{m\omega}{2\hbar}} \hat{x} \mp i \frac{\hat{p}}{\sqrt{2m\omega}}$
 Property: $a^+ |n\rangle \rightarrow |n+1\rangle$, $a^- |n\rangle \rightarrow |n-1\rangle$. Write $\hat{x} = \sqrt{\frac{\hbar}{2m\omega}} [a^+ + a^-]$
 $\Rightarrow \langle n' | x | n \rangle \propto \{ \langle n' | n+1 \rangle + \langle n' | n-1 \rangle \} = 0$ unless $n' = n \pm 1$. $\Rightarrow \Delta n = \pm 1$ Selection rule.
 (Electric dipole appear again).

* Scattering - can apply Fermi's golden rule to atomic scattering as long as scattering potential is weak

- Beam of particles. Normalize in a box volume a^3 . $\psi = A e^{ik \cdot r} \cdot I$. $|\psi|^2 = A^2 = \frac{1}{a^3}$
- Impinge on fixed potential $V(r)$. Final state $\psi' = A e^{ik' \cdot r'} \cdot I$
- Assume energy transfer negligible $\underline{k} \approx V(r)$ (Elastic collision \circlearrowleft massive). $\therefore |k'| \approx |k'|$.
- Apply FGR: Transition $k \rightarrow k' = \frac{2\pi}{a^3} |\langle \psi' | V(r) | \psi \rangle|^2 g(E_F)$ Fermi's Golden Rule
- Now $\langle \psi' | V(r) | \psi \rangle = \frac{1}{a^3} \int V(r) e^{iK \cdot r} d^3 r$ 3D Fourier Transform. c.f. Rutherford Diffraction. where $K = \underline{k} - \underline{k}'$ Note $K = 2k \sin \theta / a$
- For $g(E_F)$: Assume w.f. normalized in box size a - K values are $k_{3x} = 2\pi n_x / a$ etc...
 \therefore each state occupies volume $(\frac{2\pi}{a})^3$ g K -space. # states in solid angle $d\Omega$
 and in range $k \rightarrow k + dk = k^2 d\Omega dk / (\frac{2\pi}{a})^3 \Rightarrow g(k) = k^2 d\Omega (\frac{a}{2\pi})^3$
- Now $E_k = \frac{\hbar^2 k^2}{2m} \Rightarrow g(E_k) = g(k) \frac{dk}{dE_k} = \frac{mk dk}{\pi^2} (\frac{a}{2\pi})^3$ (g(E)dE = g(h)dh)

DEFINE "CROSS SECTION" $\sigma = \# \text{ PARTICLES SCATTERED/UNIT TIME / INCIDENT FLUX}$
 AND "DIFFERENTIAL CROSS SECTION" $\frac{d\sigma}{d\Omega} = \# \text{ SCATTERED/UNIT TIME INTO } d\Omega$

- Incident flux = $-\frac{i\hbar}{2m} (4\pi \sigma v - \sigma v^2)$ $v = \frac{a^3}{2} e^{i\hbar \omega}$ INCIDENT FLUX $\cdot dk$
- \Rightarrow Incident flux = $-\frac{i\hbar}{2m a^3} (i\hbar \omega k) = \frac{i\hbar k}{m} \cdot \frac{1}{a^3} (= \frac{p}{m} \cdot \frac{1}{a^3} = \sigma / a^3)$ particle velocity.
- $\Rightarrow \frac{d\sigma}{d\Omega} = \left(\frac{m}{2\pi \hbar^2} \right)^2 \left| \int V(r) e^{iK \cdot r} d^3 r \right|^2$ "Born Approximation"

\rightarrow For Coulomb potential get Rutherford scattering $\frac{d\sigma}{d\Omega} = \left[\frac{m^2 e^2}{8\pi \epsilon_0 p^2 \sin^2 \theta / 2} \right]^2$ QM (5)

7) Atomic Structure * Hydrogen atom - Solve SE. pr pure Coulomb potential. Solutions $R_{nl}(r)$ $\Psi_{nlm}(r)$

If $n_r = \#$ nodes of $R_{nl}(r)$ $n_r = n-l-1$. (n_l , l quantum #'s). $E_n = -\frac{R_{\infty}}{n^2}$ R_{∞} Rydberg constant.

$R_{\infty} = 13.6 \text{ eV} = \frac{e^2}{8\pi\epsilon_0 a_0} = \frac{1}{2} m_e c^2 \alpha^2$ $a_0 = \text{Bohr radius} = \frac{4\pi\epsilon_0 \hbar^2}{m_e e^2} \alpha$ "Fine structure constant" $= \frac{e^2}{4\pi\epsilon_0 \hbar c}$ $\approx \frac{1}{137}$

If nuclear charge is Ze : $E_n \rightarrow Z^2 E_n$, $a_0 \rightarrow a_0/Z$.

Conventional orbitals: $l=0, 1, 2, 3, 4, 5, \dots$ * \approx degenerate $2s, 2p$ and $3s, 3p, 3d$
 s, p, d, f, g, h for single electron atoms

* Not so pr multi-electron atoms - $E = E(n_l, l, m_l)$ not $E(n_l, \dots)$

* Real Hydrogen atom has a few corrections to above. (1) FINITE nuclear size
 - not Coulomb potential - Heavy atoms noticeable effect. (2) NUCLEAR MASS - use reduced mass m but need to take into account isotopes (3) NUCLEAR SPIN - Proton's magnetic moment interacts with electron $M_p = g_p m_N S$; $g_p \approx 5.6$ $\mu_N = \frac{e\hbar}{2m_N}$ (nuclear magneton). \Rightarrow "Hyperfine structure"
 (4) NUCLEAR SHAPE - Some non spherical i.e. deuterium - quadrupole moment \Rightarrow Not Coulomb potential
 (5) SPIN ORBIT INTERACTION - electron moves through E field of nucleus - Sees this as a B field. Energy depends on orientation of spin magnetic moment with this B .
 (6) RELATIVISTIC CORRECTIONS - $KE = (\vec{p}^2 c^2 + m^2 c^4)^{1/2} - m^2 c^2 \approx \frac{p^2}{2m} - \frac{p^4}{8m^3 c^2} \dots$ {well, spin orbit, relativistic ...}

* DIRAC EQUATION for electron in Coulomb potential solves many of these problems. {orbit, relativistic ...}

$\Rightarrow E = mc^2 \left[1 + \frac{\alpha^2}{[n-j-\frac{1}{2} + \sqrt{(j+\frac{1}{2})^2 - \frac{1}{4}}]^2} \right]^{-1/2}$. In non-relativistic limit - retaining leading terms of Dirac Hamiltonian yields:

$\hat{H} \approx \hat{H}_0 + \hat{H}' + \hat{H}_2' + \hat{H}_3'$

S.E. \hat{H}_0 RELATIVISTIC \hat{H}' SPIN ORBIT \hat{H}_2' DARWIN TERM \hat{H}_3'

DARWIN TERM HAS NO CLASSICAL INTERPRETATION - can derive spin orbit $\hat{H}_2' \propto$ by simple classical argument

* Spin orbit interaction \hat{H}_2'

$\Rightarrow U = \frac{e \hat{S} \cdot (\vec{p} \times \nabla \phi)}{mc^2}$

$= \frac{e}{m^2 c^2} \hat{S} \cdot (\vec{p} \times \vec{r} \cdot \frac{1}{r} \frac{d\phi}{dr}) = -\frac{e}{m^2 c^2} \frac{1}{r} \frac{d\phi}{dr} \hat{S} \cdot \vec{L}$ ($\vec{L} = -\vec{p} \times \vec{r}$) so $\frac{1}{2}$ out. "Thomas precession" can get this.

Energy = $-M \cdot B$ $M = \frac{e \hat{S}}{m}$ \nwarrow Magnetic moment of electron
 $B = \frac{e \times E}{c^2}$ \nwarrow Larmor motion at ω through E Now $E = -\nabla \phi$ so if Energy = U :
 Relativistic precession of coordinate frames

* Note DIRAC Eq. predicts $2p, j=\frac{1}{2}$ and $2s, j=\frac{1}{2}$ degenerate in energy. Experimentally not so. ("Lamb shift" $\sim 10\%$ of α) \Rightarrow QED - Coulomb field needs quantising.

* Multi-electron atoms - Base on solutions for Hydrogen but need to take into account e-e repulsion, nuclear screening. - connection between e-e repulsion and w.f. symmetry, exclusion principle. \rightarrow EXCHANGE INTERACTION. e.g. for ground state of He: 2+1s electrons.
 $\psi = \psi_{1s}(r_1) \psi_{1s}(r_2) \cdot \frac{1}{r_2} [\uparrow_1 \downarrow_2 - \downarrow_1 \uparrow_2]$. Antisymmetric. (C.f. - each e must have opposite spins). $\Rightarrow S=0$

entire case: (1) $\psi_S = \frac{1}{2} [\psi_{1s}(\uparrow_1) \psi_{2s}(\downarrow_2) + \psi_{2s}(\uparrow_1) \psi_{1s}(\downarrow_2)] [\uparrow_1 \downarrow_2 - \downarrow_1 \uparrow_2] \Rightarrow S=0$ 'singlet'
 (2) $\psi_T = \frac{1}{2} [\uparrow_1 \downarrow_2 - \downarrow_1 \uparrow_2] [\uparrow_1 \downarrow_2 + \downarrow_1 \uparrow_2] \Rightarrow S=1$ 'triplet'.

In Triplet case electrons further apart \Rightarrow lower energy. \therefore Coulomb energy depends on spin orientation of electrons. or $\uparrow_1 \uparrow_2$ or $\downarrow_1 \downarrow_2$

* Central field approximation $\hat{H} = \sum_i \left(-\frac{\hbar^2}{2m_e} \nabla_i^2 - \frac{Ze^2}{4\pi\epsilon_0 r_{ij}} \right) + \sum_{i,j} \frac{e^2}{4\pi\epsilon_0 r_{ij}}$; $r_{ij} = |\vec{r}_i - \vec{r}_j|$ Note contains sizeable central (spherically symmetric) component enclosed shells.

\Rightarrow write $\hat{H} = \hat{H}_0 + \hat{H}'$, treat it as a perturbation. $\hat{H}_0 = \sum_i \left(-\frac{\hbar^2}{2m_e} \nabla_i^2 - \frac{Ze^2}{4\pi\epsilon_0 r_i} \right)$ separable

$\hat{H}' = \sum_{i,j} \frac{e^2}{4\pi\epsilon_0 r_{ij}} - \sum_i U(r_i)$ - Tricky part is to find $U(r_i)$ - can only do when $\hat{H}_0 \psi = E \psi$ is solved.

* Self-consistent field approach (1) Guess $U(r_i) \rightarrow 0$ $r_i \rightarrow \infty$ (no screening) \rightarrow interpolate in $\frac{(Z-1)e^2}{4\pi\epsilon_0 r_i}$ $r_i \rightarrow 0$ (perfect screening) between.

(2) Solve $\left[-\frac{\hbar^2}{2m_e} \nabla^2 - \frac{Ze^2}{4\pi\epsilon_0 r_i} + U(r_i) \right] \psi_{nlm} = E \psi_{nlm}$ pr ψ_{nlm}

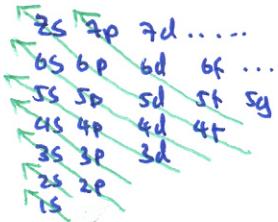
DETERMINANT $\psi = \frac{1}{\sqrt{N!}} \begin{vmatrix} \psi_{1s(1)} & \psi_{1s(2)} & \psi_{1s(3)} & \dots & \psi_{1s(N)} \\ \psi_{2s(1)} & \psi_{2s(2)} & \psi_{2s(3)} & \dots & \psi_{2s(N)} \\ \psi_{3s(1)} & \psi_{3s(2)} & \psi_{3s(3)} & \dots & \psi_{3s(N)} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \psi_{1p(1)} & \psi_{1p(2)} & \psi_{1p(3)} & \dots & \psi_{1p(N)} \\ \psi_{2p(1)} & \psi_{2p(2)} & \psi_{2p(3)} & \dots & \psi_{2p(N)} \\ \psi_{3p(1)} & \psi_{3p(2)} & \psi_{3p(3)} & \dots & \psi_{3p(N)} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \psi_{1d(1)} & \psi_{1d(2)} & \psi_{1d(3)} & \dots & \psi_{1d(N)} \\ \psi_{2d(1)} & \psi_{2d(2)} & \psi_{2d(3)} & \dots & \psi_{2d(N)} \\ \psi_{3d(1)} & \psi_{3d(2)} & \psi_{3d(3)} & \dots & \psi_{3d(N)} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \psi_{1f(1)} & \psi_{1f(2)} & \psi_{1f(3)} & \dots & \psi_{1f(N)} \\ \psi_{2f(1)} & \psi_{2f(2)} & \psi_{2f(3)} & \dots & \psi_{2f(N)} \\ \psi_{3f(1)} & \psi_{3f(2)} & \psi_{3f(3)} & \dots & \psi_{3f(N)} \end{vmatrix}$ Need to check scheme $\psi_{nlm} = \psi_{nlm_1 m_2 m_3}$

(3) work out overall wf using SLATER

(4) calculate $U(r_i) = -e^2/4r_i^2$ from $\psi \rightarrow \psi_{nlm}$ \rightarrow needs to converge

Results of self-consistent field approach → filling order to predict ground state of multi-electron atoms.

Nemonic:



Filling order for ground state.

"Fill from bottom" - Aufbau principle. - understand periodic table - elements with similar outer electron configurations (i.e. $Kr \rightarrow$) - valency have similar chemical properties.

* Coupling schemes - Need to consider spin-orbit interaction as well as electron repulsion for outer electrons.

Hamiltonian has general form $\hat{H} = \hat{H}_0 + \sum_{ij} \frac{e^2}{|r_{ij}|} + \sum_i \Sigma_i(r_i) \hat{L}_i \cdot \hat{S}_i$

\hat{H}_0 - central field, shielded above \hat{H}_1 - residual coupling between unpaired electrons \hat{H}_2 - spin-orbit interaction.

2 scenarios (1) $\hat{H}_1 \gg \hat{H}_2$: light atoms usually; \Rightarrow eigenstates of $\hat{H}_0 + \hat{H}_1$, \hat{H}_2 as perturbation \Rightarrow LS coupling. (2) $\hat{H}_2 \gg \hat{H}_1$: v. heavy atoms / highly ionized light atoms - relativistic effects more important. \Rightarrow jj coupling.

* LS coupling. \hat{H}_1 commutes with $S_z, L_z, S \cdot L$ \Rightarrow energy levels characterised by TOTAL A.M. quantum numbers L, S, J . Order via MUND'S RULES.

(1) combine spin \Rightarrow possible values of S . MAX S lies lowest in energy.

Reason: MAX $S \Rightarrow$ spin w.r.t. symmetric \Rightarrow spatial w.r.t. antisym. \Rightarrow electron repulsion reduced.

(2) combine orbital A.M. pr each S , accounting for exclusion principle \Rightarrow L MAX L lies lowest in energy. Reason: MAX L keeps electron apart $\uparrow \downarrow \uparrow \downarrow$

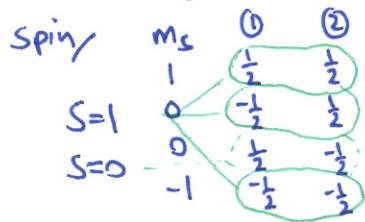
(3) combine $L, S \Rightarrow J$ If shell $\leq \frac{1}{2}$ FULL \Rightarrow MIN J lies lowest $\stackrel{\text{MAX}}{L} \stackrel{\text{MIN}}{J}$
" " $> \frac{1}{2}$ " \Rightarrow MAX J " "

Reason: result of \hat{H}_2 treated as perturbation. "Wigner-Eckart Theorem" used.

* Use Spectroscopists notation to label states $|S, L, J\rangle \equiv {}^{2S+1}L_S$ ($L = S, P, D, F, G$ etc)

Method: Given $|N, L\rangle$ states of no electrons/holes (1) work out

multiplicity of spin states. e.g. pr electron (1) = $3P \equiv |3, 1\rangle$, electron (2) = $4S \equiv |4, 0\rangle$



⇒ 4 states due to spin splitting.

SAME ORBITAL, each S state will have different L values. In other case do L calculation separately for $S=0, S=1$. Note pr $S=0$ w.r.t. antisym so spatial w.r.t. can be the same to satisfy Pauli E.P. \Rightarrow any pair of m_L ok. For $S=1$ only different m_L pairs are allowed.

equivalent \rightarrow L same pr $S=0, S=1$ states.

Simply sum all m_L combinations. e.g. m_L

For each $|S, L\rangle$

State work out S values using $S = L + S - |L - S|$

(1)	(2)	m_L
-1	0	-1 $\Rightarrow L=1$
0	0	0
-1	0	-1

Spectroscopists notation (j_1, j_2, j_3)

→ e.g. States $|S, L\rangle = |0, 1\rangle$ and $|1, 1\rangle$ so far. $|0, 1\rangle : S=1 \quad |1, 1\rangle : S=2, L=0$

→ Hund's rule pr ordering: $\Rightarrow E \uparrow$ Total, Total A.M?

- Hamiltonian commutes with S^2 and with Σ_i^2 since $L_i \cdot S_i = \frac{1}{2}(\hat{L}_i^2 - \hat{L}_i \cdot \hat{S}_i)$ \therefore characteristic eigenstates by $|S, j_1, j_2, j_3, \dots, j_N\rangle$ * compute allowed j pr individual electrons - split according to S-O interaction. * compute allowed $S \rightarrow (j_1, j_2)$ set with split according to S. No MUND rules expect however $j_1 > S$ ($j_1 > E$). i.e. $(\frac{3}{2}, \frac{3}{2}) > E$ then $(\frac{1}{2}, \frac{1}{2})$. Example: $(3P)^2$ - each electron can have $j = l + s$; $|l-s| = 3, 2, 1, 0$. Possibilities: $j_1 = j_2 = \frac{3}{2} \Rightarrow S = j_1 + j_2 = 3, 2, 1, 0$

→ each electron must be distinct (Pauli) so S must be even. (Antisymmetric). $\Rightarrow S=2, 0$ $j_1 = j_2 = \frac{1}{2} \Rightarrow S=1, 0 \Rightarrow S=2 \Rightarrow$ allowed. $j_1 = \frac{3}{2}, j_2 = \frac{1}{2} \Rightarrow S=2, 1$. $j_1 \neq j_2$ so all S allowed.

$$E \uparrow = \begin{matrix} 0^E \\ 1^E \\ 2^E \\ 3^E \\ 4^E \end{matrix} \quad A = (\frac{1}{2}, \frac{1}{2})_0 \quad B = (\frac{3}{2}, \frac{1}{2})_1 \quad C = (\frac{3}{2}, \frac{1}{2})_2 \quad D = (\frac{3}{2}, \frac{3}{2})_2 \quad E = (\frac{3}{2}, \frac{3}{2})_0$$

* Atomic Spectra - transitions between electronic states of atoms via emission/absorption of photons. - rate of spontaneous emission $\propto \omega^3 |\langle \mathbf{k} | \mathbf{d}|\psi\rangle|^2$ in electric dipole approx.

- Selection rules: (Dipole transitions) * Parity change $\Delta S = \pm 1, 0 \rightarrow 0 \times \Delta M_J = \pm 1, 0$

Now for LS coupling $\Delta S=0 \Rightarrow \Delta L = \pm 1, 0 \rightarrow 0 \times \Delta M_L = \pm 1, 0 \times \Delta l_i = \pm 1$ (if only electron involved in transition). (Parity change rule satisfied).

8) Molecular Structure - examples of H_2^+ ion and H_2 molecule to illustrate techniques used.

* H_2^+ ion (Exact solution possible but hard) If protons have coords. R_{a1}, R_{b1} , electron r

→ Full Hamiltonian is: $\hat{H} = -\frac{\hbar^2}{2m_p} (\nabla_{R_a}^2 + \nabla_{R_b}^2) - \frac{\hbar^2}{2me} \nabla_r^2 + \frac{e^2}{4\pi\epsilon_0} \left(\frac{1}{|R_a - R_b|} - \frac{1}{|r - R_{a1}|} - \frac{1}{|r - R_{b1}|} \right)$

- Simplify using "Born-Oppenheimer Approx". $m_p \approx 1835me$ ∴ br gives fast acceleration of electrons \gg protons. ∴ assume electrons respond to changes in proton position INSTANTANEOUSLY.

⇒ Try to solve $\hat{H}\psi = E\psi$ for R_{a1}, R_{b1} fixed ⇒ Energy of system. E ↗ Should we add $|R_{a1} - R_{b1}| = R$ hypercorr.

∴ In B.O. approx S.E. is $\left\{ -\frac{\hbar^2}{2me} \nabla_r^2 + \frac{e^2}{4\pi\epsilon_0} \left\{ \frac{1}{R} - \frac{1}{|r - R_{a1}|} - \frac{1}{|r - R_{b1}|} \right\} \right\} \psi(r, R_{a1}, R_{b1}) = E\psi(r, R_{a1}, R_{b1})$

Note only need to solve $\frac{1}{4\pi\epsilon_0} \left\{ \frac{1}{R} - \frac{1}{|r - R_{a1}|} - \frac{1}{|r - R_{b1}|} \right\} \psi(r, R_{a1}, R_{b1})$ ∵ $\frac{\hbar^2}{2me} \nabla_r^2$ is negligible!!

→ No obviously small term so use variational method. Trial function: - when e is very close to one p-orbit will have negligible effect ⇒ expect w.f. close to p's to be \propto atomic orbitals. Ground state $\psi_{LSLR} = \left(\frac{1}{\pi a_0}\right)^{1/2} e^{-r/a_0}$. Try $\psi = \alpha \psi_a + \beta \psi_b$ where $\psi_{a,b} = \left(\frac{1}{\pi a_0}\right)^{1/2} e^{-|R-R_{a,b}|/a_0}$. Variational $E_0 \leq \frac{\langle \psi | \hat{H} | \psi \rangle}{\langle \psi | \psi \rangle}$ minimize w.r.t R . Now let $S = \langle \psi_a | \psi_b \rangle$ "overlap integral".

and let $H_{ab} = \langle \psi_a | \hat{H} | \psi_b \rangle \Rightarrow E_0 \leq \frac{\alpha^2 H_{aa} + \beta^2 H_{bb} + 2\alpha\beta H_{ab}}{\alpha^2 + \beta^2 + 2\alpha\beta S}$. Now \hat{H} symmetric w.r.t R_{a1}, R_{b1}

⇒ $H_{aa} = H_{bb}$. + Potential symmetric about mid point between two protons. $\alpha = \beta, \alpha = -\beta$. $\Rightarrow E_0 \leq \frac{H_{aa} \pm H_{ab}}{1 \pm S}$ W.F. are normalised $\psi \Rightarrow \psi_g = \alpha(\psi_a + \psi_b) \Rightarrow \alpha_g = \frac{1}{\sqrt{1 \pm S}}$; $\psi_u = \alpha(\psi_a - \psi_b) \Rightarrow \alpha_u = \frac{1}{\sqrt{2(1 \mp S)}}$ Let α ... i.e. $\langle \psi | \psi \rangle = 1$

→ Not such a good fit to real data but could improve trial function with adjustable parameters etc. → Rayleigh Ritz...

* Labelling MO's - g,u notation applies to all homonuclear diatomic molecules (H_2, O_2, \dots) Note diatomic not spherically symmetric → eigenstates of \hat{L}_z^2 not \hat{I}^2 . Label l_{MO} values by Greek letters: $l_{MO} = 0, 1, 2, \dots$ → in case of H_2^+ ion: $\psi_g \leftrightarrow 1S5g, \psi_u \leftrightarrow 1S0_u^*$ \leftrightarrow ANTI-BONDING.

* H_2 Molecule - In B.O. approx: $\hat{H} = -\frac{\hbar^2}{2me} (\nabla_{r_1}^2 + \nabla_{r_2}^2) + \frac{e^2}{4\pi\epsilon_0} \left\{ \frac{1}{r_{12}} + \frac{1}{r_{1a}} + \frac{1}{r_{1b}} + \frac{1}{r_{2a}} + \frac{1}{r_{2b}} + \frac{1}{r_{2c}} \right\}$ where $r_{1a} = |r_1 - R_{a1}|$ etc... ⇒ 2 hfs H_2^+ ion Hamiltonians + SMALL TERM $\frac{e^2}{4\pi\epsilon_0} \left\{ \frac{1}{r_{12}} + \frac{1}{r_{1a}} + \frac{1}{r_{1b}} + \frac{1}{r_{2a}} + \frac{1}{r_{2b}} + \frac{1}{r_{2c}} \right\}$ (ee, pp repulsion). ∴ Product of 2 H_2^+ MO's could be a good start.

- particularly for ground state. i.e. $\psi = \psi_g(r_1)\psi_g(r_2) \rightarrow \langle \psi | \hat{H} | \psi \rangle / \langle \psi | \psi \rangle$ minimise w.r.t $R = R_{ab}$

- NOT THAT GOOD. Why? $\psi_g(r_1)\psi_g(r_2) \sim (\psi_{a1}(r_1) + \psi_{b1}(r_1))(\psi_{a1}(r_2) + \psi_{b1}(r_2)) \sim \{ \psi_{a1}(r_1)\psi_{b1}(r_2) + \psi_{b1}(r_1)\psi_{a1}(r_2) \}$

→ 1st term ψ_{VB} - "Valence Bond" - electrons shared between protons + $\{ \psi_{a1}(r_1)\psi_{a1}(r_2) + \psi_{b1}(r_1)\psi_{b1}(r_2) \}$

→ 2nd term ψ_{IB} - "Ionic Bond" - electrons ^{not} shared but associated with same proton. i.e. $1t - 1t'$

Equal stability/ionization unattainable so try $\psi = \psi_{VB} + \gamma \psi_{IB}$ (γ is small - actually $\approx \frac{1}{6}$).

→ put into $\langle \psi | \hat{H} | \psi \rangle / \langle \psi | \psi \rangle$ and minimise w.r.t. γ and R . Fairly good agreement.

* MO's have their own "spectroscopic notation" $2S+1 \Lambda^u/g$

9) Molecular Spectra - analogous to electronic transitions in atom spectra though $\Delta E \sim \text{eV}$ (optical)

Molecules can rotate and vibrate as well as radiate photons during "orbital transitions".

Rotation $\sim \frac{L^2}{2I} \sim \frac{\hbar^2}{2I} \ll 10^{-4} \text{ eV}$ Vibration $\Delta E \sim 0.1 \text{ eV}$

$2S+1 = \# \text{ spin sublevels } (m_s \text{ values})$

$\Lambda = 1L21 \quad \begin{matrix} 0 & 1 & 2 \\ \Pi & \Delta \end{matrix}$

$1u/g$ Note u/g product $\rightarrow u$

Λ^u/g symmetry. $g_u = uu \rightarrow g$.

i.e. $\psi_{a1}\psi_{b1}(r_1)X_{00} \leftrightarrow \sum_g$

↑ Spin w.f. QM(8)

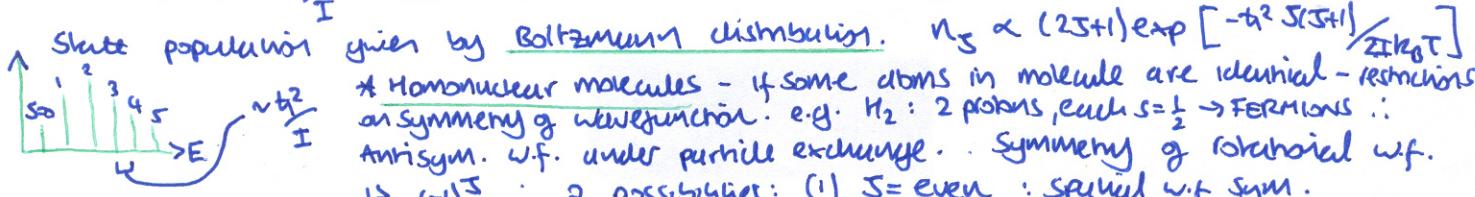
Molecular rotation - Hamiltonian has $\frac{1}{2}I\omega^2 = \frac{1}{2}\frac{L^2}{I}$ ($L=I\omega$). S.E. is $\frac{1}{2I}L^2\ell = E_{\text{rot}}$
- eigenstates of L^2 are spherical harmonics $Y_{\ell m}$, energy $E_{\text{rot}} = \ell(\ell+1)\hbar^2/2I$ (c.f. $L^2\ell = \ell(\ell+1)\hbar^2$)
Now $I = \mu R_0^2$, R_0 eq. separation, μ reduced mass.

→ Recurrent transitions between 1st. states: Dipole transitions PARITY, $\Delta S = \pm 1, 0$

But parity = $(-1)^S \Rightarrow \Delta S = \pm 1$. \therefore Possible energies of emitted/absorbed photons $S+1 \leftrightarrow S$ ($S=0, 1, 2, 3, \dots$) $\Delta E = \pm \frac{h^2}{4} [(S+1)(S+2) - S(S+1)] = \pm \frac{h^2}{4} (S+1)$. Spontaneous emission law

(FGR- energy small) - absorbing better method & Shelly. Absorption spectrum:

lines spaced $\propto \frac{t^2}{T}$. (I sometimes varies $\rightarrow R_0$ at large S - centrifugal potential)



\Rightarrow Spin w.f. antisym. $\Rightarrow S=0$. "PARAHYDROGEN". (a) $S=\text{odd}$: spatial w.f. antisym \Rightarrow spin w.f. sym. $\Rightarrow S=1$ "ORTHOHYDROGEN". Now - Electric dipole selection rules $\Rightarrow \Delta S=0$. $\therefore S=0 \rightarrow S=1$ not possible. Hence H_2 must have $\Delta S=\pm 1$ pure tot. transision. (possible for HCl , HD etc. though). - transitions via collision. relative populations of S,S states $\propto (2S+1)(2S+1) \exp[-S(S+1)h^2/2kT]$ so at high temp ratio ortho/pure $\rightarrow 3:1$. CO_2 : nuclear spin 0^{16} zero. \Rightarrow boson! sym w.f. overall. But spin w.f. sym. $\Rightarrow S$ even. More complicated: $O_2 \rightarrow$ electron w.f. symmetry needed as well. \rightarrow turns out S must be odd for O_2 .
- diatomic atom atom molecule - usually some proton energy change.

* Raman Spectroscopy - Scattering of photon from molecule - usually some photon energy change.
 - arises from second order perturbation theory (time var. dependent). Matrix element in FG.R is
 $\langle \tilde{L}(\tilde{\mu}^1)ij | \sum_k c(k)\tilde{\mu}^1(k) \times \epsilon(k) \tilde{\mu}^1(k) \rangle$ i.e. transition $j \rightarrow k, k \rightarrow i$. (ϵ over k pr. rule). If $\tilde{\mu}^1 =$ Electric dipole operator
 $E_i - E_j$ $\therefore +1, 0, -1$ unit ΔE . $|k\rangle, |i\rangle$ same. $\Rightarrow |i\rangle, |j\rangle : \Delta J = \pm 2, 0$

- Selection rules: $|E_j - E_k|$ differ by $\pm 1, \Phi_R$ parity units. $(\Delta j = \pm 1) \Rightarrow$ (Parity $\Rightarrow \pm 1$).
 Also parity change in both $j \rightarrow k$ and $n \rightarrow i \Rightarrow$ RAMAN SELECTION RULES $\Delta S = \pm 2, 0$, NO PARITY CHANGE

$$\Delta E_{\text{photon}} = 0 \text{ if } \Delta S=0; J \rightarrow J+2 \Rightarrow \Delta E = \frac{\hbar^2}{2I} (S(J+1) - S(J+2)(J+3)) = \frac{-\hbar^2}{2I} (4S+6) \quad S=0,1,2,3, \dots$$

$\Rightarrow \Delta E = \frac{\hbar^2}{2I} [4S-2], S=2,3,4, \dots \leftarrow \text{ANTI STOKES LINES.}$

$J \rightarrow J-2$ - good way to observe rot. transitions (Any photon can be used - pick convenient wavelength) $E_g \uparrow$

* Vibrational transitions - diatomic molecules - expand $E_0(R)$ about R_0
- approximate as harmonic oscillator potential $\frac{1}{2}M\omega^2(R-R_0)^2$. $\omega^2 = \frac{1}{M} \frac{\partial^2 E_0}{\partial R^2}|_{R_0}$
Energies are $(n+\frac{1}{2})\hbar\omega + E_0(R_0)$. Typical energy $\sim \frac{1}{50}$ eV/atom. Selection rules: Dipole $\rightarrow \Delta n \neq 1 \Rightarrow \Delta E = \hbar\omega$
(pure v.v.). Note pot. not merely harmonic $\sim (R-R_0)^3$ etc... $\Delta n = \pm 1$ not smoothly true. Polyatomic
molecules - more modes ($3N-6$) pr N atoms. \rightarrow Normal mode analysis.

* Vib. rotation of diatomic molecule - Vib. transitions often accompanied by rot. Because interatomic potential is not exactly harmonic $\Rightarrow R_0 \sim$ different in two different vib. states. $\Rightarrow I$ larger in state of higher n . Consider absorption: $n \rightarrow n+1$ $\Delta E = \hbar \omega$. Define $B = \frac{\hbar^2}{2I}$, $B' = \frac{\hbar^2}{2I'}$, $B' = B + \delta B$ ($\delta B < 0$)
+ for. change $I = I' \Rightarrow \delta B = 0$

$$(2) \Delta S = -1, \Delta E = \hbar\omega - 2BS + \delta B(S-1)S \quad S=1,2,3\dots \quad (3) \Delta S = 0, \Delta E = \hbar\omega + \delta B(S+1)S \quad S=1,2,3\dots$$

Now $|S| \ll B \rightarrow$ observe: ~equally spaced lines above & below "R Branch" $\Delta S = +1$
 ~" " " " below trw "P Branch" $\Delta S = -1$; "Q Branch" $\Delta S = 0$
 \rightarrow In P-Branch spacing increases \sim with S , decreased with S for R.
 (S.B.) NOT SEEN (P trw)
-No parity change.

* Electronic Transitions. - if electrons move transitions in a molecule \rightarrow include vib and rot.

- Franck -康登 principle: Many nuclei - Electronic transiting leaves nuclei unaffected.

E

→ initial - final shift: not vib level: $\Delta E \approx \frac{1}{2} M \omega^2 (R_0^+ - R_0^-)^2 \cdot \Delta J$ sizeable. - FB not effect small. (Analysis above still valid though). ΔE in UV spectrum. - absent in emission

- Note Q-Branch may be seen if electronic transition satisfies parity selection rule.
- even/odd R branch may increase @ small $J +$ decrease at large $J \Rightarrow$ "BAND HEAD".

$\Delta E_1 S \Rightarrow$ "FORTRAN DIAGRAM" $S \uparrow$

* E.M. Fields in QM * Hamiltonian operator in QM based on Hamiltonian formulation of classical mechanics. For a system of N particles (coordinates x_i ($i=1 \dots 3N$), p_i = momentum)

$H = T + V$. $\frac{dx_i}{dt} = \frac{\partial H}{\partial p_i}$, $\frac{dp_i}{dt} = -\frac{\partial H}{\partial x_i}$ if replace p_i by $p_i - gA_i$ (vector potential A)
 \Rightarrow generate Lorentz form Hamilton's equations - $\dot{m}\dot{x}_i = qE_i + qv_i \times B$. So in general
 e.g. use $p = -i\hbar\nabla - qA$ in S.E. if $v = q\phi \Rightarrow$ S.E. $\frac{1}{2m}(-\hbar^2 \nabla^2 + i\hbar q(2A \cdot \nabla + q\phi) + q^2 A^2) + q\phi\hat{t} = E\hat{t}$ (or $i\hbar \frac{\partial}{\partial t}$)
 - can neglect A_z^2 term in weak fields.

* Uniform magnetic field - choose potentials $A = -\frac{1}{2}r \times B$, $\phi = 0$. (Note $B = \nabla \times A$ ✓). Note "Gauge Gauge" $\nabla \cdot A = 0$ if $\nabla \cdot r = 0$ or pr. ad).

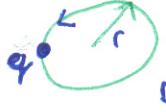
Take $B = (0, 0, B_z) \Rightarrow A = \frac{B_z}{2}(-y, x, 0)$. Note "Gauge Gauge" $\nabla \cdot A = 0$ if $\nabla \cdot r = 0$ or pr. ad).

$$\Rightarrow \hat{H} = -\frac{\hbar^2}{2m} \nabla^2 - \frac{q}{2m} B_z \hat{l}_z + \frac{q^2 B_z^2}{8m} (x^2 + y^2) \text{ noting } \hat{l}_z = i\hbar \left(x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x} \right) = x \hat{p}_y - y \hat{p}_x$$

$$\Rightarrow 3D \text{ generalise to } \hat{H} = -\frac{\hbar^2}{2m} \nabla^2 - \frac{q}{2m} B_z \hat{l}_z + \frac{q^2}{8m} [B^2 r^2 - (B_z r)^2]$$

Now B^2 term negligible for $B \lesssim 10^4 T$

\Rightarrow important term is $-\frac{q}{2m} B_z \hat{l}_z$. "magnetic moment associated with e (stationary) sum interacting with B " i.e. $-\mu \cdot B$ with $\hat{\mu} = \frac{q}{2m} \hat{l}_z$. Denote μ classically considering purhill in circular motion \Leftrightarrow current loop.

q  $L = mr^2\omega$. current = $I = qr\omega/2\pi \Rightarrow \mu = I\pi r^2 = \pi r^2 \frac{q^2}{mr^2/2\pi} = \frac{q^2 L}{2m}$.

Define BORN MAGNETON $\mu_B = \frac{e\hbar}{2me} \therefore \hat{\mu}_{\text{magnetic}} = -\frac{q}{2m} B_z \hat{l}_z = \mu_B B_z \hat{l}_z$

\therefore If $\hat{\mu}_{\text{magnetic}} = \gamma \hat{q}$, $\gamma = \mu_B B_z m_e \quad \{ m_l = -l, -l+1, \dots, +l \}$.

* Larmor precession and random levels - magnetic moment associated with \hat{l}
 \rightarrow classically expect PRECESSION about B (Euler equation etc). Also occurs in QM.

If $B = B_z \Rightarrow \frac{d}{dt} \langle \hat{l}_x \rangle = \frac{i}{\hbar} \langle [\hat{H}, \hat{l}_x] \rangle = \frac{i}{\hbar} \frac{eB_z}{2m} \langle [\hat{l}_y, \hat{l}_x] \rangle = -\frac{eB_z}{2m} \langle \hat{l}_y \rangle = -\omega_L \langle \hat{l}_y \rangle$
 $\omega_L = \frac{eB_z}{2m} = \text{'larmor frequency'}$. Also $\frac{d}{dt} \langle \hat{l}_y \rangle = \omega_L \langle \hat{l}_x \rangle \Rightarrow \langle \hat{l}_y \rangle = A \cos(\omega_L t + \alpha)$ i.e. precesses

Now consider electron in eigenstate of \hat{l}_z in uniform $B = B_z \hat{z}$.

NOT SMALL. $\hat{H} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial z^2} + \mu_B B_z \hat{l}_z + \left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + \frac{e^2 B_z^2}{8m} x^2 \right) + \left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial y^2} + \frac{e^2 B_z^2}{8m} y^2 \right)$

Harmonic oscillator parts have main in z direction \hat{l}_z, B interaction

frequency $\frac{1}{2}m\omega^2 = \frac{e^2 B_z^2}{8m} \Rightarrow \omega = \frac{eB_z}{2m} = \omega_L$. Energies $(n_x + \frac{1}{2})\hbar\omega_L, (n_y + \frac{1}{2})\hbar\omega_L$.

$\therefore E_{\text{tot}} = \frac{p_z^2}{2m} + (m_l + n_x + n_y + \frac{1}{2})\hbar\omega_L$ "LANDAU LEVELS". Now by purity of \hat{l}_z ,

Harmonic oscillator w.f. $\Rightarrow m_l + n_x + n_y = \text{even} = 2n$. Also $E > 0$ ($\hat{H} = \frac{(p - qA)^2}{2m}$ is a square)
 $\Rightarrow n \geq 0$. So $E = \frac{p_z^2}{2m} + (2n + 1)\hbar\omega_L, n \geq 0$.

* Spin magnetic moment - Dirac equation predicts POINTLIKE spin $\frac{1}{2}$ purhill has magnetic moment $\mu = g \cdot \frac{e}{2m} \vec{S}$ with $g = 2$. (QED $\Rightarrow g = 2.002$).

\therefore for electron $\mu_z = \pm \frac{e}{m_e} \frac{\hbar}{2} = \pm \mu_B$. \therefore modify (H) from SE to $\hat{H} = -\frac{\hbar^2}{2m} \nabla^2 - \mu \cdot B + \frac{q^2}{8m} (B^2 r^2 - B_z^2)$ where $\mu = \frac{g}{2m} (\hat{l}_z + 2\hat{S}_z)$. Note μ is not \parallel to $\vec{S} = \hat{l}_z + \hat{S}_z$ in general. Note proton has $g = 5.58$ i.e. NOT POINT LIKE.

* ZEEMAN EFFECT - when atoms are placed in a magnetic field spectral lines are observed to split. Including SPIN ORBIT INTERACTION (not inc. in (H)) \rightarrow atom in B_z field has Hamiltonian: $\hat{H} = \hat{H}_0 + \xi \hat{l}_z \cdot \hat{S}_z + \frac{e}{2m} B_z (\hat{l}_z + 2\hat{S}_z)$. \hat{H}_0 commutes with $\hat{l}_z^2, \hat{l}_z \cdot \hat{S}_z^2, \hat{S}_x^2, \hat{S}_y^2$ but $\hat{l}_z \cdot \hat{S}_z$ commutes with \hat{l}_z^2 or \hat{S}_z^2 in general. \Rightarrow TWO REGIMES:

(1) STRONG FIELD $\mu_B B \gg \xi \hat{l}_z \cdot \hat{S}_z$

B large or $\hat{l}_z \gg \hat{S}_z$ (paramagnetic limit).

(2) WEAK FIELD $\mu_B B \ll \xi \hat{l}_z \cdot \hat{S}_z$

Find eigenstates of $\hat{H}_0 + \xi \hat{l}_z \cdot \hat{S}_z$

'ANOMALOUS ZEEMAN EFFECT'

then weak $\mu_B \hat{l}_z \cdot \hat{S}_z$ as perturbation

QM(1)

* STRONG FIELD EXAMPLE $2p \rightarrow 1s$ transition in H.

$1s$	$m_s = +\frac{1}{2}$	$E - E_0 = +\mu_B B_z$	$2p$	$m_l = +1$	$m_s = \frac{1}{2}$	$E - E_0 = 2\mu_B B_z$
	$-\frac{1}{2}$	$-\mu_B B_z$		$+1$	$-\frac{1}{2}$	0
		↑ can write it all down!		0	$\frac{1}{2}$	$\mu_B B_z$
				-1	$+\frac{1}{2}$	0
				-1	$+\frac{1}{2}$	$-2\mu_B B_z$

- Selection rules: $\Delta m_s = 0$, $\Delta m_l = \pm 1, 0$

\Rightarrow allowed transitions are $\Delta E = |E_0 - E|$, $|E_0 - E| \pm \mu_B B_z$

i.e. spectral line split into 3 components.

$| \rightarrow \xrightarrow{\mu_B B_z} | E_0 - E | \rightarrow E$

* WEAK FIELD ANALYSIS + EXAMPLE

Firstly consider $\Sigma \leftarrow \Sigma$ term. Write $\Sigma \leftarrow \Sigma = \frac{1}{2}(\Sigma^2 - \Sigma_L \cdot \Sigma)$ since $\Sigma^2 = \Sigma^2 + \Sigma^2 + 2\Sigma \cdot \Sigma$

\Rightarrow eigenstates of Σ & $\Sigma \leftarrow \Sigma$ are eigenstates of $\Sigma^2, \Sigma^2, \Sigma^2, \Sigma^2$ BUT NOT Σ^2, Σ^2 . \Rightarrow state $|jLS\rangle$.

$\therefore \Delta E$ due to $\Sigma \leftarrow \Sigma = \langle jLS | \Sigma \leftarrow \Sigma | jLS \rangle = \frac{\epsilon_{jL}^2}{2} (j(j+1) - S(S+1) - L(L+1))$

Now using 1st order perturbation theory ΔE due to $\frac{e}{2m} B_z (\hat{L}_z + 2\hat{S}_z)$ is $\langle jLS | \frac{e}{2m} B_z (\hat{L}_z + 2\hat{S}_z) | jLS \rangle$ needed.

Evaluate this using 'vector model'. - represent state $|LSjmj\rangle$ by:

- lengths $g \perp \Sigma, \Sigma$ fixed by (e.g.) $|\Sigma| = \sqrt{j(j+1)}$
- angles between Σ and \hat{z} fixed and between any pair
- $g \Sigma \perp L, S$ e.g., $\cos\alpha = \frac{\Sigma_z}{|\Sigma|} = \frac{m_j}{\sqrt{j(j+1)}} \quad - M = \frac{e}{2m} (\hat{L} + 2\hat{S}) \perp \Sigma$
-
- $\therefore g = \frac{3j(j+1) - L(L+1) - S(S+1)}{2j(j+1)} \quad \text{so } M \text{ not fixed. BUT angle between } M, \Sigma (\beta) \text{ is fixed so } M \text{ moves around with } \Sigma$
- $\text{Prob.: } |\mu||J| \cos\beta = M \cdot \Sigma \propto (\hat{L} + 2\hat{S}) \cdot (\hat{L} + \hat{S}) = L^2 + 2S^2 + 3L \cdot S = \frac{1}{2}(3S^2 - L^2 + S^2) \text{ Fixed.}$

- Can i. calculate M_2 as $M_2 = |\mu| \cos\beta \cdot \cos\alpha = M \cdot \Sigma \frac{S_z}{|\Sigma|^2} \Rightarrow \langle M_2 B_z \rangle |LSmj\rangle = g M_B M_j B_z$

where $g = \frac{3j(j+1) - L(L+1) - S(S+1)}{2j(j+1)}$ LANDÉ g-FACTOR. (using $M \cdot \Sigma = \frac{1}{2}(3S^2 - L^2 + S^2) \cdot \frac{e}{2m}$)

so (e.g) for S state of H: ($\Rightarrow, S=\frac{1}{2} \Rightarrow j=\frac{1}{2} \Rightarrow g=2$) so NS: $E = E_0 + \Delta E$

$\Delta E = 2 \times g M_B \cdot M_{jmax} \cdot B_z = 4 M_B \frac{1}{2} B_z = 2 M_B B_z$ {Note Selection rules $\Delta m_s = \pm 1, 0$ for transitions} $\Delta E = \frac{\epsilon_{jL}^2}{2} \left(\frac{1}{2}(\frac{1}{2}+1) - \frac{1}{2}(\frac{1}{2}+1) \right) = 0$

* Polarisation in Zeeman effect

- Selection rules $\Delta m_s = 0$ for z-dipoles i.e. $\Sigma \parallel \hat{z}$ $\Delta m_J = \pm 1$ for x,y dipoles.
- photon travelling $\parallel z$ can only have x,y components $\& \Sigma \therefore$ if view spectrum $\parallel B$ only see $\Delta m_J = \pm 1$ transitions. (4, 2 lines)
- photon travelling $\parallel x$ can have y,z components $\& \Sigma \Rightarrow$ all Δm_J possible. $\frac{N_B}{R_{22}} =$
- view $\parallel B$, $\Delta m_J = \pm 1$ transitions will have OPPOSITE CIRCULAR POLARISATIONS.
- $\Delta m_J = +1 \Rightarrow \alpha + iy, \Delta m_J = -1 \Rightarrow \alpha - iy$.
- view $\perp B$ i.e. using x: $\Delta m_J = \Rightarrow$ plane polarised $\parallel z, \parallel B$. \Rightarrow MORE ABOUT THIS IN ATOMIC PHYSICS!

* Magnetic susceptibility

- magnetic moments of atoms have discrete values $g \mu_B B_m$, typical energy for $B \approx 1T \approx 10^{-23} J \ll k_B T$ @ room temp. ($\sim k_B T$ @ 1K). \therefore Normally m_s values populated ~ equally. - probability of being in state $m_s = \exp(-g \mu_B B_m / k_B T) / \sum_i \exp(-g \mu_B B_m / k_B T)$
- \Rightarrow average $M \parallel B$ is $\langle M_z \rangle = -g \mu_B \sum_i m_s e^{um_s} / \sum_i e^{-um_s} \propto \frac{g \mu_B B}{k_B T} = \chi$. m_s if we can get Curie's law for χ_c /atom.
- $\chi_c = \frac{M_0 \langle M_z \rangle}{B} = \frac{\mu_0 g^2 \mu_B^2}{B} \frac{1}{j(j+1)}$. Note have neglected B^2 term in $H \Rightarrow$ DIAMAGNETISM with $\Delta E \propto B^2$. Unbalance $3k_B T$ even if $(=S=1)$. Energy $\langle e^2 B^2 (x^2 + y^2) \rangle = \frac{e^2 B^2}{8m} \langle r^2 \rangle$ so if $\langle x^2 \rangle = \langle y^2 \rangle = \langle z^2 \rangle = \frac{1}{3} \langle r^2 \rangle \Rightarrow$ induced dipole energy $= -\frac{1}{2} \mu B$ with $\mu = \frac{e^2 B}{6m} \langle r^2 \rangle$
- \therefore DIAMAGNETIC SUSCEPTIBILITY / ATOM $\chi_0 = \frac{\mu_0 e^2}{6m} \langle r^2 \rangle$ DM(11) all electrons in atom.

Extra // * More correct treatment of EM waves in QM.

Consider travelling wave polarized $\parallel z$ $\epsilon_z = \epsilon_0 e^{i(hz-wt)}$ (travels in α direction)
vector potential is $\underline{A} = (0, 0, A_z)$ $A_z = A_0 e^{i(hz-wt)}$; $\phi = 0$
 $\underline{\xi} = -\frac{\partial \underline{A}}{\partial t} \Rightarrow \epsilon_0 = i\omega A_0 \Rightarrow \underline{B} = (0, B_y, 0); B_y = i\omega A_0 e^{i(hz-wt)} = \frac{k}{\omega} \epsilon_z = \frac{1}{c}$

Now, neglecting A_z^2 term in \hat{H} , $\hat{H} = \hat{H}_0 + \hat{H}'$ for Hydrogen like atom
with $\hat{H}' = i\frac{qz}{m} \cdot \underline{A} \cdot \nabla = i\frac{qz}{m} A_0 \frac{\partial}{\partial z}$. (Note $\nabla \cdot \underline{A} = 0$; electric dipole approx
 $\Rightarrow \text{no } \propto 0$; also omit e^{iwt} for clarity).

- use TRICK $[\hat{H}_0, z] = -\frac{\hbar^2}{m} \frac{\partial^2}{\partial z^2}$

$$\Rightarrow \langle k | \hat{H}' | i \rangle = -i\frac{qz}{m} \cdot \frac{mA_0}{\hbar^2} \langle k | \hat{H}_0 z - z \hat{H}_0 | i \rangle$$

$$= -i\frac{qz}{\hbar} A_0 (E_k - E_i) \langle k | z | i \rangle$$

$$= -i\omega q A_0 \langle k | z | i \rangle \quad \text{with } \hbar\omega = E_k - E_i$$

$$= -q \epsilon_0 \langle k | z | i \rangle$$

As postulated previously (QM ④ - ⑤).