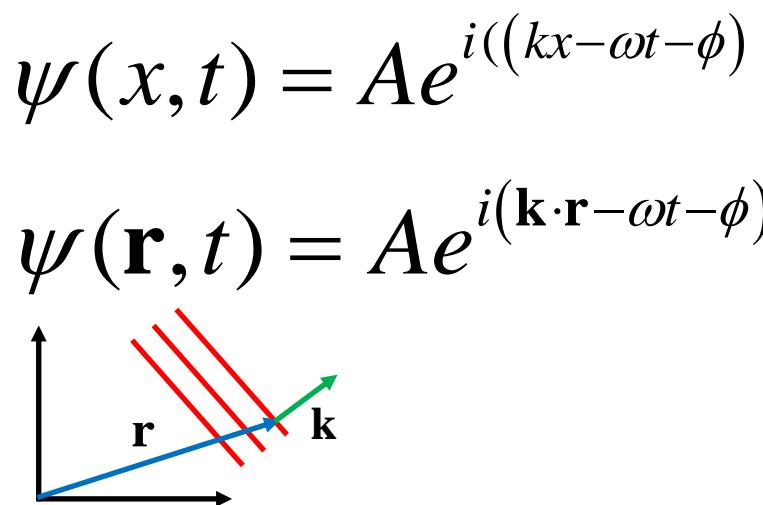
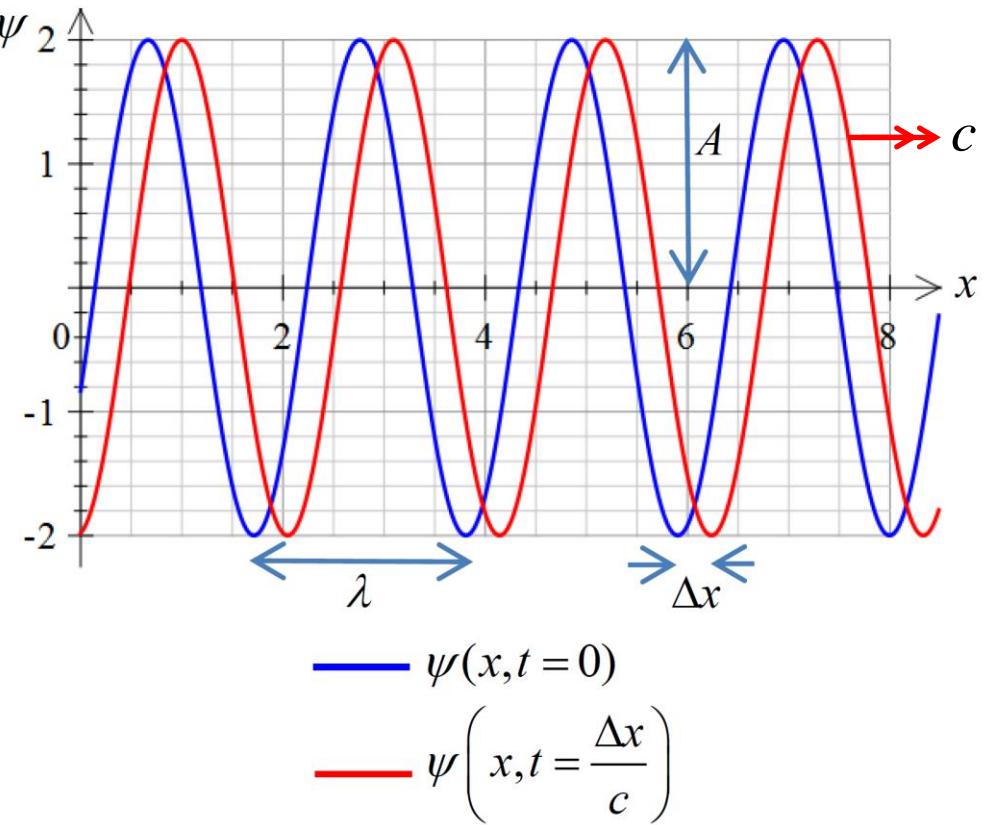


# BPhO

## Computational Challenge

# Waves and optics

Dr Andrew French.  
December 2023.



Wave speed  $c = f\lambda$

Frequency = 1/period  $f = \frac{1}{T}$

Wavenumber  $k = \frac{2\pi}{\lambda}$

$$\psi(x, t) = A \cos\left(2\pi \frac{x - ct}{\lambda} - \phi\right)$$

$$\psi(x, t) = A \cos(kx - \omega t - \phi)$$

Wave Fourier components are **translated sinusoids**

$\omega = ck$

```

A = 1; %Amplitude
lambda = 0.5; %Wavelength /m
c = 340; %Wavespeed /ms^-1
f_unit = 'kHz'; %Frequency unit
t_unit = 'ms'; %Time unit

%Spatial sinusoid separation (fraction of wavelength)
dx = lambda/8;

fsize = 18; %Fontsize
N = 1000; %Number of data points
Nwaves = 3; %Number of waves to plot
linewidth = 2; %Line width

%
%Determine period /s and frequency /Hz
f = c/lambda;
T = 1/f;

%Determine time delay associated with dx
dt = dx/c;

%Generate x,y and t,y data vectors
x = linspace( 0, Nwaves*lambda, N );
t = linspace( 0, Nwaves*T, N );

yx1 = A*sin( 2*pi*x/lambda );
yx2 = A*sin( 2*pi*( x- dx)/lambda );
yt = A*sin( 2*pi*t/T );

```

```

%Plot y vs x graph
figure('color',[1 1 1],'name','waves anatomy')
plot(x,yx1,'b-',x,yx2,'r-','linewidth',linewidth);
legend( {'t = 0','t = \Delta t'},'fontsize',fsize )
xlabel( 'x /m','fontsize',fsize )
title([' A = ',num2str(A),' , \lambda = ',num2str(lambda),...
',m, c = ',num2str(c),'m/s'], 'fontsize',fsize )
grid on;
set( gca,'fontsize',fsize )
axis tight
print( gcf,'wave y vs x.png','-dpng', '-r300' )
clf;

%Plot y vs t graph
plot(t/tuf( t_unit ),yt,'linewidth',linewidth);
xlabel( ['t /',t_unit], 'fontsize',fsize )
title([' A = ',num2str(A),' , \lambda = ',num2str(lambda),...
',m, c = ',num2str(c),'m/s, f = ',...
num2str( f/fuf( f_unit ),5),f_unit,', T = ',...
num2str(T/tuf( t_unit ),5),t_unit], 'fontsize',fsize )
grid on;
set( gca,'fontsize',fsize )
axis tight
print( gcf,'wave y vs t.png','-dpng', '-r300' )
close(gcf);

%%

```

```

%Frequency unit factor
function m = fuf( f_unit )
if strcmp( f_unit, 'THz' )==1
    m = 1e12;
elseif strcmp( f_unit, 'GHz' )==1
    m = 1e9;
elseif strcmp( f_unit, 'MHz' )==1
    m = 1e6;
elseif strcmp( f_unit, 'kHz' )==1
    m = 1e3;
else; m = 1; end

```

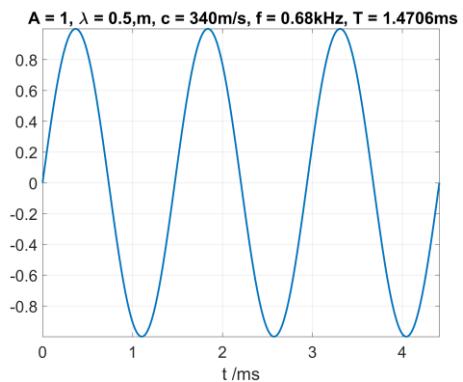
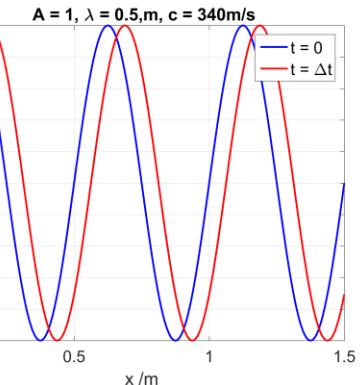
```

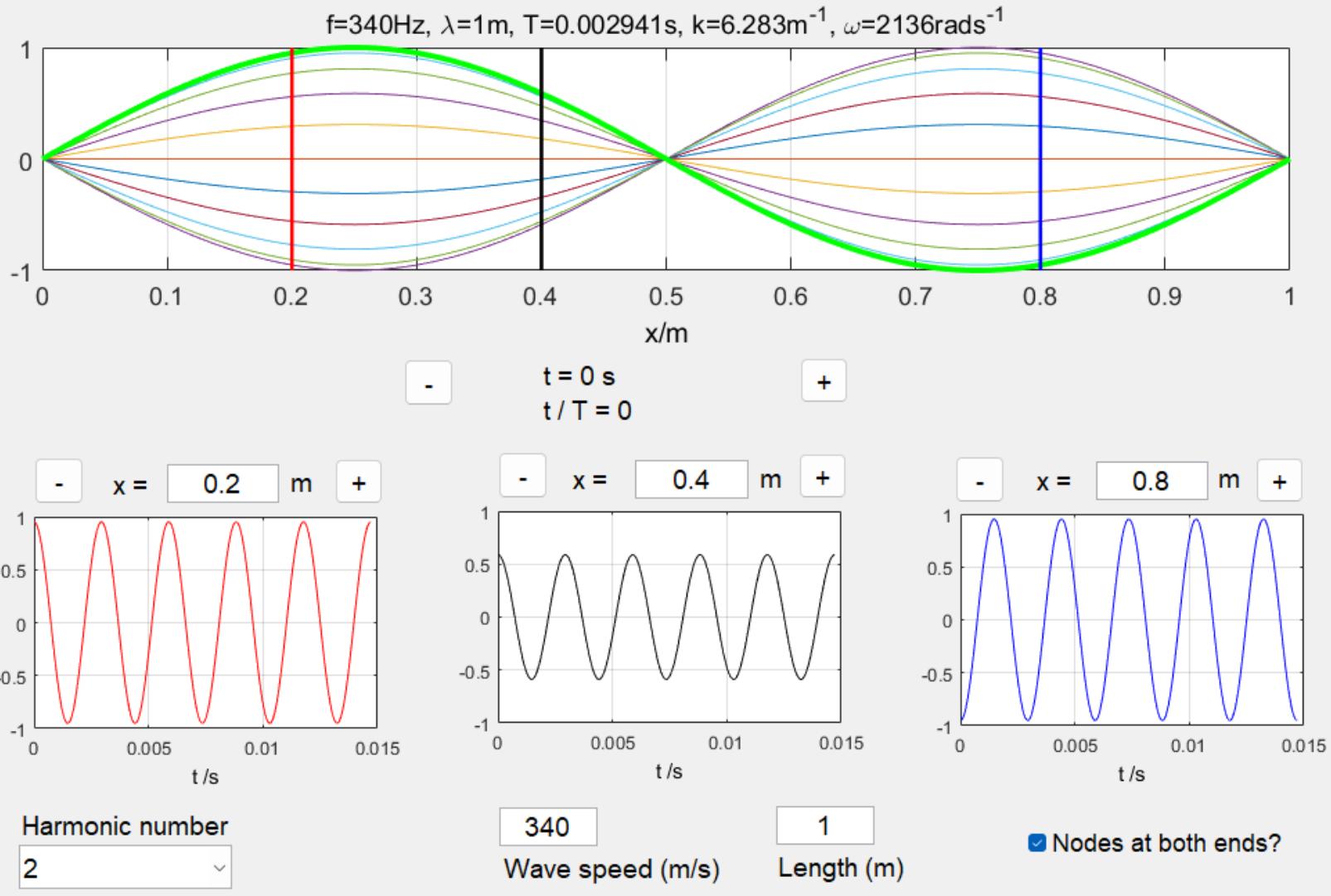
%
```

```

%Time unit factor
function m = tuf( t_unit )
if strcmp( t_unit, 'ps' )==1
    m = 1e-12;
elseif strcmp( t_unit, 'ns' )==1
    m = 1e-9;
elseif strcmp( t_unit, '\mu s' )==1
    m = 1e-6;
elseif strcmp( t_unit, 'ms' )==1
    m = 1e-3;
else; m = 1; end

```

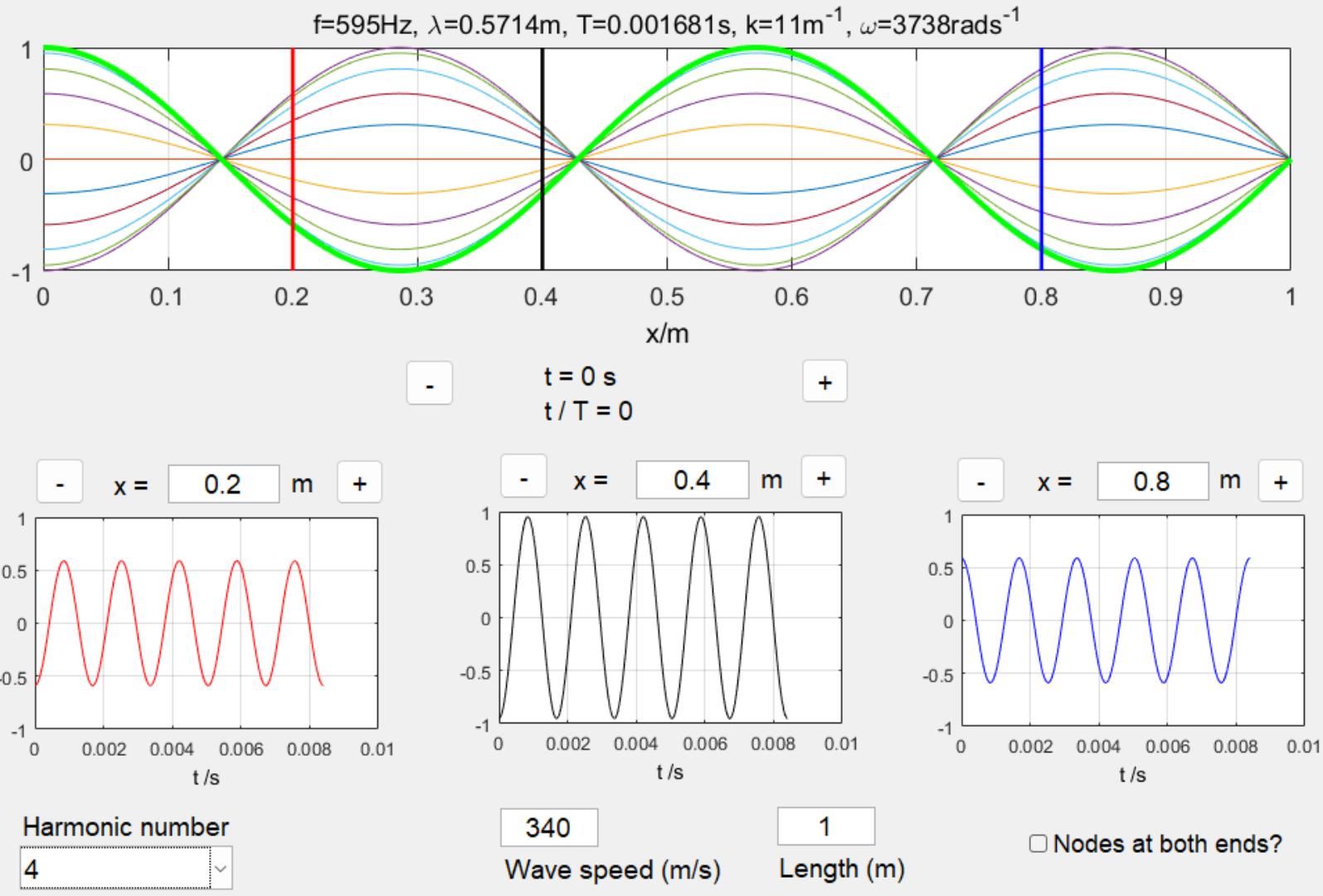




Length is a whole number of half wavelengths

$$L = n \frac{1}{2} \lambda$$

$$n = 1, 2, 3, 4, \dots$$

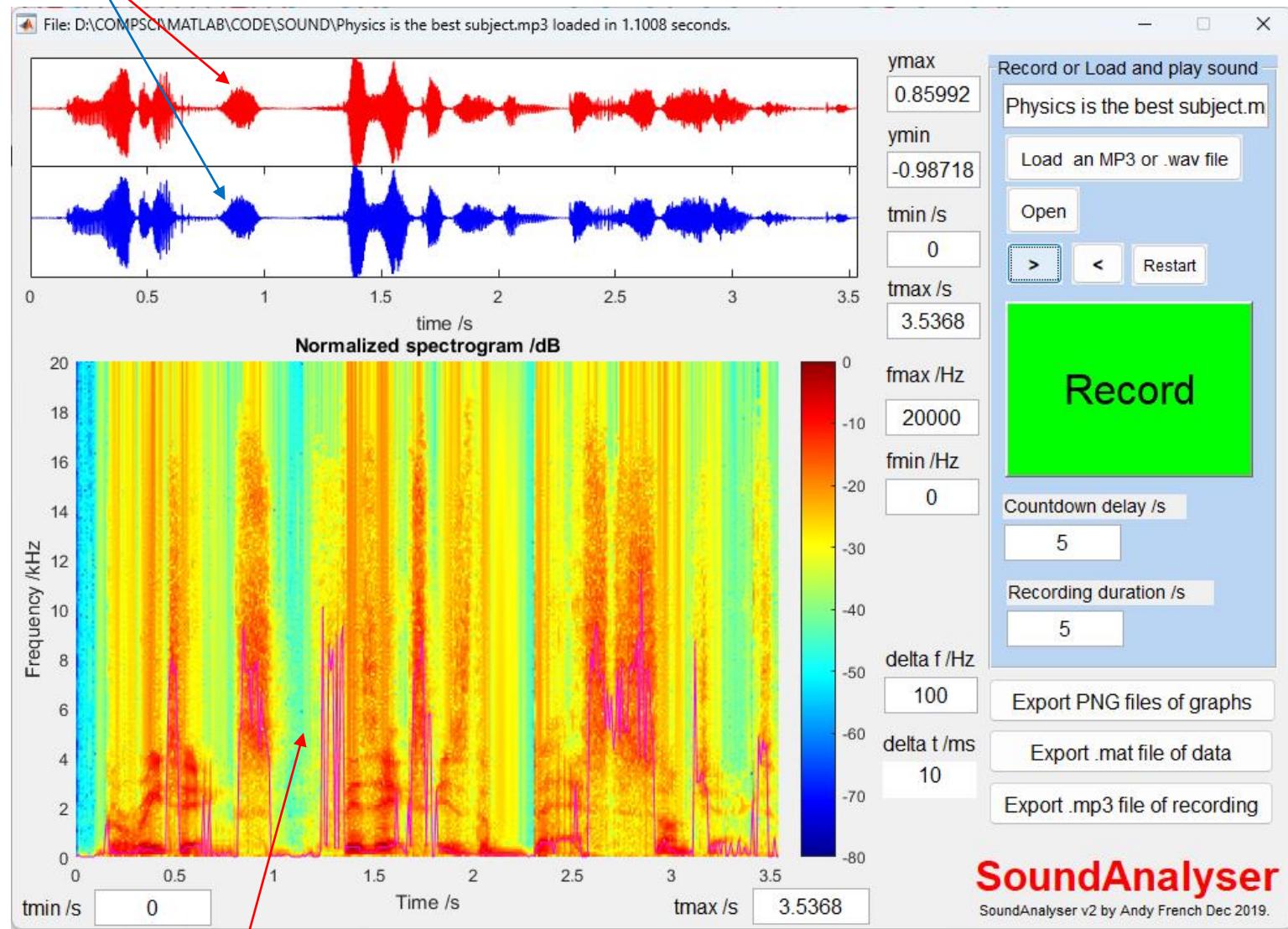


Length (plus end correction) is an odd number of quarter wavelengths

$$L + \frac{2}{3} r = (2n-1) \frac{1}{4} \lambda$$

$$n = 1, 2, 3, 4\dots$$

# Left and right stereo channels of a sound wave recording



Plot of **frequency content** of waves vs time.  
Colour scale is **decibels** (dB) with max power set at 0dB.

$$dB = 10 \log_{10} (\text{signal power})$$

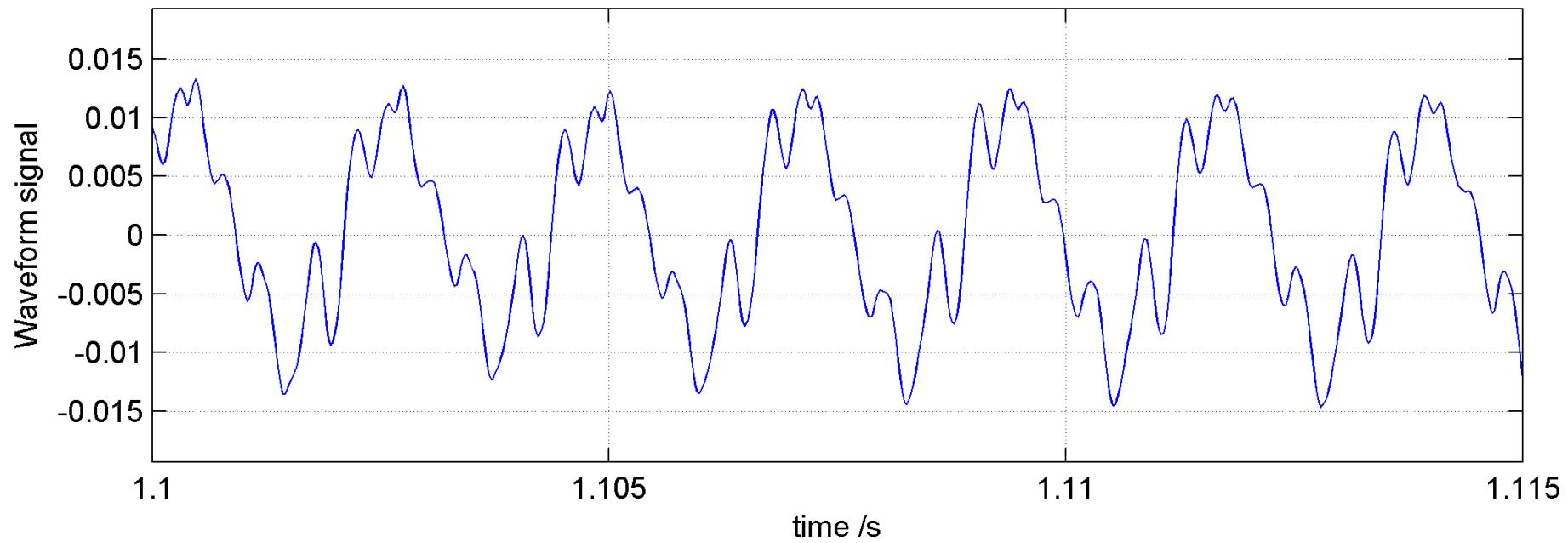
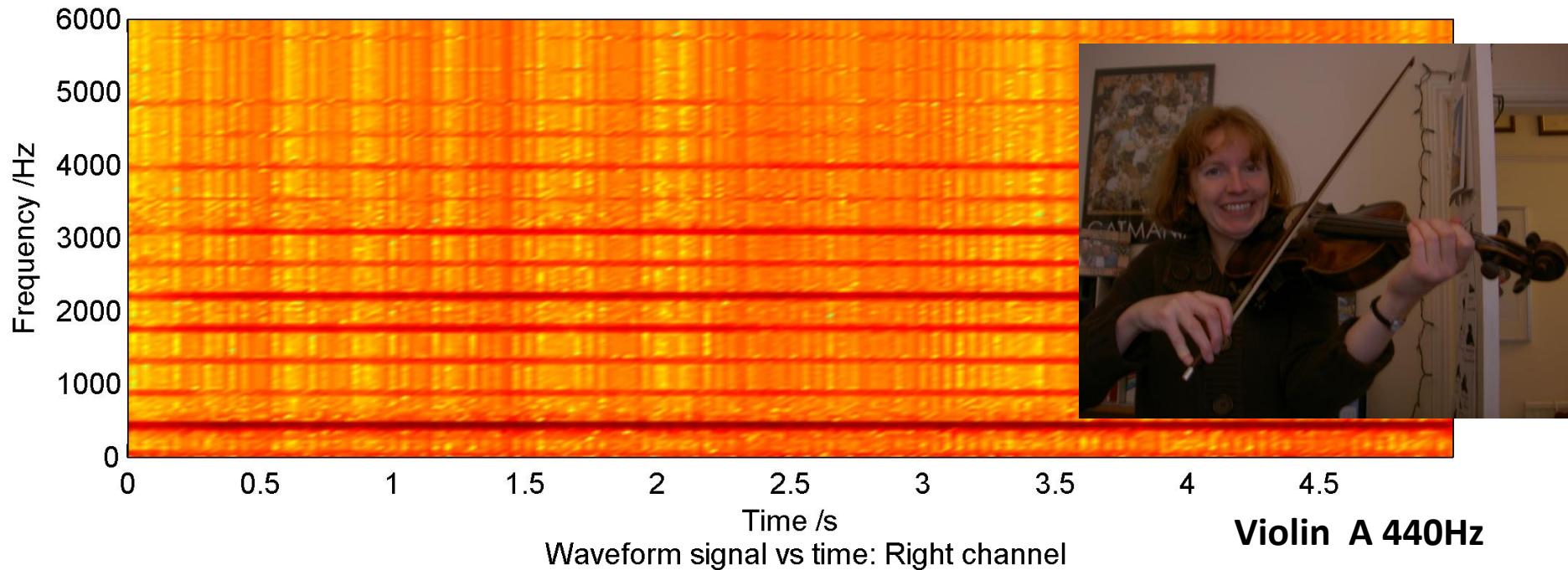
# Musical harmony



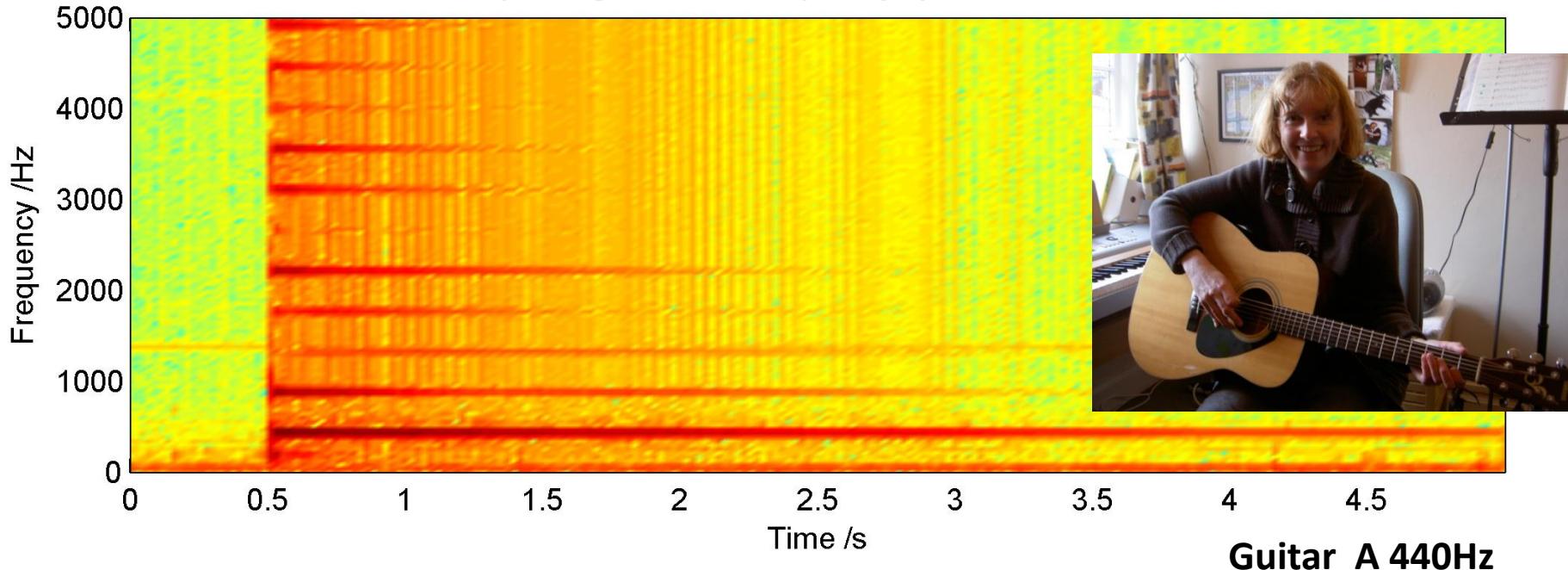
- The mathematics of music has been known since the time of Pythagoras, 2500 years ago
- Frequency intervals of simple fractions e.g. 3:2 (a fifth) yield ‘harmonious’ music
- An **octave** means a **frequency ratio of 2**. An octave above concert A (440Hz) is therefore 880Hz. An octave below is 220Hz.
- The modern ‘equal-tempered scale’ divides an octave (the frequency ratio 2) into twelve parts such that

$$f_n = 2^{n/12} = \sqrt[12]{2}$$

Normalized spectrogram /dB: Frequency spectrum variation with time

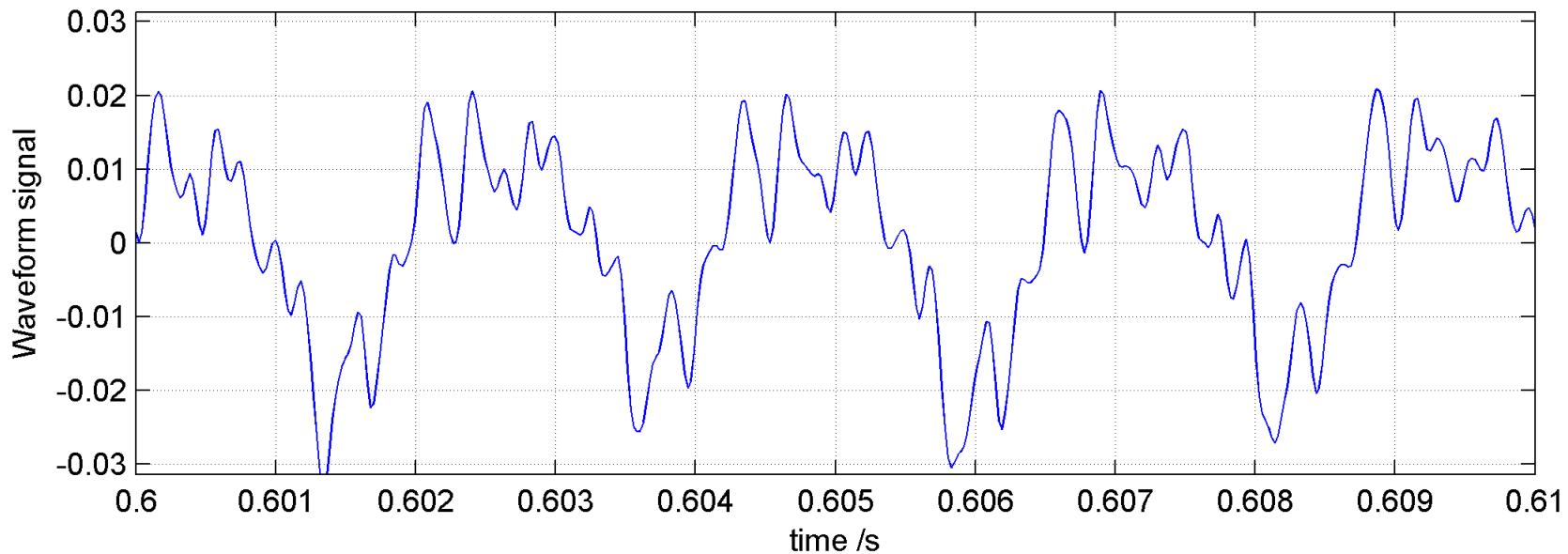


Normalized spectrogram /dB: Frequency spectrum variation with time

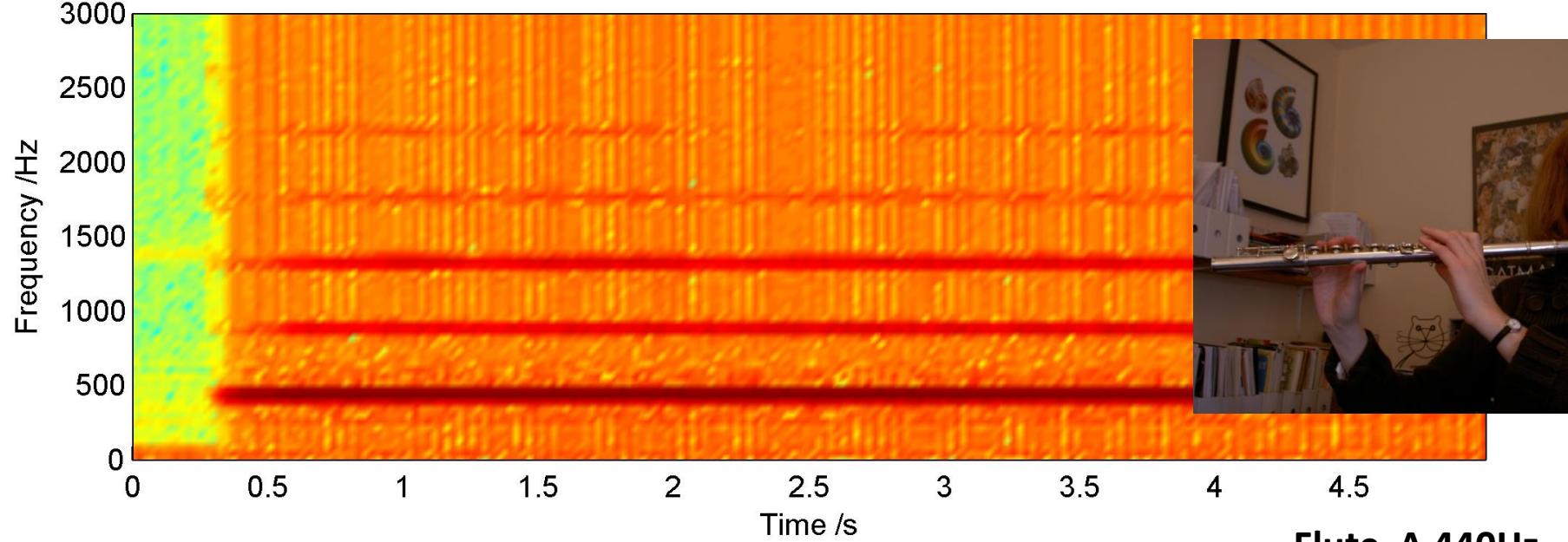


**Guitar A 440Hz**

Waveform signal vs time: Right channel

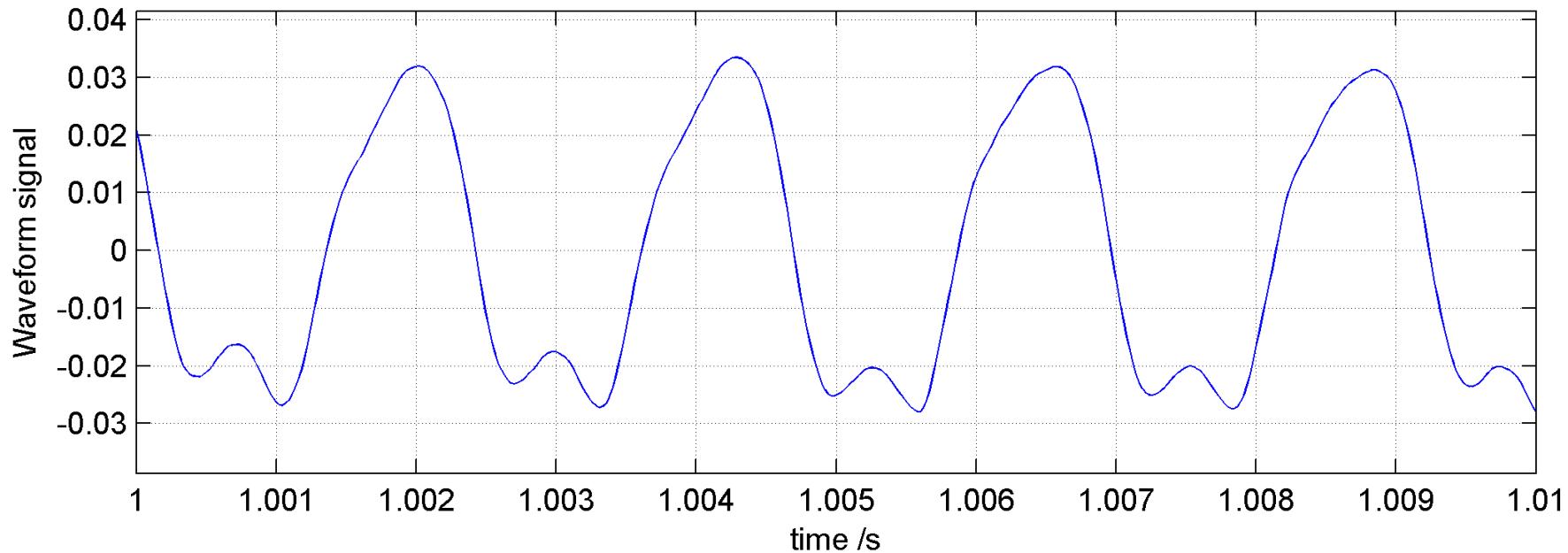


Normalized spectrogram /dB: Frequency spectrum variation with time

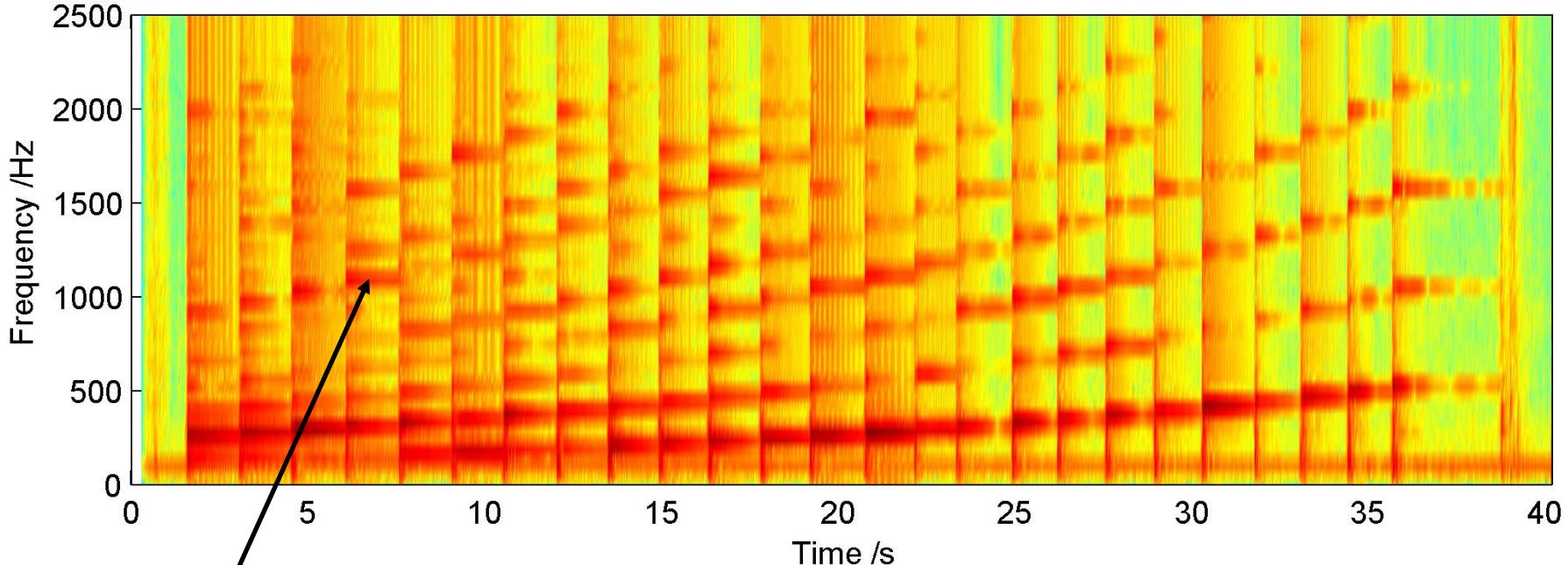


**Flute A 440Hz**

Waveform signal vs time: Right channel

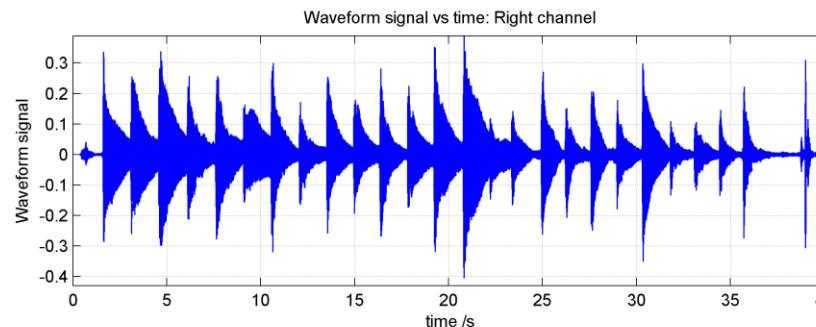
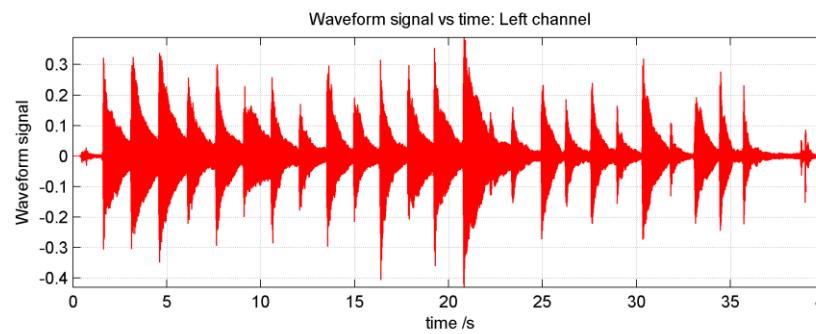


Normalized spectrogram /dB: Frequency spectrum variation with time



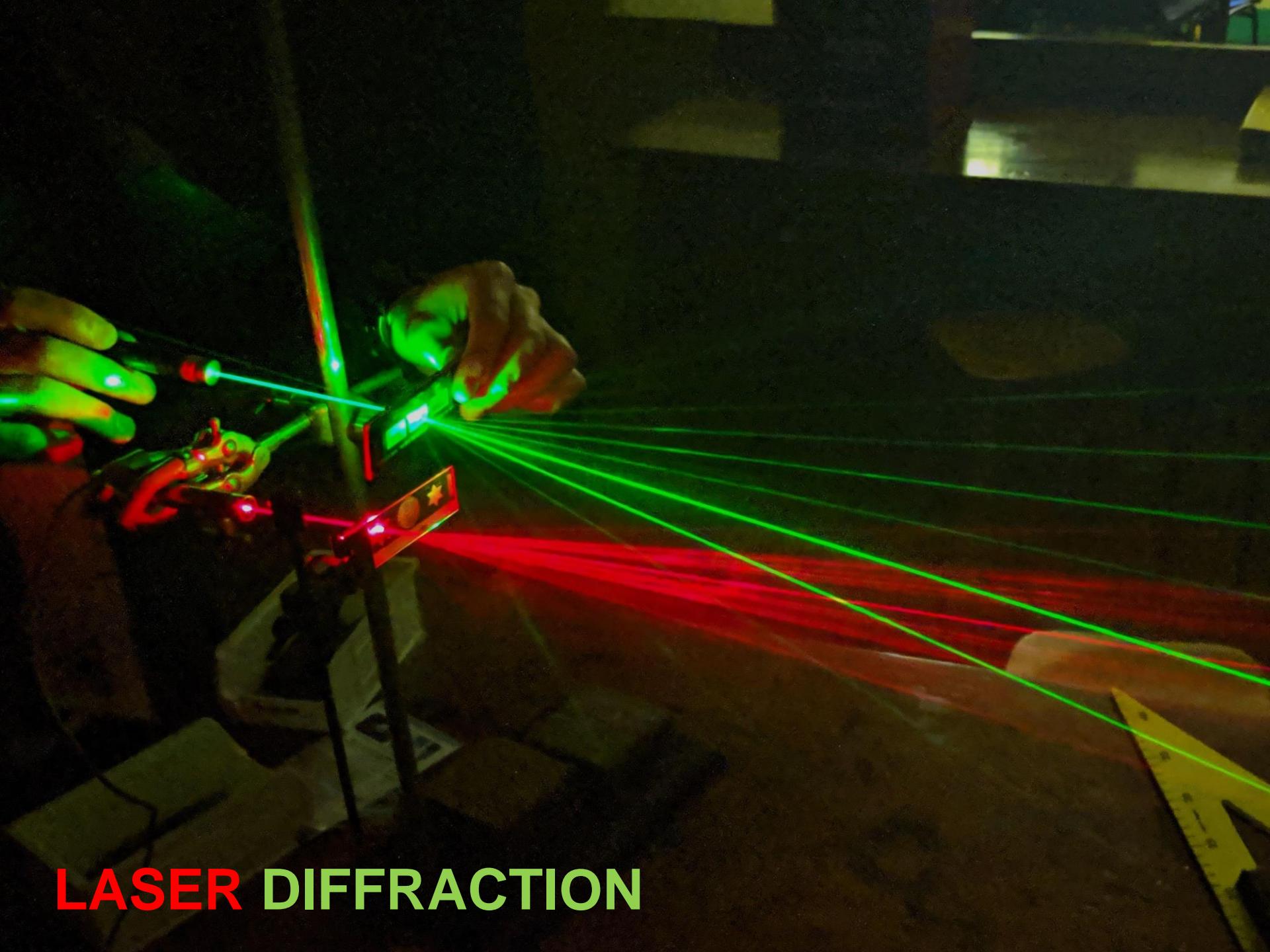
$$f_n = 2^{n/12}$$

Notice  
power law of  
harmonics



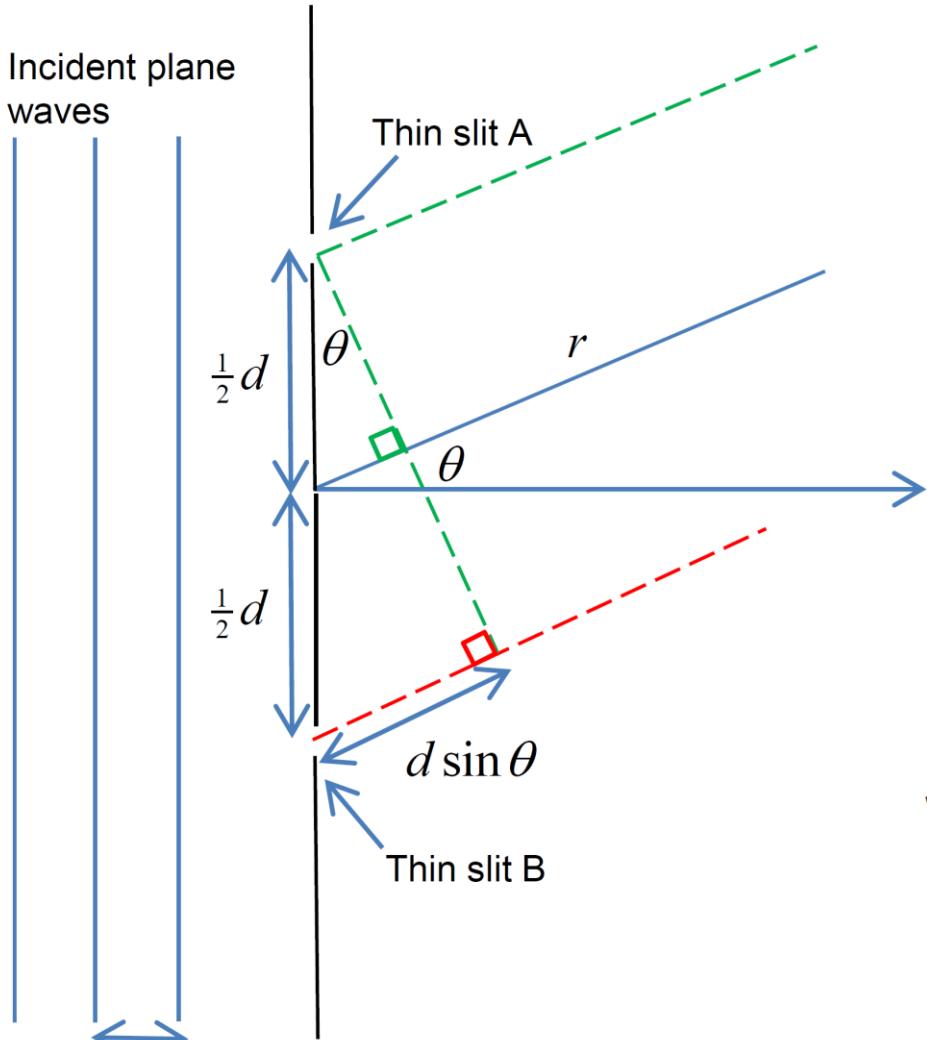
*Chromatic scale (played on a piano) over two octaves*

C,C#...Bb,B,  
C,C#...Bb,B,C



**LASER DIFFRACTION**

## Key geometrical idea from two infinitesimally thin slits ('Young's Slits')



**Aperture**  
(in this case an opaque wall with two infinitesimally thin slits)

Spherical waves will emanate from the slits, and interfere with each other.

For distances such that:

(we call this the **Far Field**)

we can assume waves from each slit are **plane waves**, for any given observational angle  $\theta$ .

**Constructive interference** occurs when the *phase difference* between the waves from slits A and B is an integer multiple of  $2\pi$  radians.

$$\frac{2\pi}{\lambda} d \sin \theta = 2\pi n$$

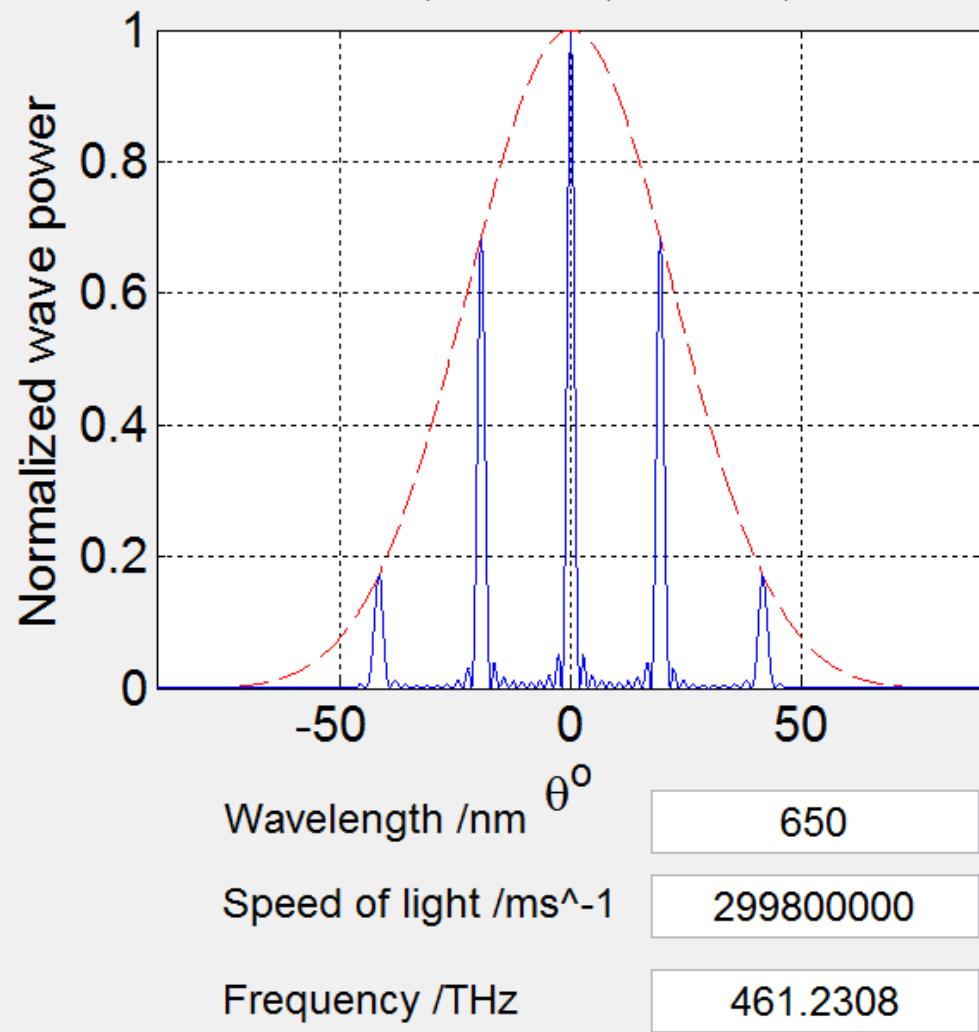
Wavenumber  $k$       Path difference between waves from A and B      Integer  $n$

Hence expect **maxima** in the resulting **Far Field Diffraction pattern** (e.g. spots of a laser on a wall) at angles

$$\sin \theta = \frac{n\lambda}{d}$$

# Grating Fraunhofer far field diffraction

$\lambda = 650\text{nm}$ ,  $s = 3\lambda$ ,  $w = 1\lambda$ ,  $N = 10$



Number  
of slits

10

Slit spacing  
/wavelength

3

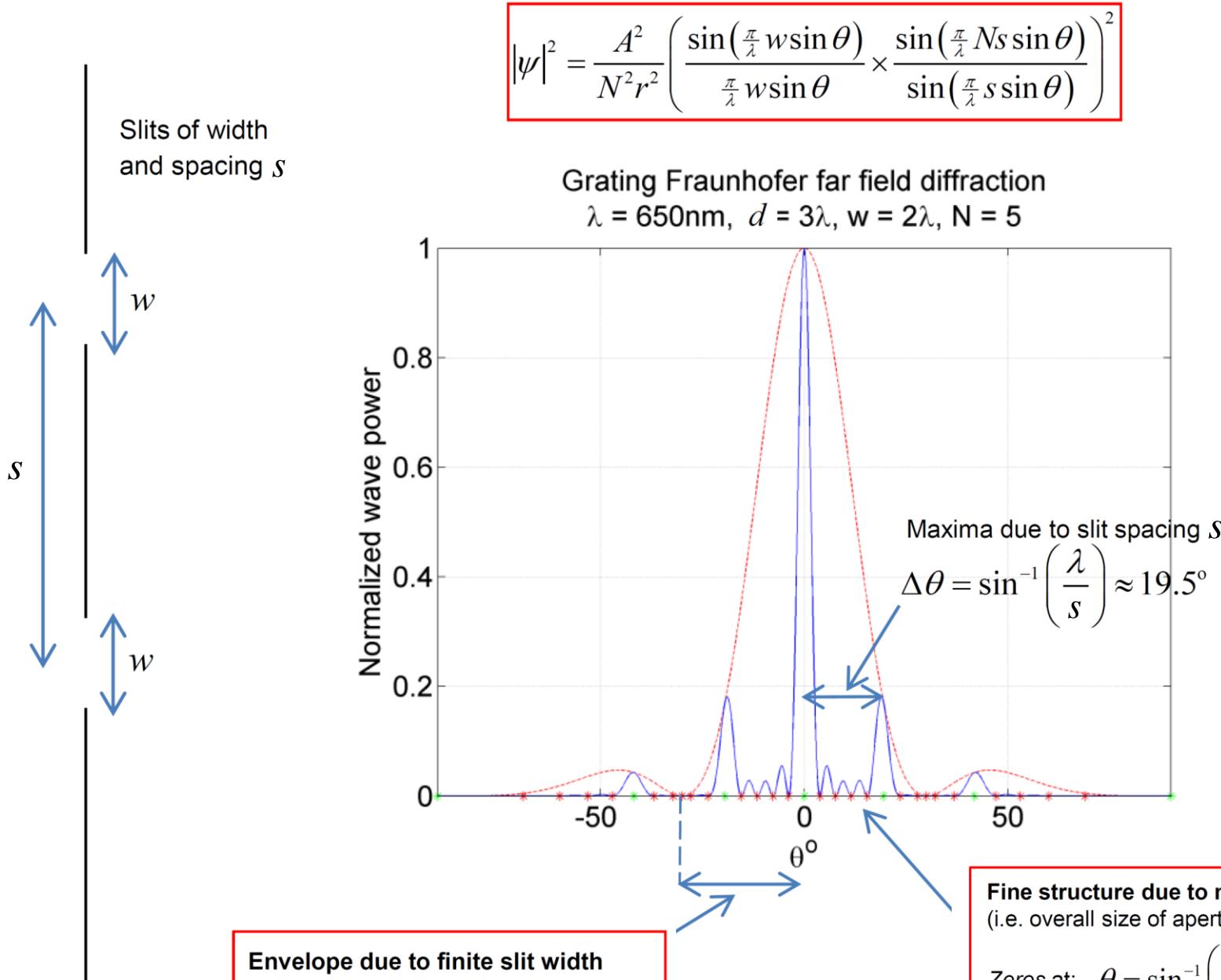
Slit width  
/wavelength

1

Save .PNG

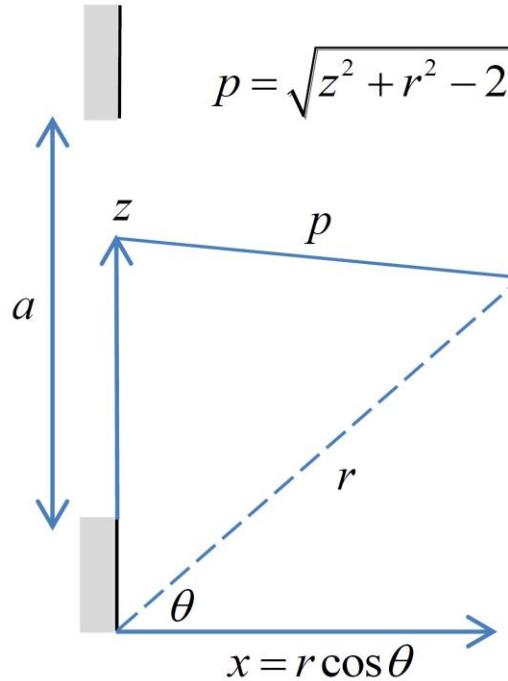
dB

FRAUNHOFER DIFFRACTOR by A. French 2016



## Modelling general diffraction effects from a finite width slit

We can use a computer to evaluate the wavefield in the vicinity of a finite width slit which is uniformly illuminated. We are therefore not restricted to the limitations of the Fraunhofer and Fresnel regimes



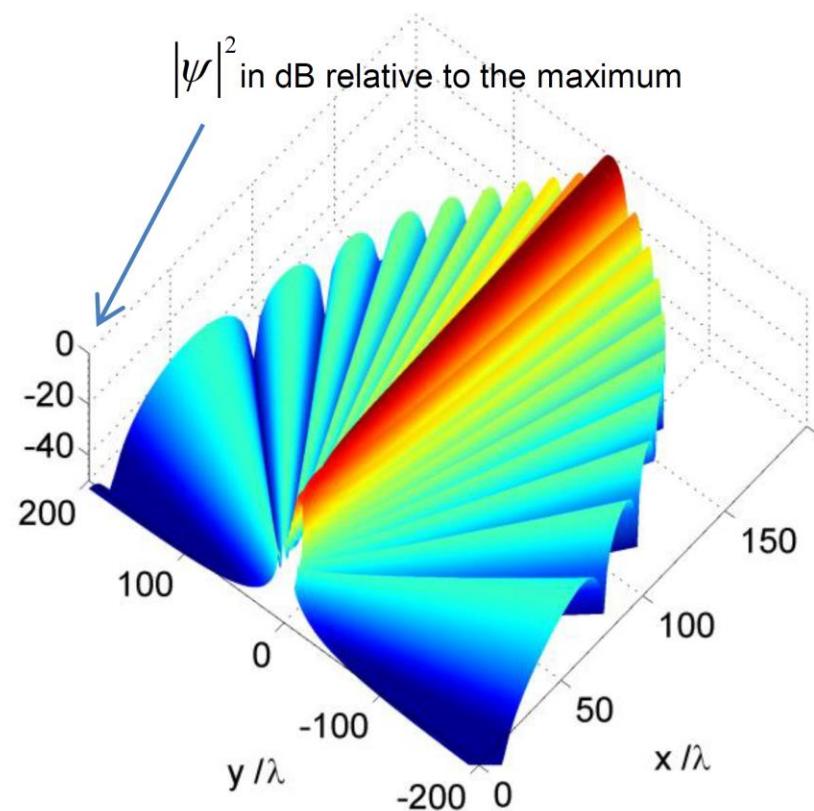
$$p = \sqrt{z^2 + r^2 - 2zr \cos(\frac{1}{2}\pi - \theta)} = \sqrt{z^2 + r^2 - 2zr \sin \theta}$$

$$\psi(r, t) = \frac{A}{a} e^{-i\omega t} \int \frac{e^{ikp}}{p} dz$$

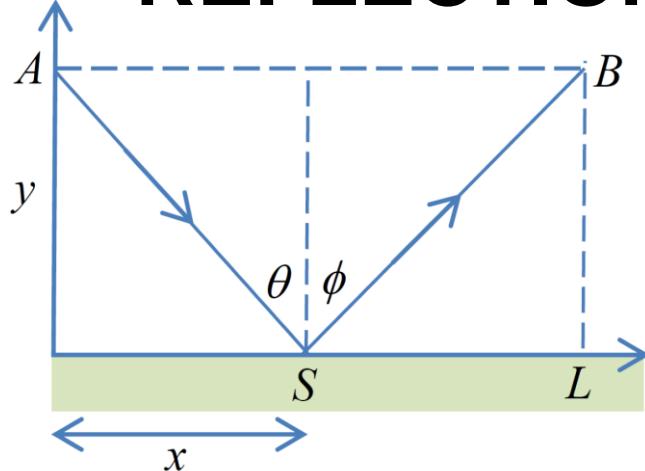
The idea is to evaluate the wavefield power at various  $r$  and  $\theta$  values. The integral is approximated by a sum based on a large number of finite  $dz$  values

In the examples below :

$$dz = \frac{1}{500} a$$



# REFLECTION



$$x = y \tan \theta$$

$$L - x = Y \tan \phi$$

$$t = \frac{\sqrt{x^2 + y^2}}{c/n} + \frac{\sqrt{(L-x)^2 + y^2}}{c/n}$$

$$\frac{\partial t}{\partial x} = \frac{n}{c} \left( \frac{\frac{1}{2} 2x}{\sqrt{x^2 + y^2}} + \frac{\frac{1}{2} 2(L-x)(-1)}{\sqrt{(L-x)^2 + y^2}} \right) \quad \frac{1}{\cos^2 \theta} = 1 + \tan^2 \theta$$

$$\tan \theta = \frac{\sin \theta}{\cos \theta}$$

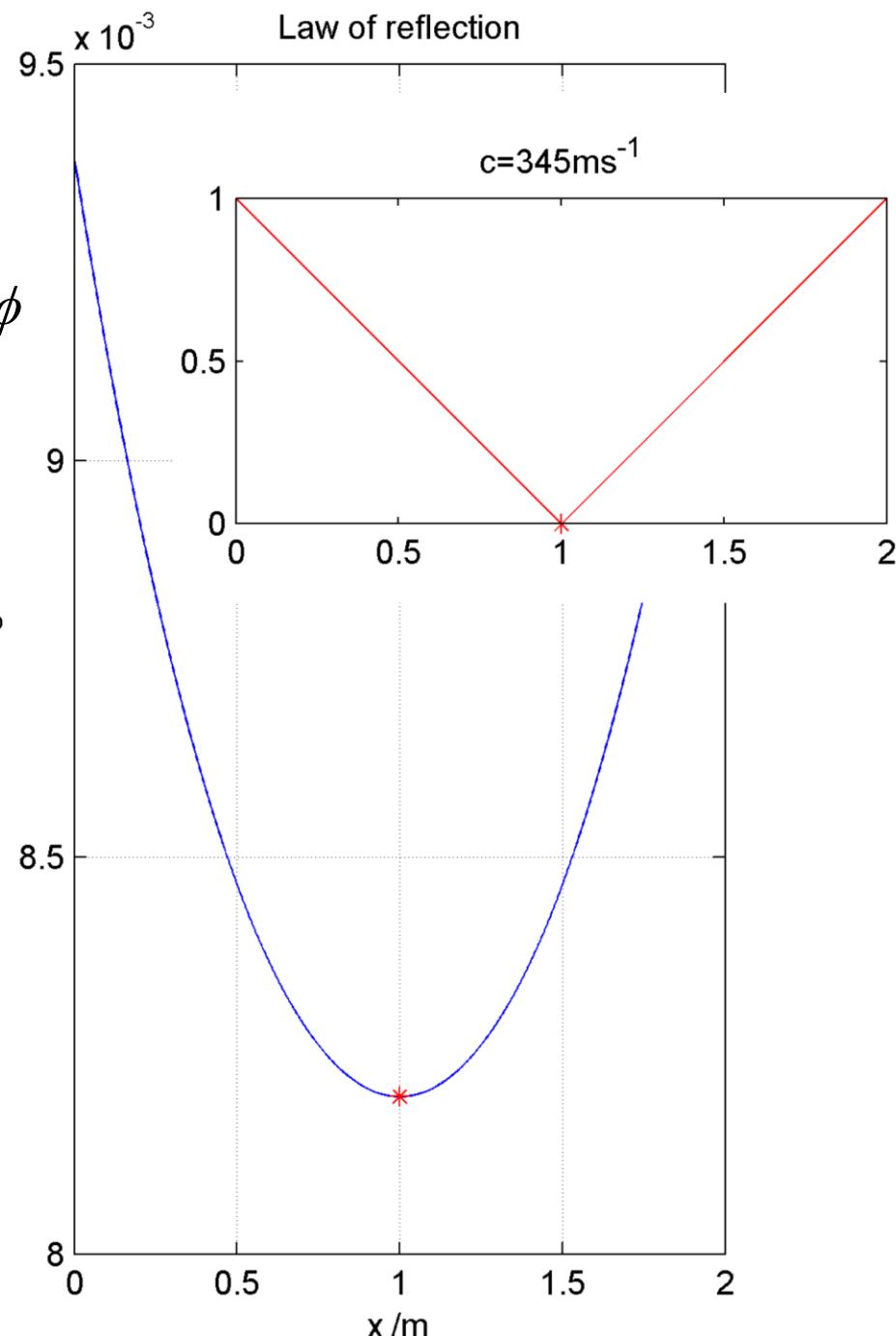
$$\therefore \frac{\partial t}{\partial x} = \frac{n}{c} \left( \frac{y \tan \theta}{\sqrt{y^2 \tan^2 \theta + y^2}} - \frac{y \tan \phi}{\sqrt{y^2 \tan^2 \phi + y^2}} \right)$$

$$\therefore \frac{\partial t}{\partial x} = \frac{n}{c} (\sin \theta - \sin \phi)$$

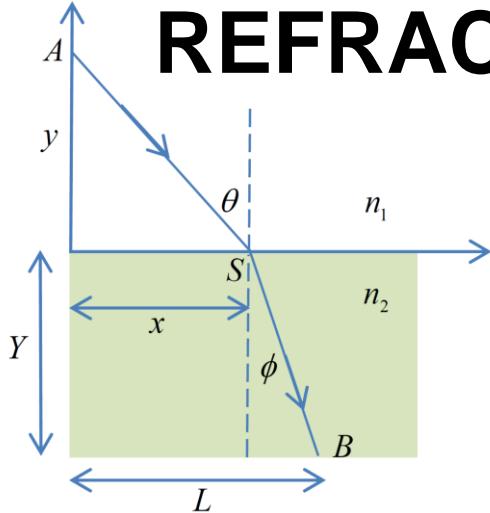
The travel time is minimized when  $\frac{\partial t}{\partial x} = 0$

$$\therefore \sin \theta = \sin \phi$$

$$\therefore \theta = \phi$$



# REFRACTION



$$x = y \tan \theta$$

$$L - x = Y \tan \phi$$

$$t = \frac{\sqrt{x^2 + y^2}}{c/n_1} + \frac{\sqrt{(L-x)^2 + Y^2}}{c/n_2}$$

$$\frac{\partial t}{\partial x} = \frac{1}{c} \left( \frac{\frac{1}{2} 2x n_1}{\sqrt{x^2 + y^2}} + \frac{\frac{1}{2} 2(L-x)(-1)n_2}{\sqrt{(L-x)^2 + Y^2}} \right)$$

$$\frac{\partial t}{\partial x} = \frac{1}{c} \left( \frac{y \tan \theta n_1}{y \sqrt{\tan^2 \theta + 1}} + \frac{-Y \tan \phi n_2}{Y \sqrt{\tan^2 \phi + 1}} \right)$$

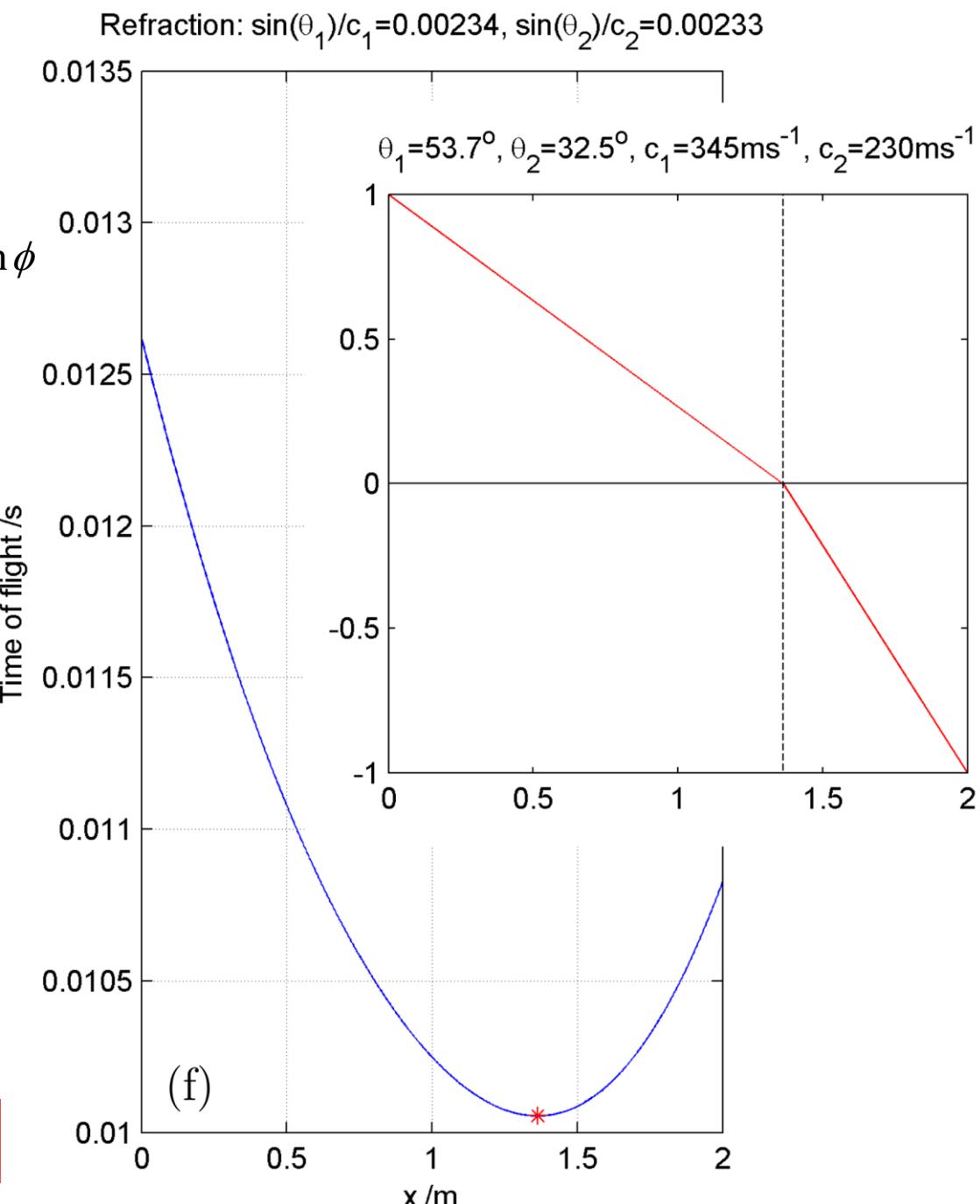
$$\frac{\partial t}{\partial x} = \frac{1}{c} (\cos \theta \tan \theta n_1 - \cos \phi \tan \phi n_2)$$

$$\frac{\partial t}{\partial x} = \frac{1}{c} (n_1 \sin \theta - n_2 \sin \phi)$$

The travel time is minimized when

$$\frac{\partial t}{\partial x} = 0$$

$$\therefore n_1 \sin \theta = n_2 \sin \phi$$



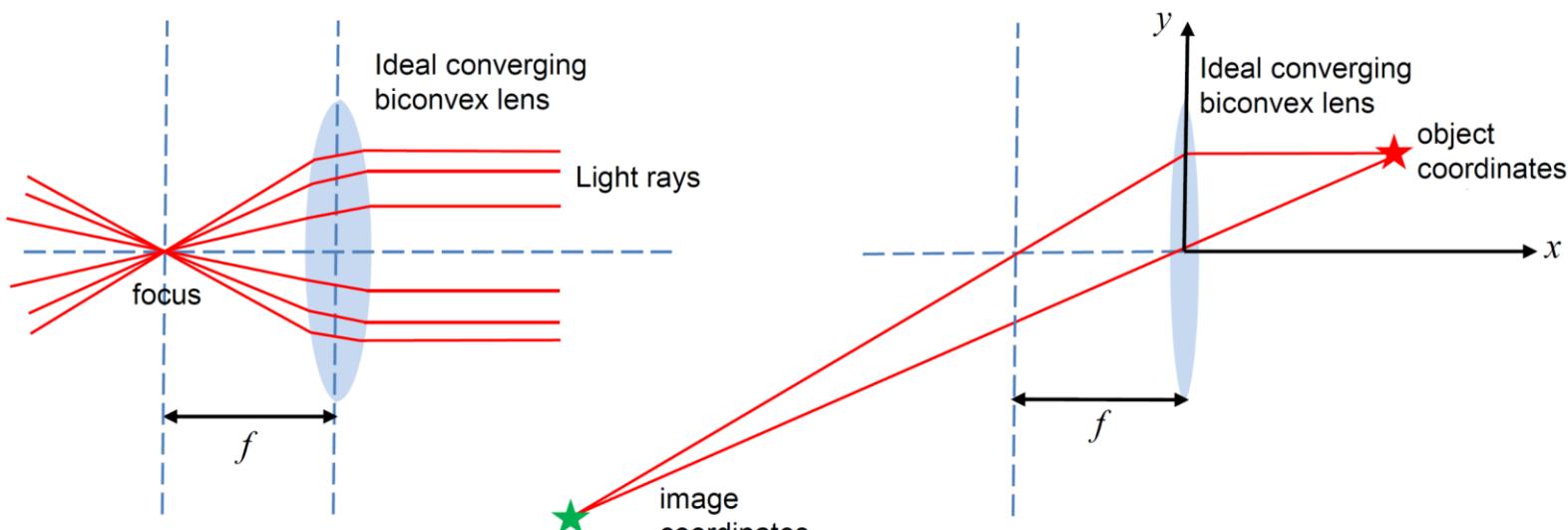
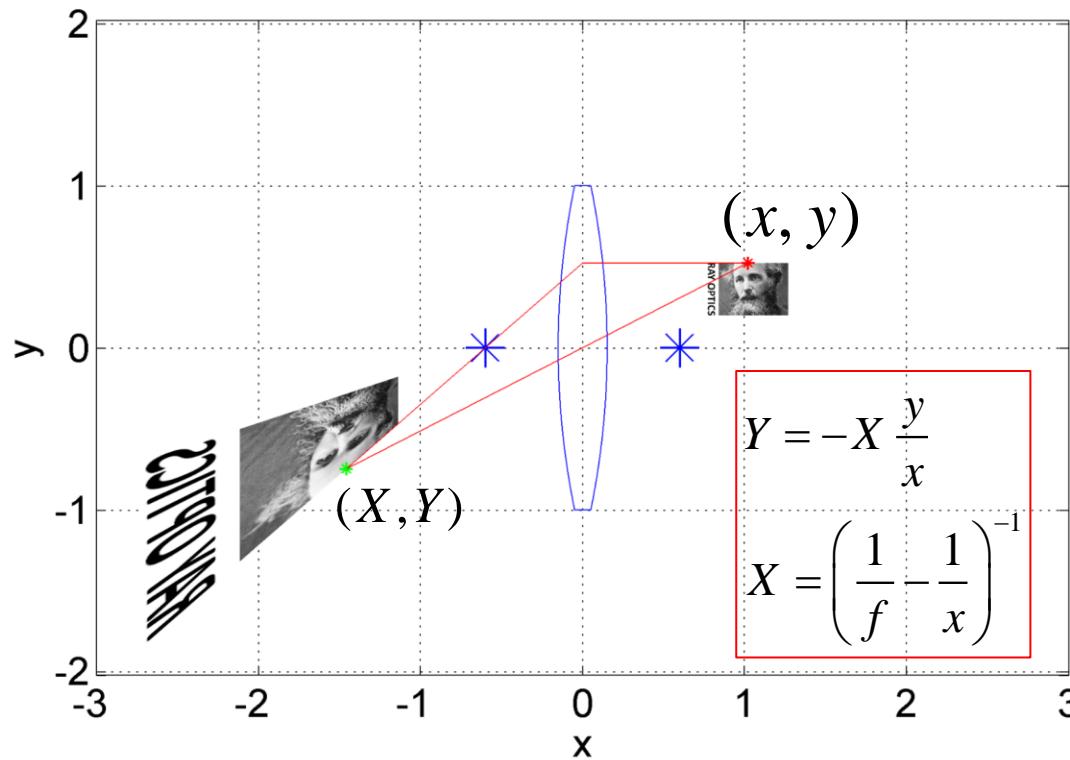
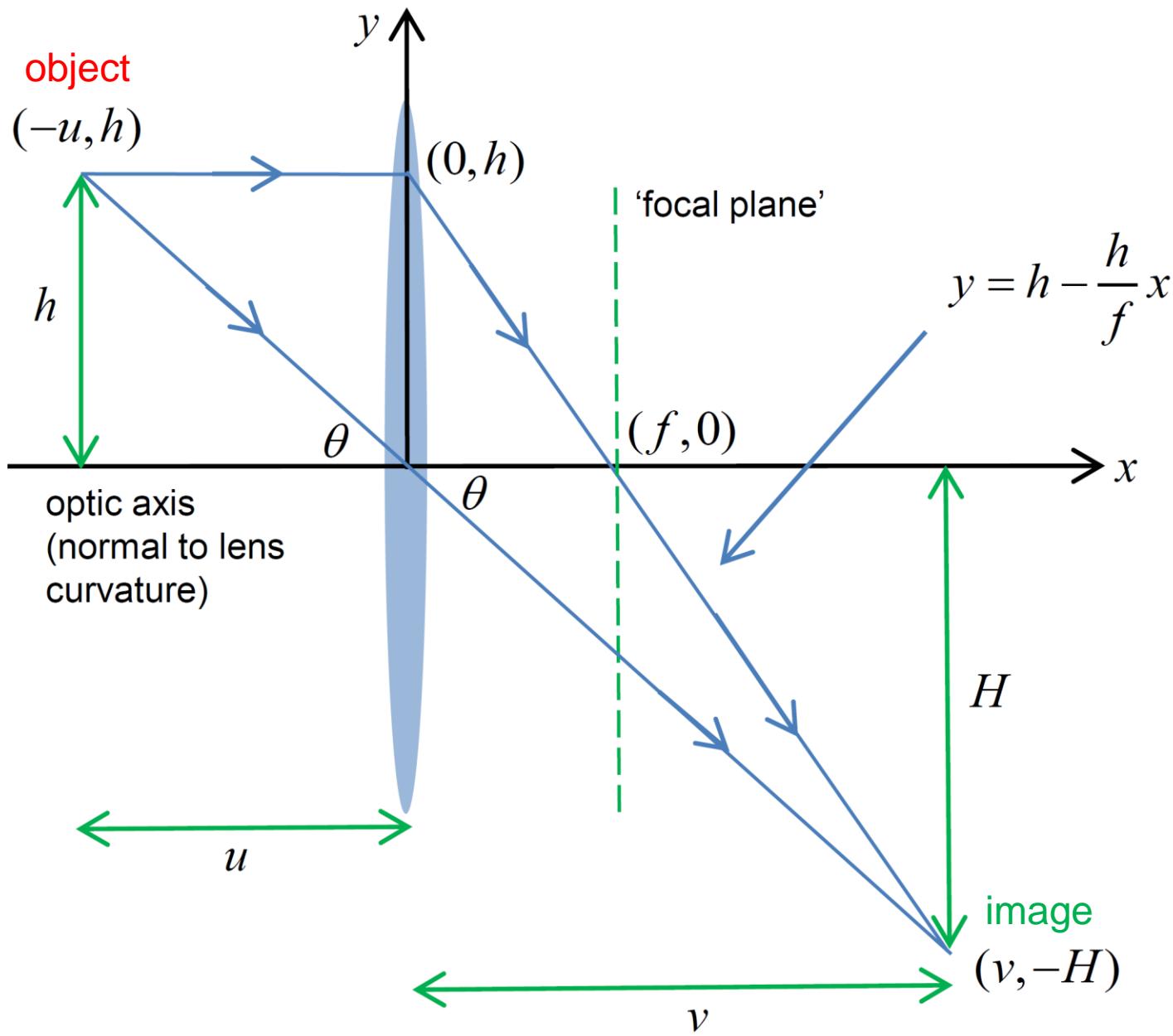


Image of object in a converging lens



# RAY OPTICS: THIN BICONVEX LENS

# THIN LENS EQUATION



$$\frac{1}{u} + \frac{1}{v} = \frac{1}{f}$$

$$\frac{H}{v} = \frac{h}{u}$$

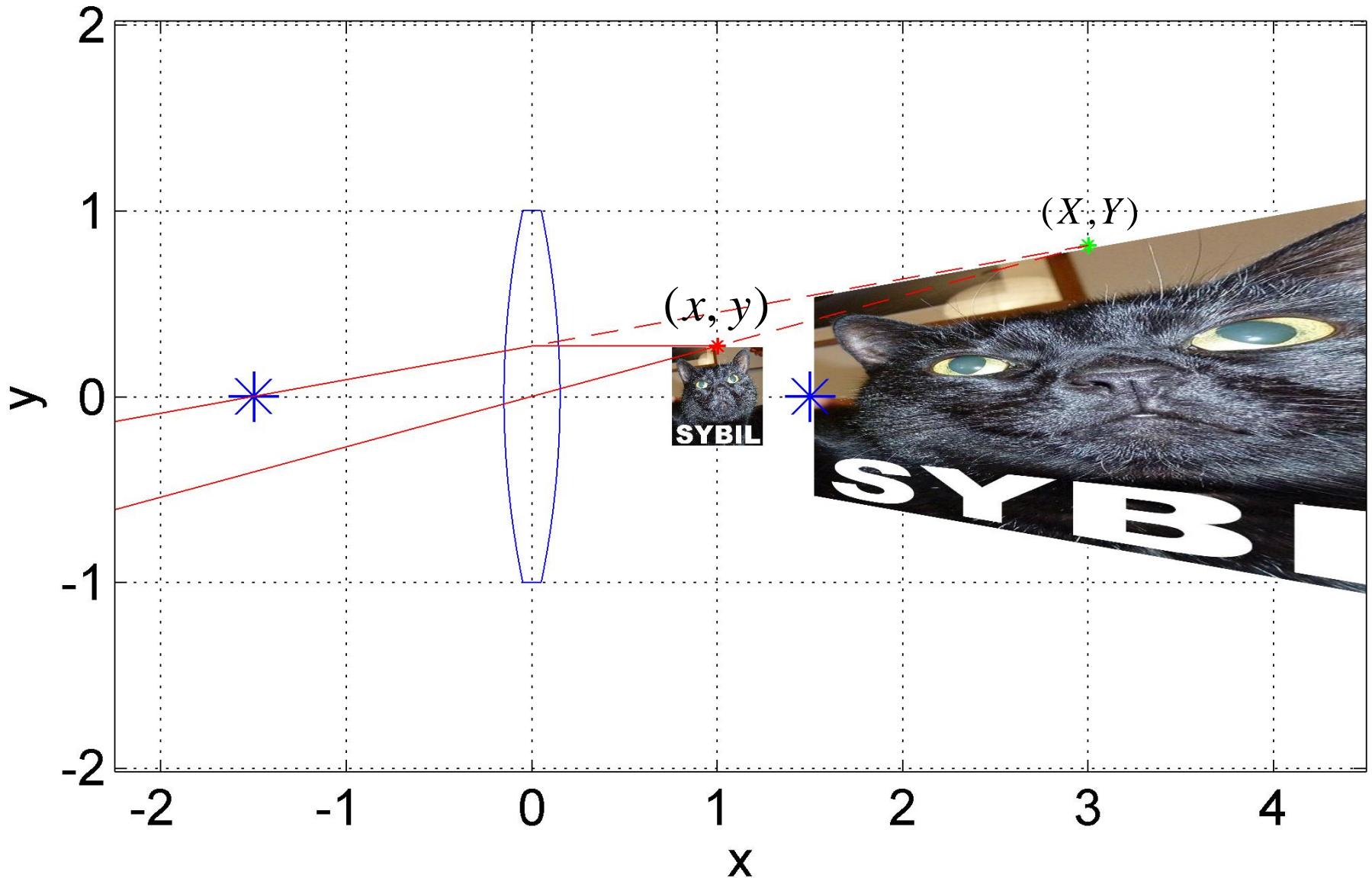
$$v = \left( \frac{1}{f} - \frac{1}{u} \right)^{-1}$$

$$Y = -X \frac{y}{x}$$

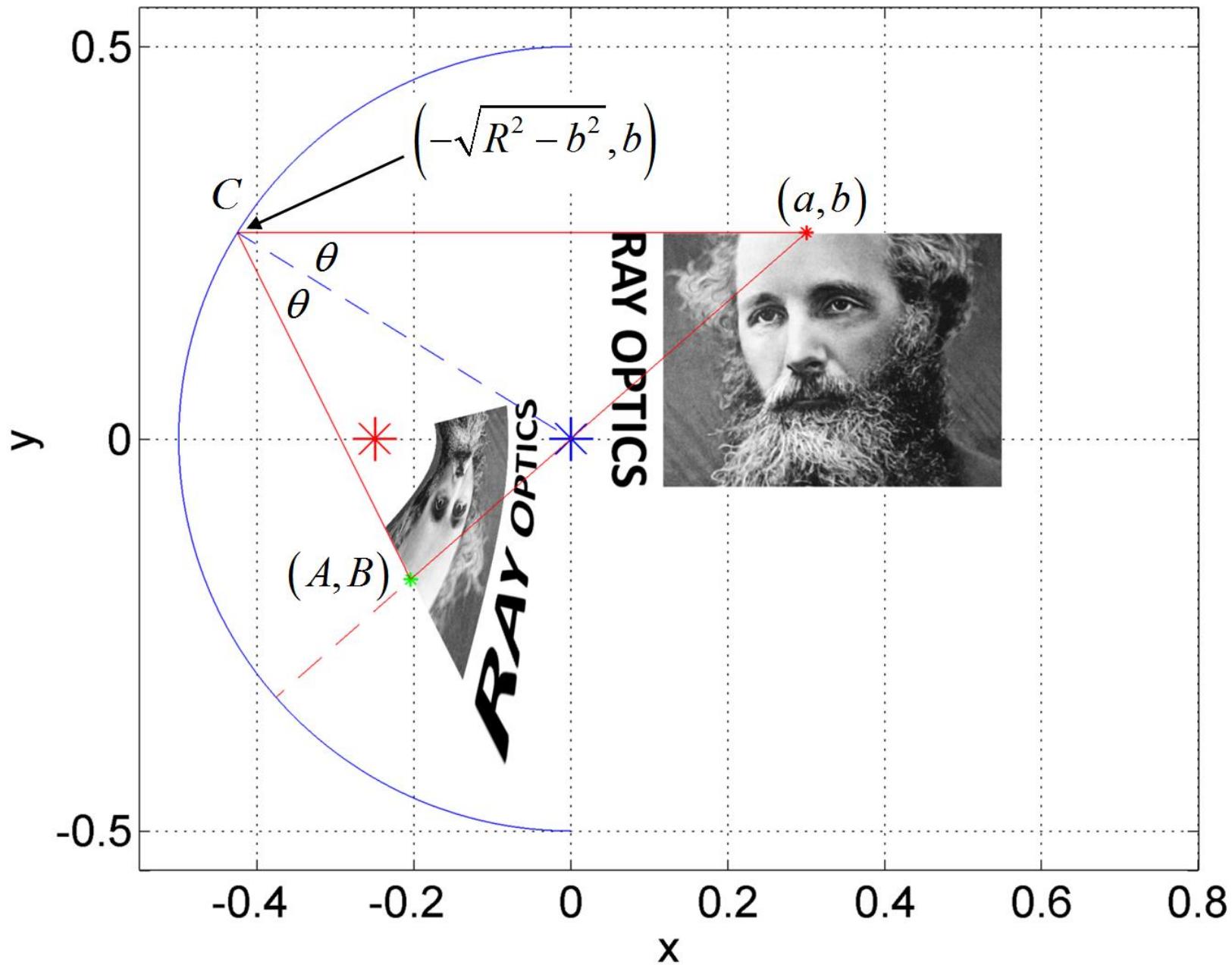
$$X = \left( \frac{1}{f} - \frac{1}{x} \right)^{-1}$$

Image is a  
*coordinate  
transformation*

# Virtual image of object in a magnifying lens



# Reflection in a concave mirror



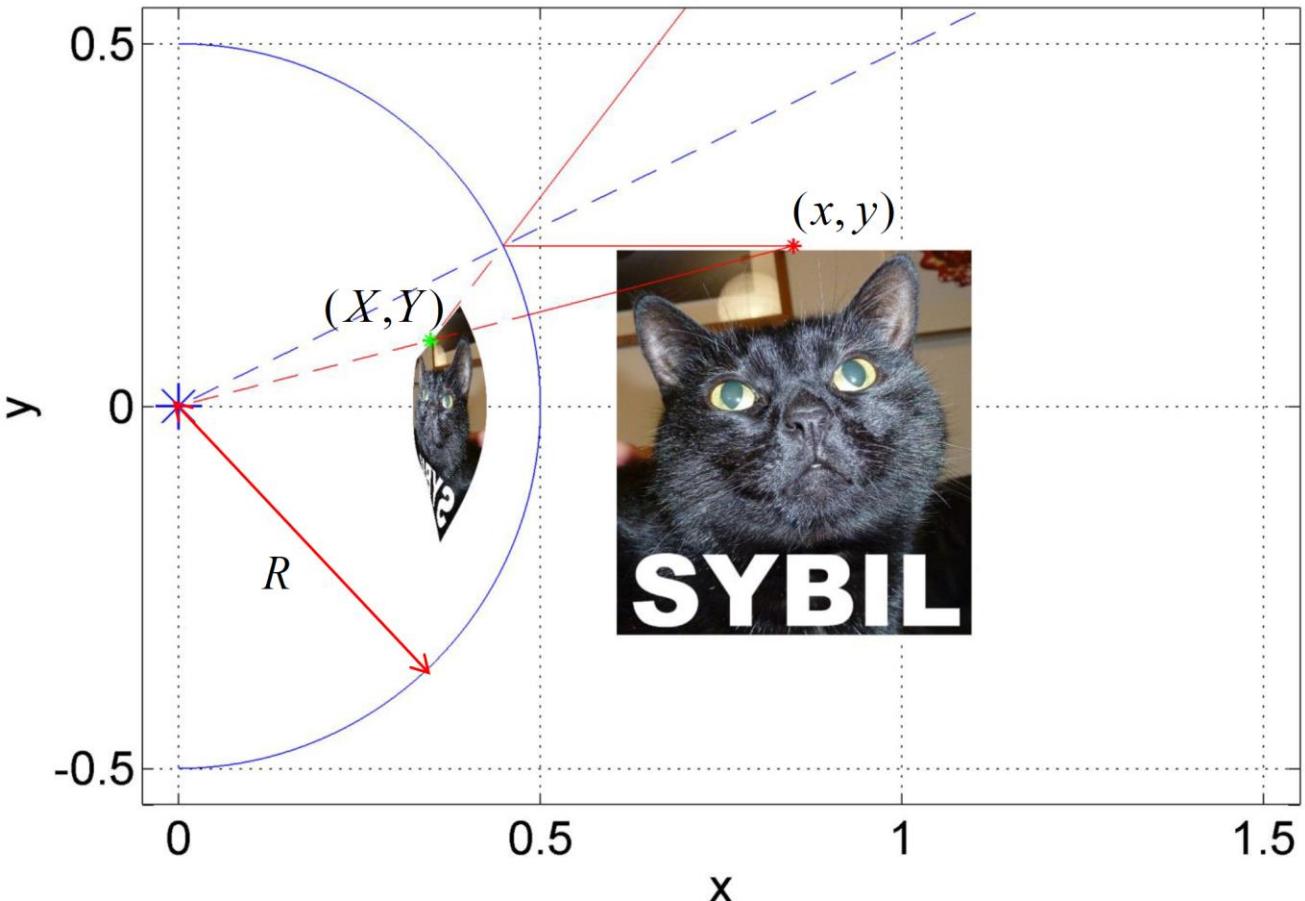
(a)



(b)



## Reflection in a convex mirror



We see an upright, distorted *virtual* image in a cylindrical mirror.

$$\alpha = \frac{1}{2} \tan^{-1} \left( \frac{y}{x} \right)$$

$$k = \frac{x}{\cos(2\alpha)}$$

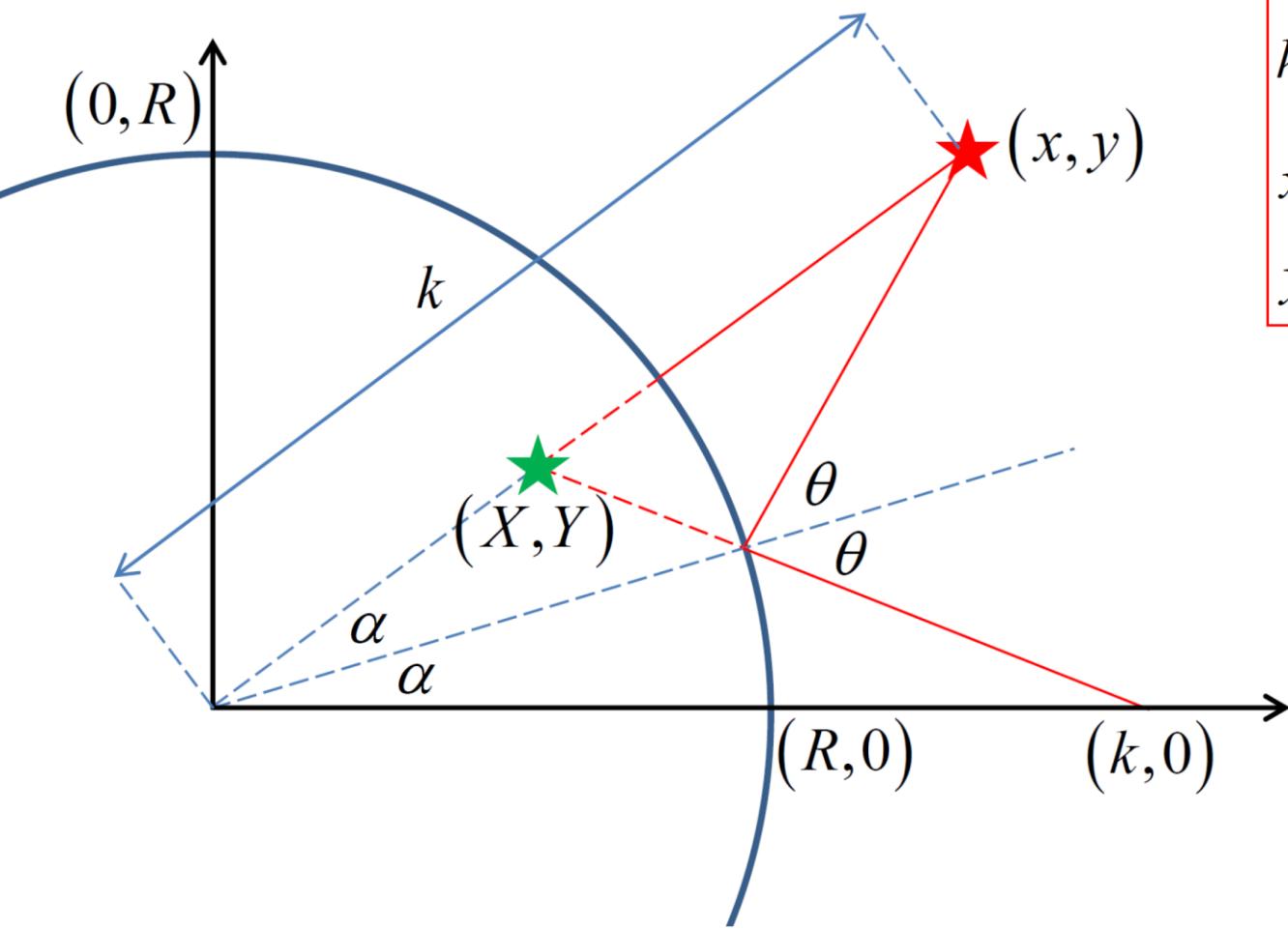
$$Y = \frac{k \sin \alpha}{\frac{k}{R} - \cos \alpha + \frac{x}{y} \sin \alpha}$$

$$X = x - \frac{Y}{y}$$

Virtual  
image from  
object  
coordinates

i.e. the *apparent source* of  
(diverging) light rays from  
the mirror

# **Convex mirror** object to virtual image transformation (and reverse transformation)



$$\alpha = \frac{1}{2} \tan^{-1} \left( \frac{Y}{X} \right)$$

$$k = \frac{R(Y \cos \alpha - X \sin \alpha)}{Y - R \sin \alpha}$$

$$x = k \cos(2\alpha)$$

$$y = k \sin(2\alpha)$$

$$\alpha = \frac{1}{2} \tan^{-1} \left( \frac{y}{x} \right)$$

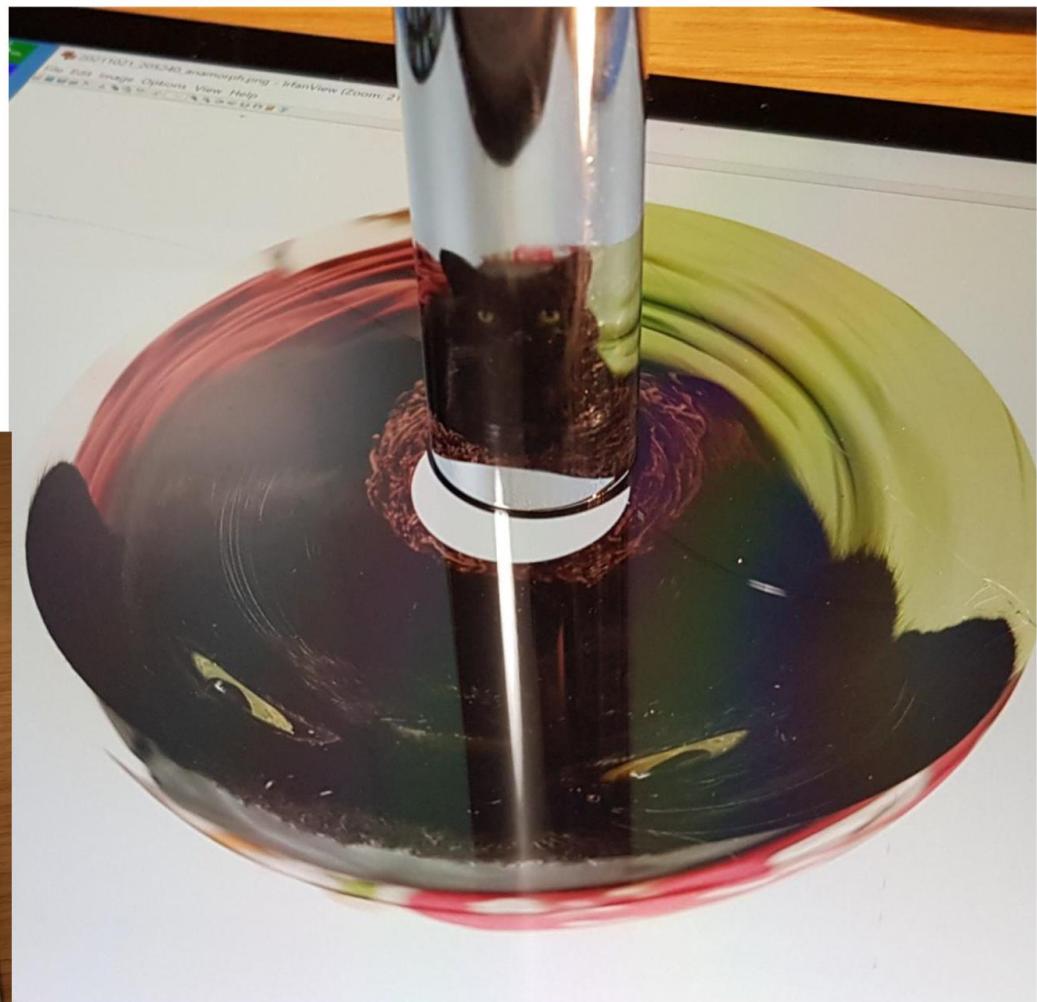
$$k = \frac{x}{\cos(2\alpha)}$$

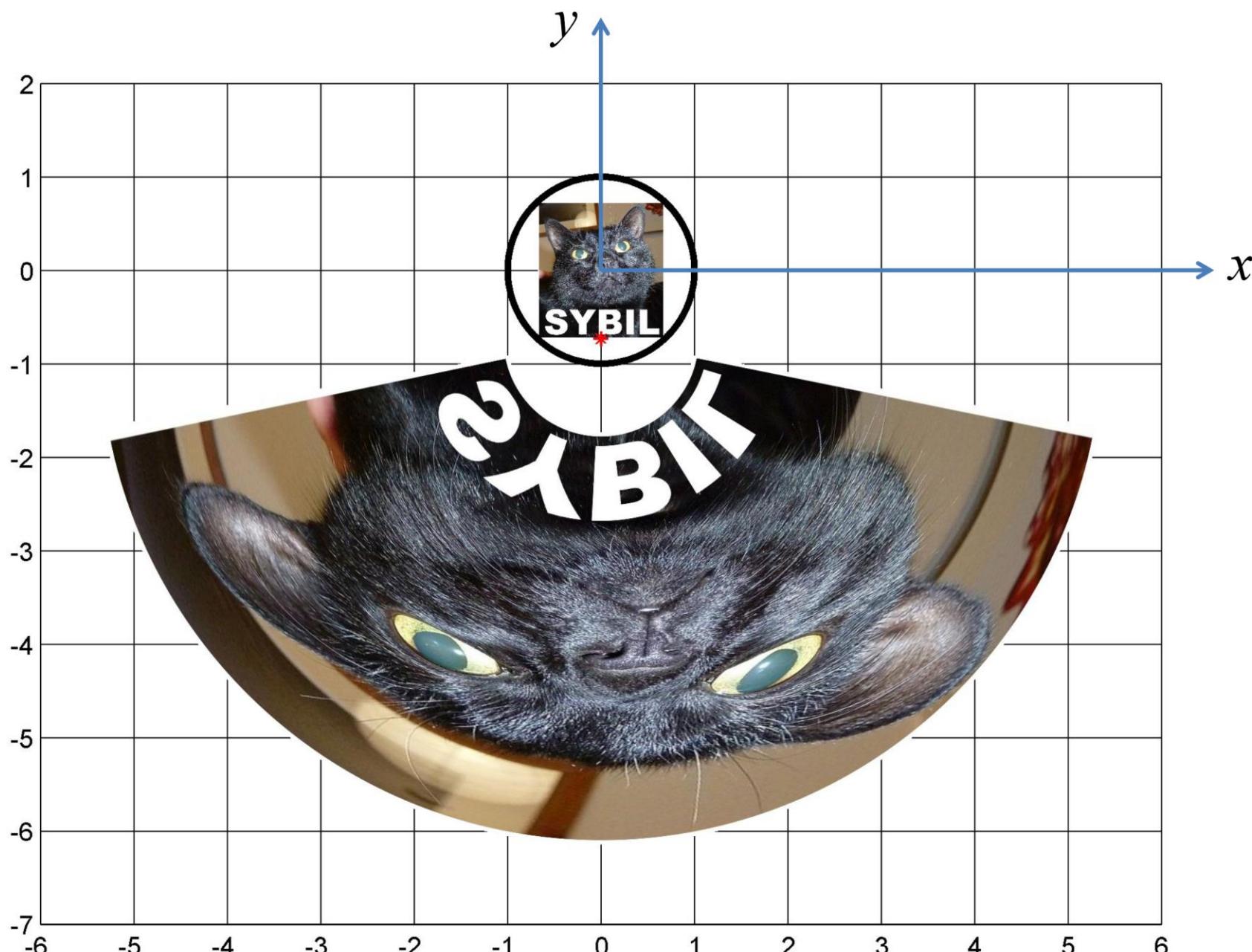
$$Y = \frac{k \sin \alpha}{\frac{k}{R} - \cos \alpha + \frac{x}{y} \sin \alpha}$$

$$X = x - \frac{Y}{y}$$



Sybil the cat was unperturbed by this anamorphic transformation.

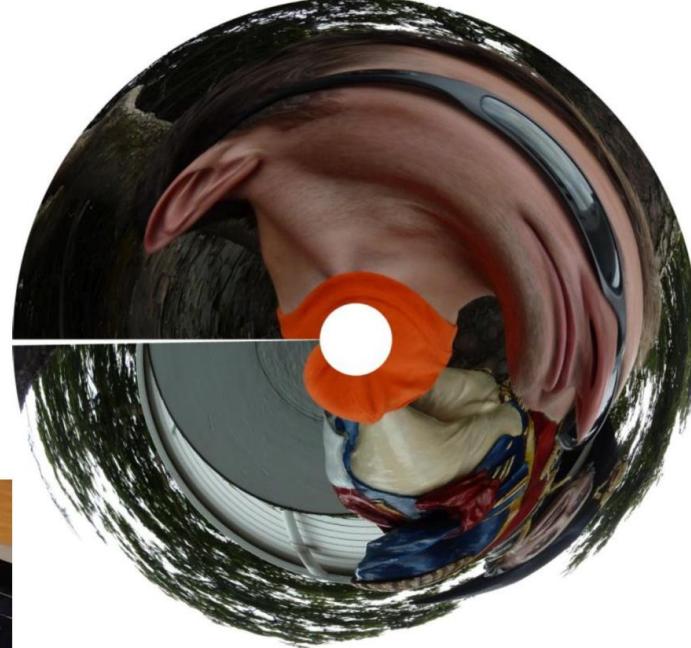
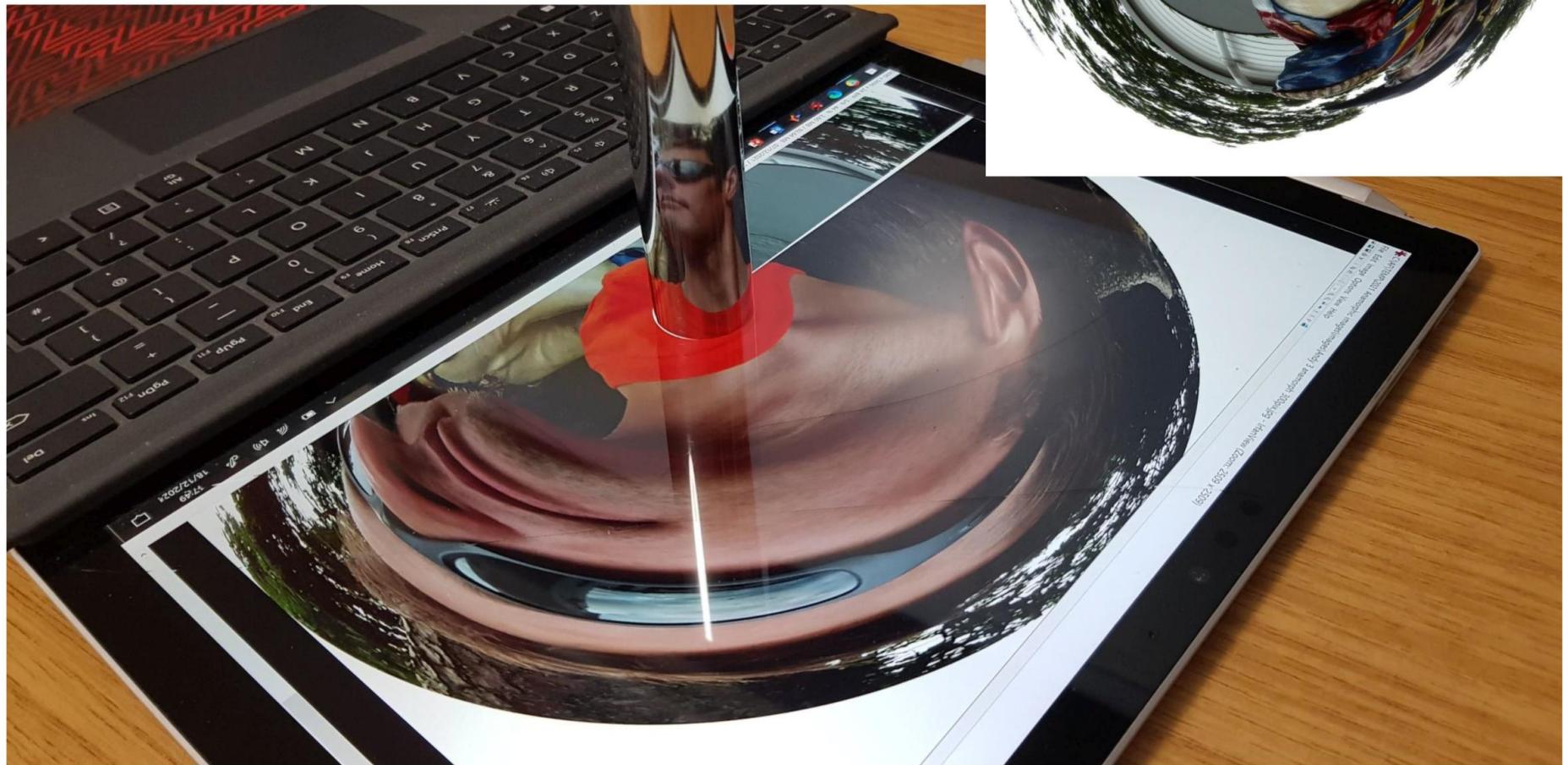




arc\_deg = 160

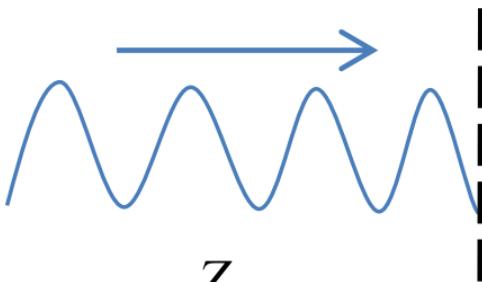
Rf = 3

'Fear and loathing' in Portmeirion ...



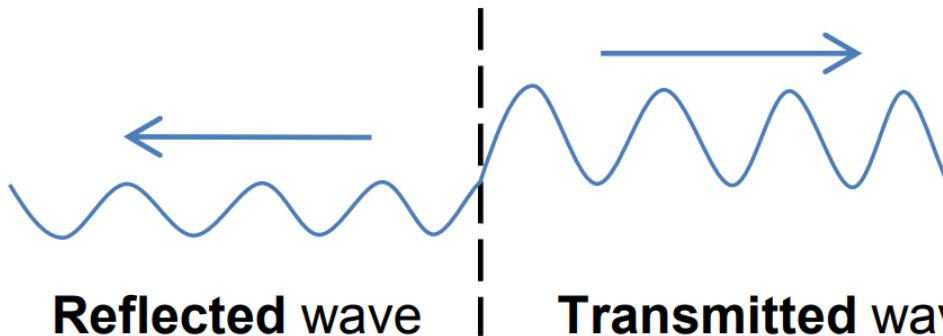
# Reflection and transmission of waves on boundaries

$$Z = \rho c \quad \text{acoustic impedance}$$



**Incident wave**

$$\psi_I = A_I e^{i(k_1 x - \omega t)}$$



**Reflected wave**

$$\psi_R = A_R e^{i(-k_1 x - \omega t)}$$

$Z_1$

**Transmitted wave**

$$\psi_T = A_T e^{i(k_2 x - \omega t)}$$

$Z_2$

$$\frac{A_R}{A_I} = \frac{Z_1 - Z_2}{Z_2 + Z_1}$$

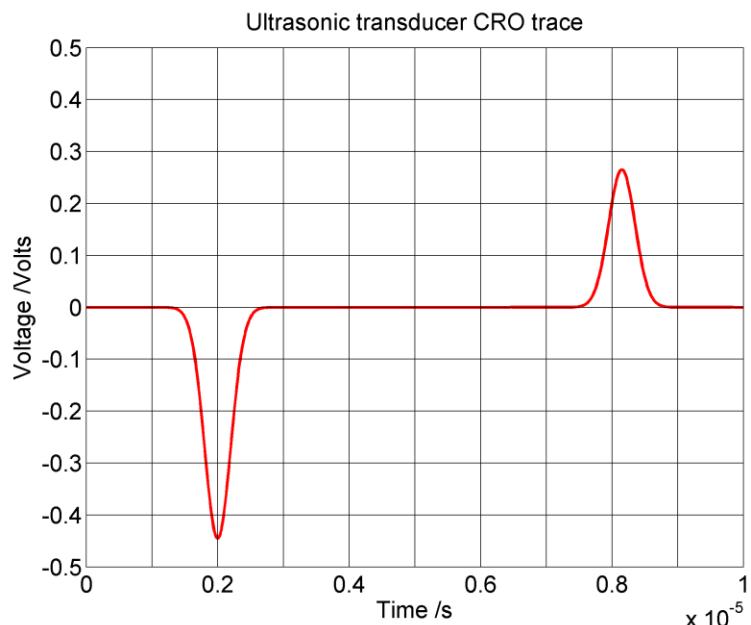
$$\frac{A_T}{A_I} = \frac{2Z_1}{Z_2 + Z_1}$$

$$P = \frac{1}{2} Z A^2 \omega^2$$

Wave power

$$\frac{P_R}{P_I} = \left| \frac{Z_1 - Z_2}{Z_1 + Z_2} \right|^2$$

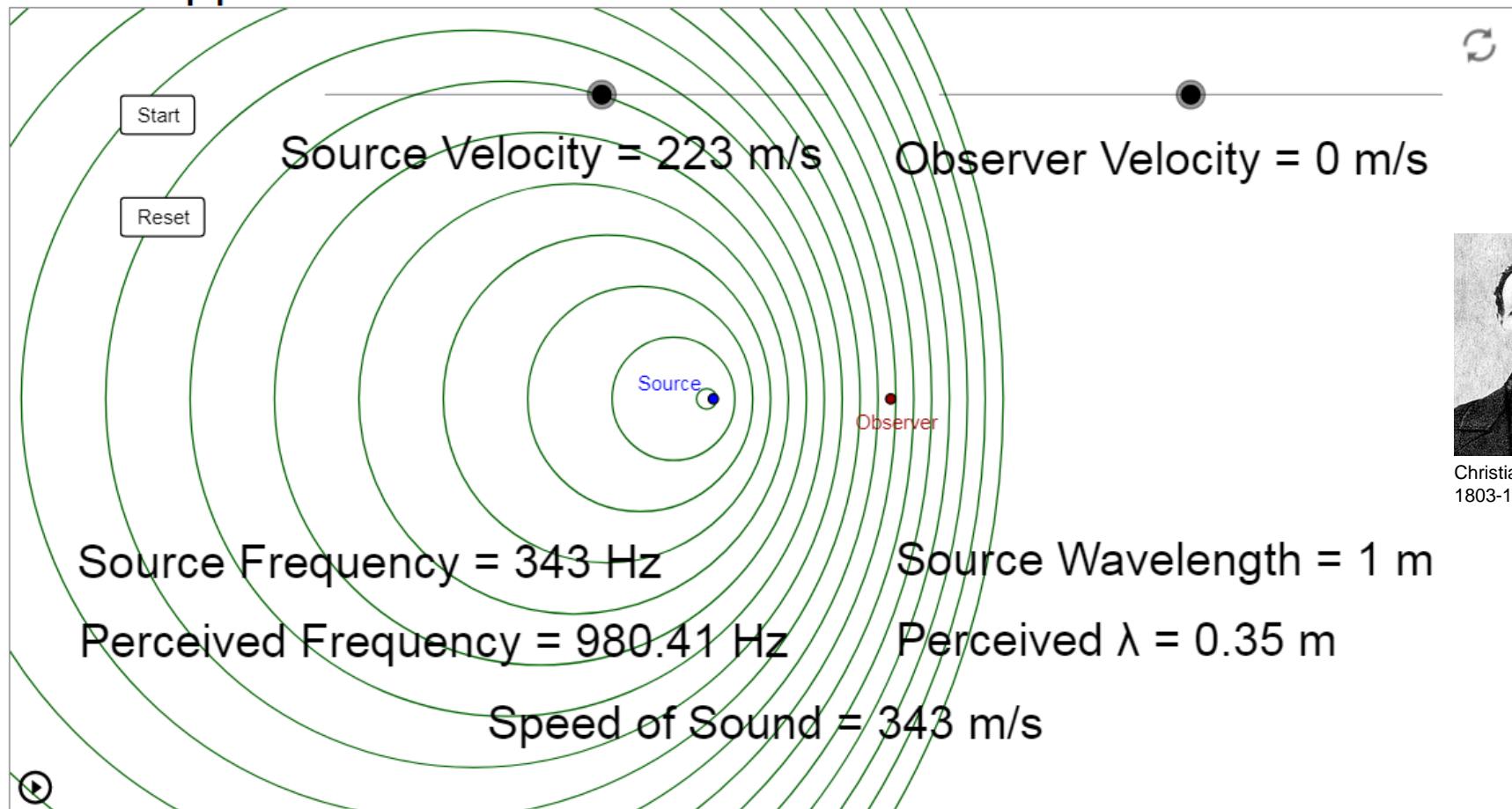
$$\frac{P_T}{P_I} = 1 - \left| \frac{Z_1 - Z_2}{Z_1 + Z_2} \right|^2$$



# *oPhysics: Interactive Physics Simulations*

Home Kinematics Forces Conservation Waves Light E & M Rotation Fluids Modern Drawing Tools Fun Stuff

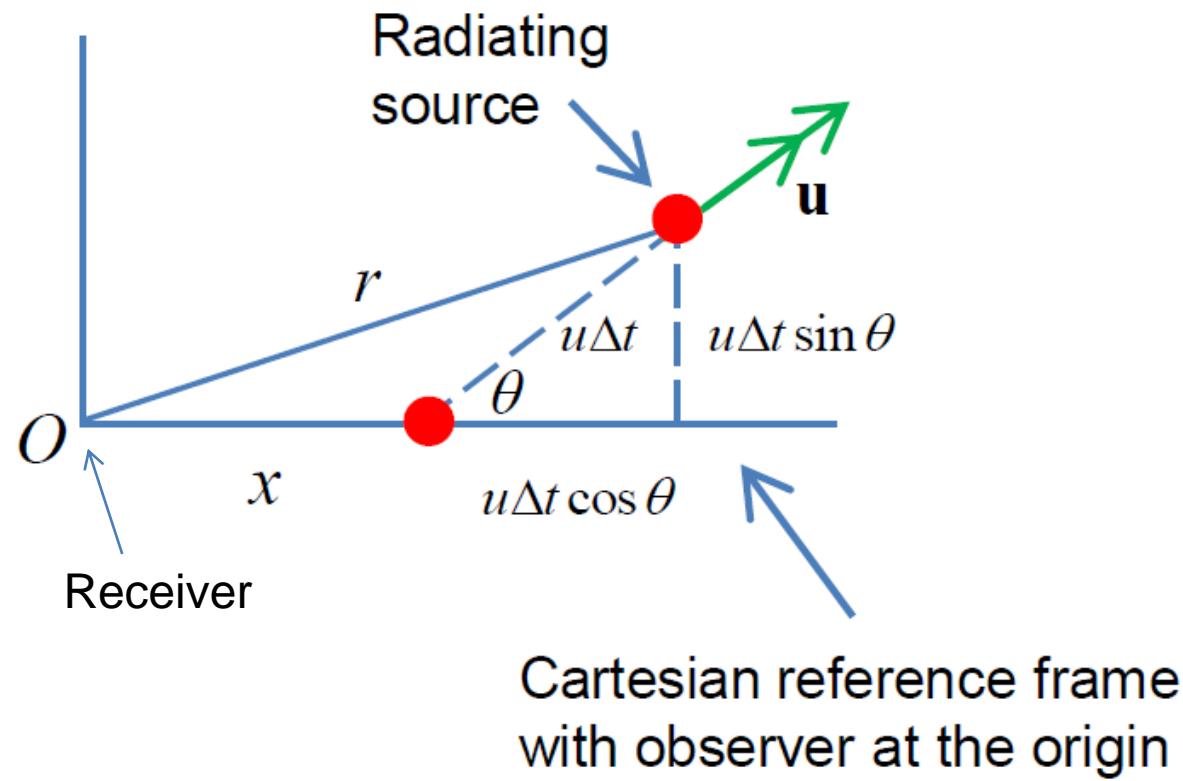
## The Doppler Effect & Sonic Boom



Christian Doppler  
1803-1853

Consider a **receding** wave source of frequency  $f$ . It crosses the  $x$  axis of a Cartesian reference frame at angle  $\theta$  with speed  $u$ . The receiver of the waves is stationary at the origin of the Cartesian frame.

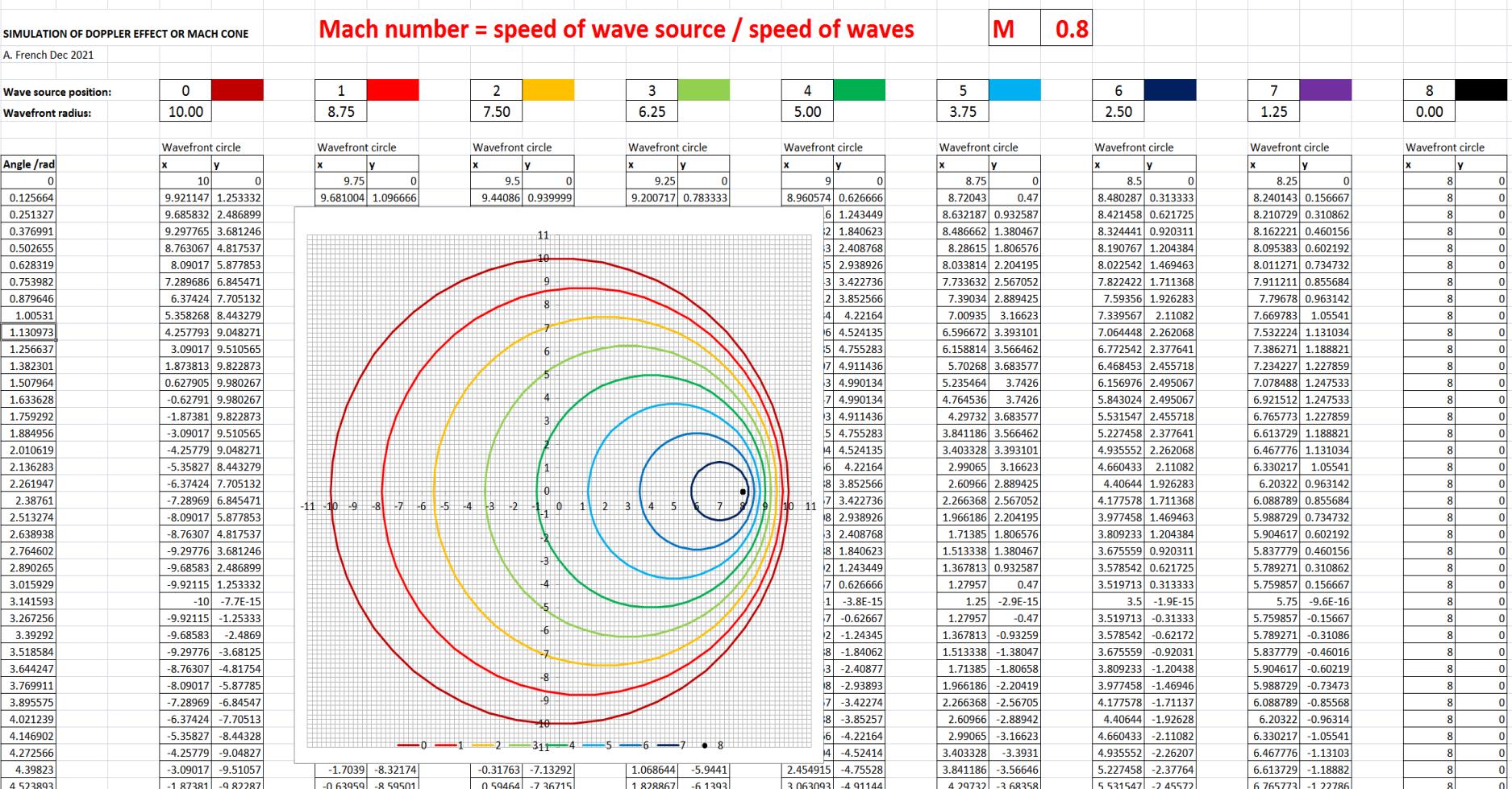
The speed of waves, relative to the observer, is  $w$ .



Depending on the velocity of the wave source relative to the observer, the observer will experience a *frequency shift* from  $f$ . If the source *recedes*, the frequency *diminishes* and the *wavelength increases* ('**redshift**'). If the source is *approaching*, the observed *frequency will increase* and the *wavelength will decrease* ('**blueshift**').

# Excel model of Doppler Shift

- Plot the circular wavefront(s) when the source has got to position 8.
- i.e. you see a wavefront radius which increases as the *initial position* decreases, and also a *shift of wavefront centre*. The combination of these two effects causes the bunching up of wavefronts

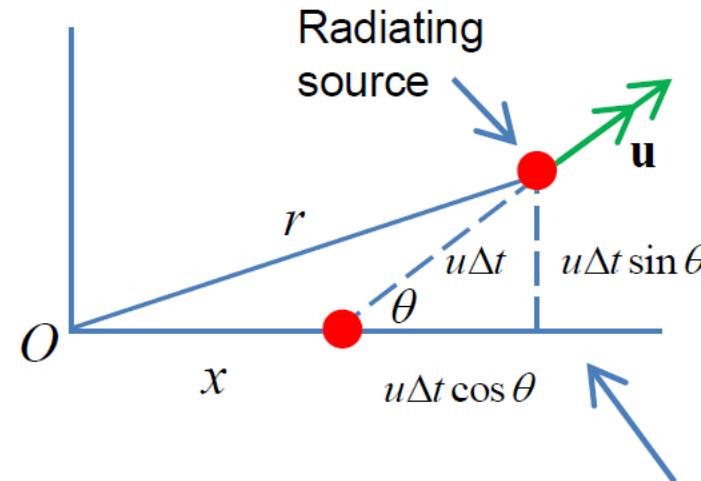


The period  $T$  of waves received by an observer (in the  $x$  direction) at the frame origin  $O$  is:

$$T = \Delta t + \frac{r - x}{w}$$

time between wave crests at source      wave speed

extra distance travelled by source between wave crests



From geometry:

$$r = \sqrt{(x + u\Delta t \cos \theta)^2 + u^2 \Delta t^2 \sin^2 \theta}$$

$$r = \sqrt{x^2 + u^2 \Delta t^2 \cos^2 \theta + 2ux\Delta t \cos \theta + u^2 \Delta t^2 \sin^2 \theta}$$

$$r = \sqrt{x^2 + u^2 \Delta t^2 + 2ux\Delta t \cos \theta}$$

$$r = x \sqrt{1 + 2\cos \theta \frac{u\Delta t}{x} + \left(\frac{u\Delta t}{x}\right)^2}$$

Cartesian reference frame with observer at the origin

If  $u\Delta t \ll x$

$$r \approx x \sqrt{1 + 2\cos \theta \frac{u\Delta t}{x}} \approx x \left(1 + \cos \theta \frac{u\Delta t}{x}\right) = x + u\Delta t \cos \theta$$

$\therefore r - x \approx u\Delta t \cos \theta$

Hence frequency of radiation received at O is  $F = 1/T$  where:

$$\frac{1}{F} = \Delta t + \frac{u\Delta t \cos \theta}{w} = \Delta t \left( 1 + \frac{u \cos \theta}{w} \right)$$

In a *Classical* scenario, where  $u, w$  are much less than the speed of light:

$$f = 1/\Delta t$$

Received frequency

$$\therefore F = \frac{1}{1 + \frac{u}{w} \cos \theta} f$$

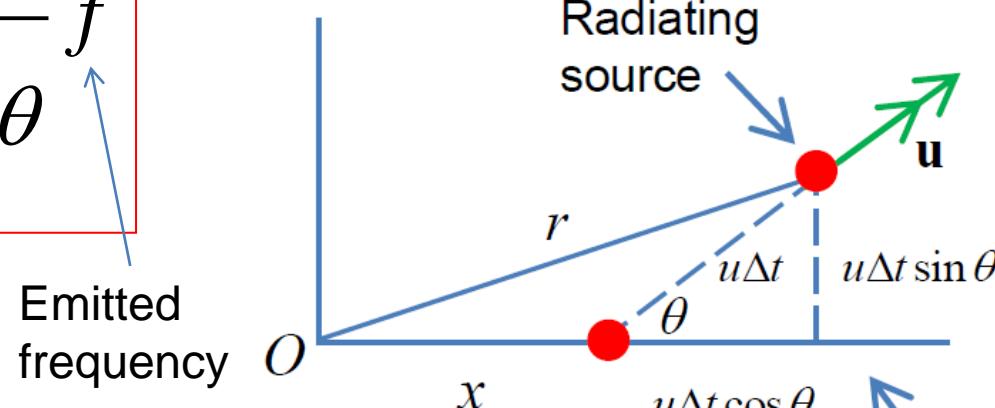
Emitted frequency

If  $w \gg u, \theta \approx 0$

$$F \approx f - \frac{u}{w} f$$

i.e. a **Doppler Shift** of:

$$\Delta f = \frac{u}{w} f$$



Cartesian reference frame  
with observer at the origin

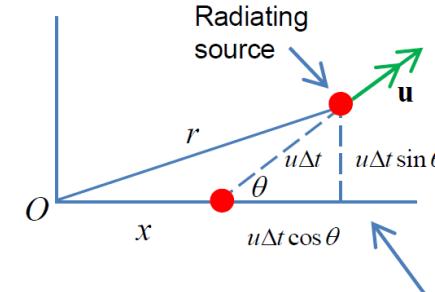
Recession velocity from observer

Observed wavelength

Emitted wavelength

$$\frac{\lambda_o}{\lambda_e} = 1 + \frac{u}{w} \cos \theta$$

wave speed



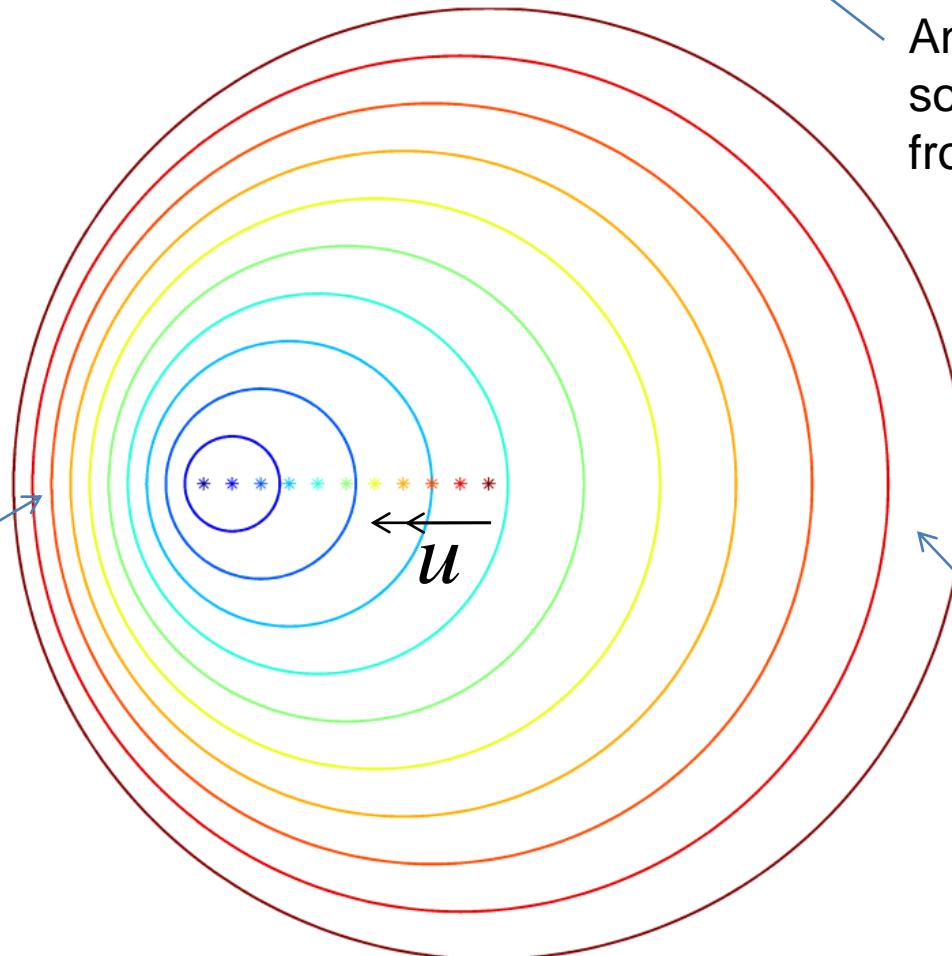
In this example:

$$\theta = 0$$

$$u = 0.6w$$

$$\lambda_o = \lambda_e (1 - 0.6)$$

Angle of source velocity from horizontal



$$\lambda_o = \lambda_e (1 + 0.6)$$

## Mach's construction

$c$  is the wave speed

$u$  is the velocity of the object making the waves

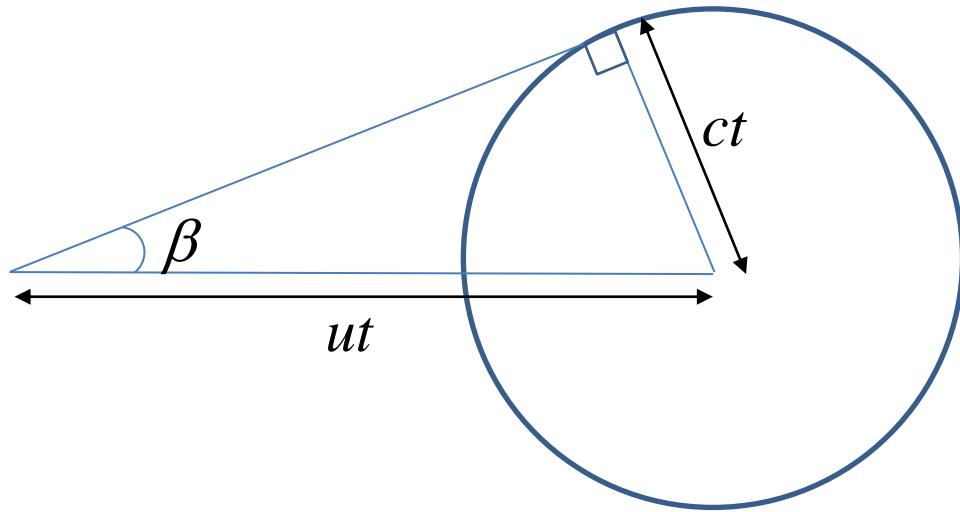
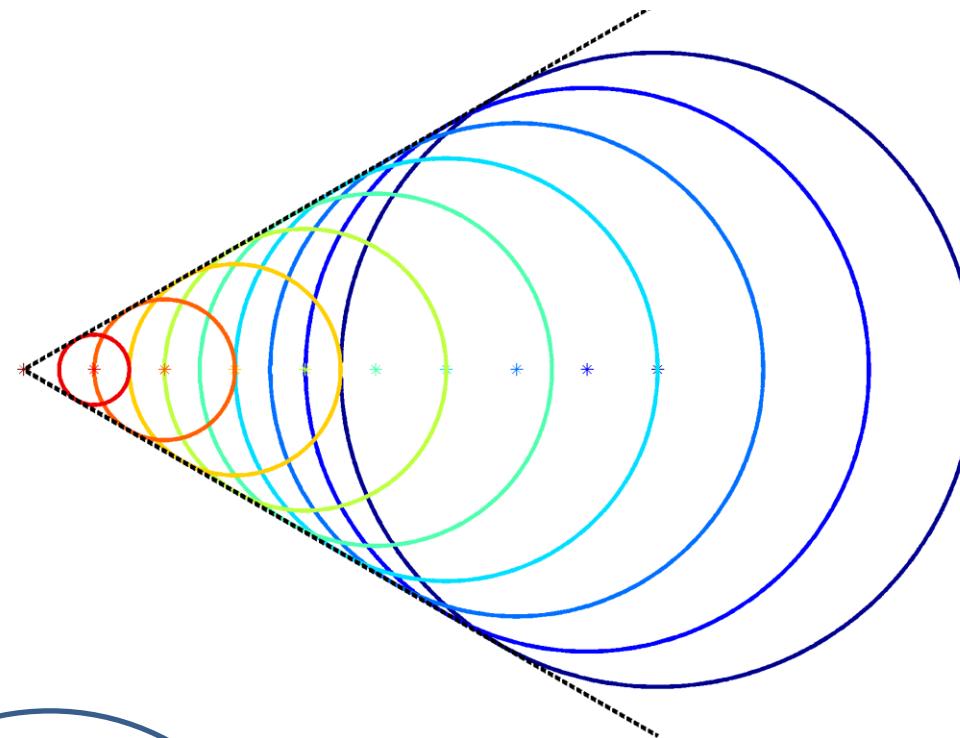
$$v/c = 2, \sin^{-1}(c/v) = 30^\circ$$



Ernst Mach 1838-1916

**Mach number**

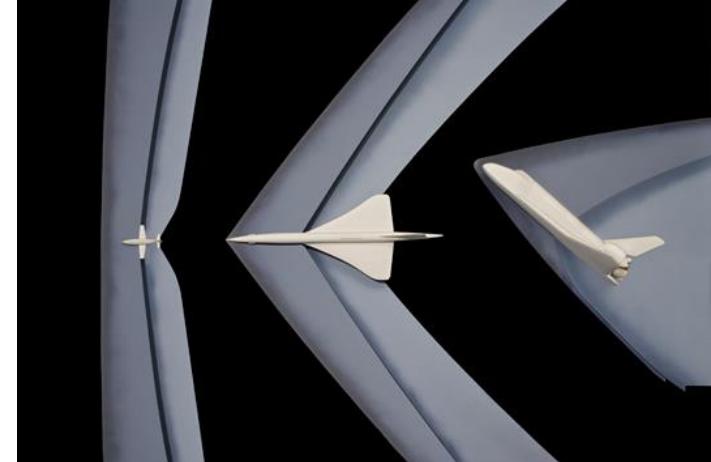
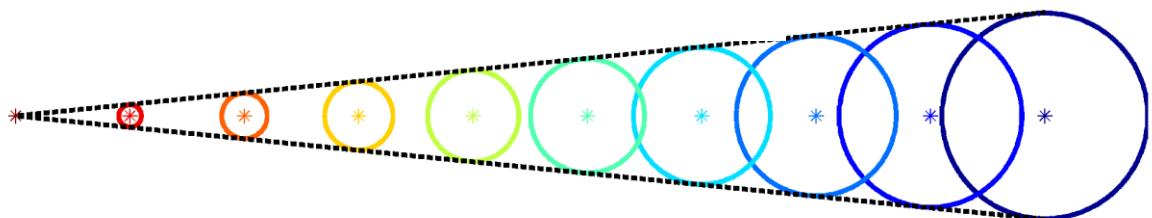
$$M = \frac{u}{c}$$



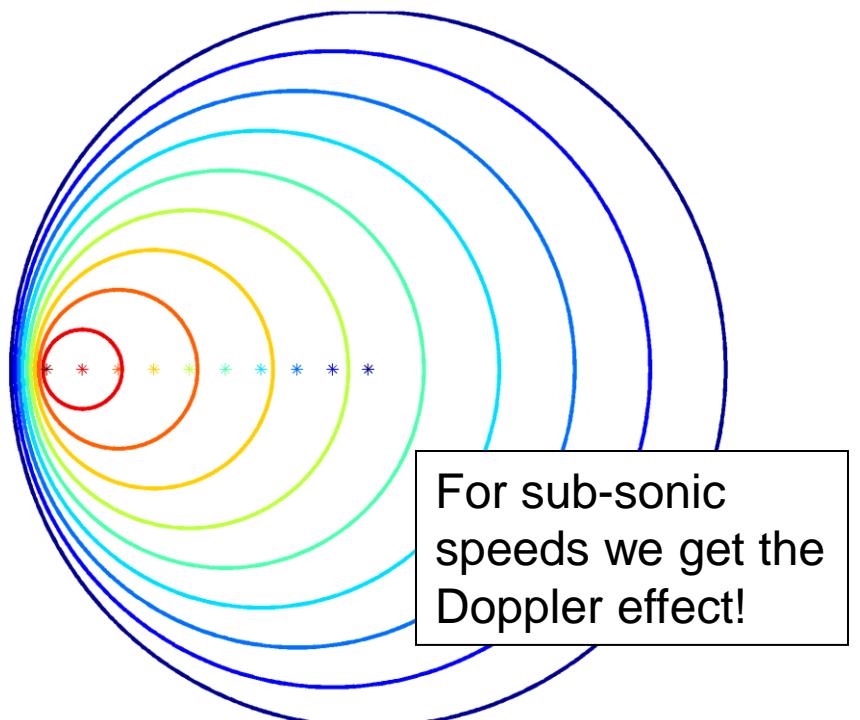
$$ut \sin \beta = ct$$

$$\therefore \beta = \sin^{-1} \left( \frac{c}{u} \right) = \sin^{-1} \frac{1}{M}$$

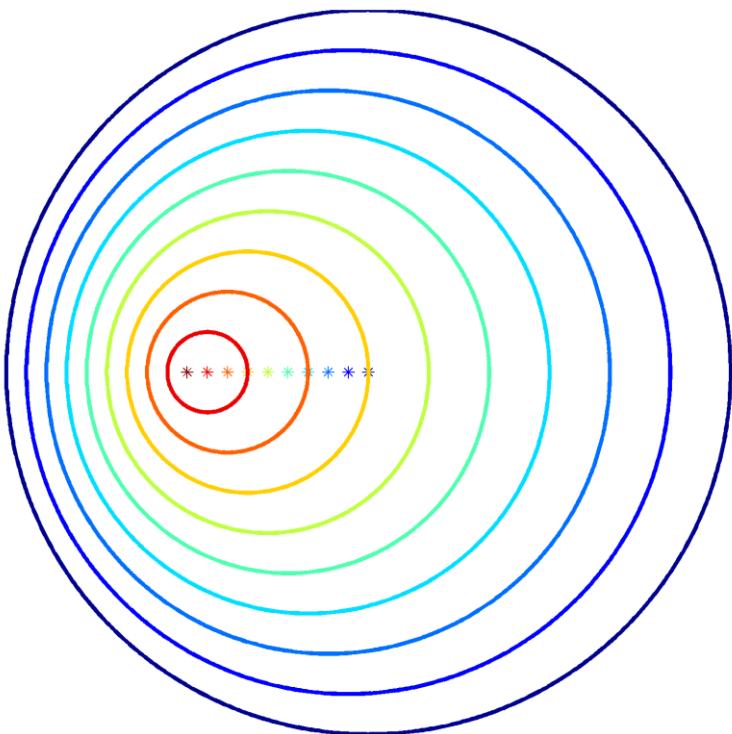
$$v/c = 10. \sin^{-1}(c/v) = 5.7392^\circ$$



$$v/c = 0.9. \sin^{-1}(c/v) = \text{NaN}^\circ$$



$$v/c = 0.5. \sin^{-1}(c/v) = \text{NaN}^\circ$$

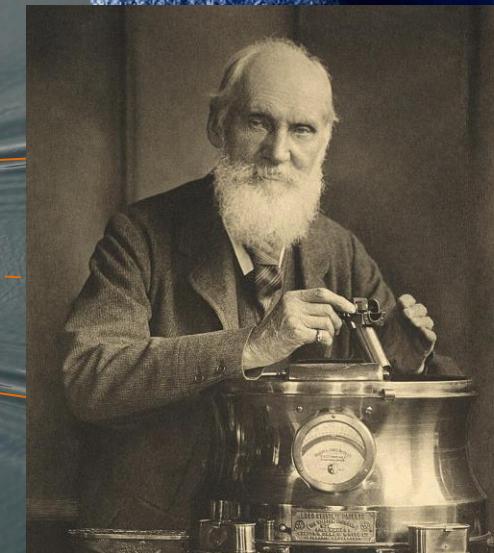
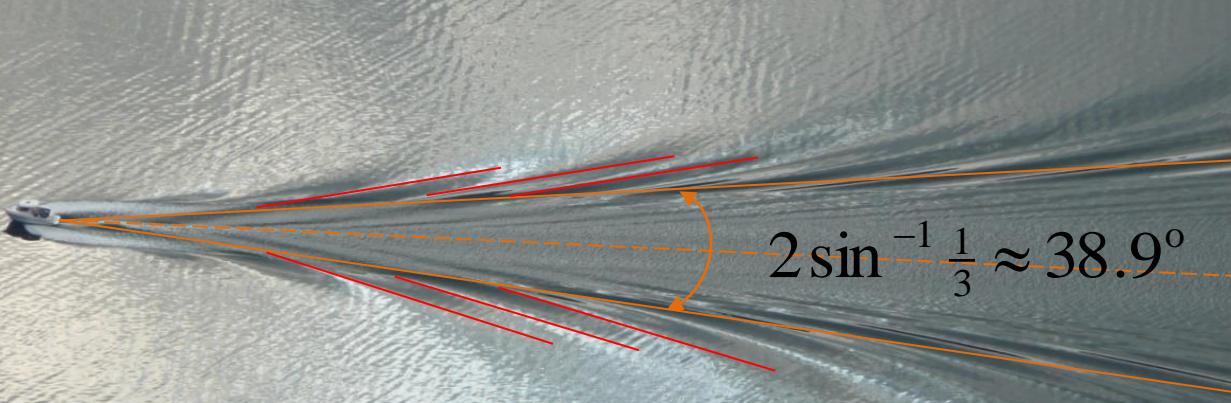
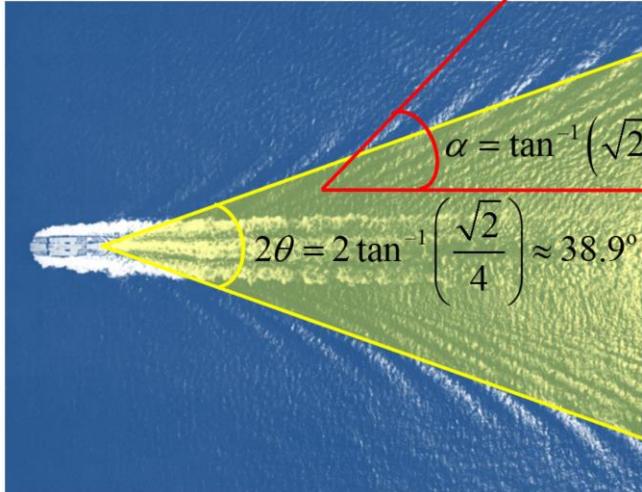
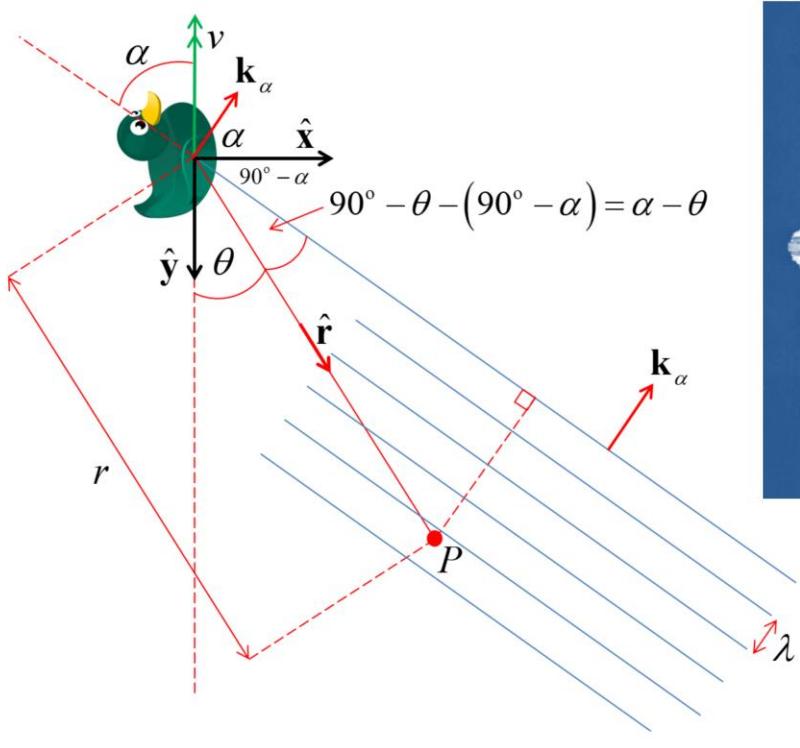




A wake is an *interference pattern* of waves formed by the motion of a body through a fluid. Intriguingly, the angular width of the wake produced by ships (and ducks!) in deep water is the same (about  $38.9^\circ$ ). A mathematical explanation for this phenomenon was first proposed by [Lord Kelvin](#) (1824-1907). The triangular envelope of the wake pattern has since been known as

<http://en.wikipedia.org/wiki/Wake>

**the Kelvin Wedge**



# The Kelvin Wedge

William Thomson, 1<sup>st</sup> Baron Kelvin  
1824-1907