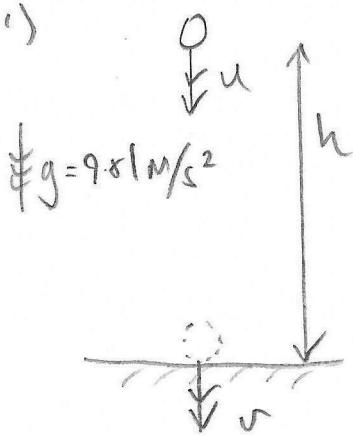


- a) A steel ball is thrown down with a speed of 3.0 m s^{-1} on to a hard surface from a height of 2.0 m . It retains 70% of its energy on each bounce. Calculate
- the speed at which it hits the ground for the first time, and
 - the maximum height it reaches after the 4th bounce.

[2]

(i)



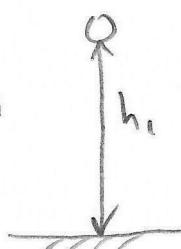
If ball has mass m kg and no energy is lost during the fall...

...By conservation of energy: $\frac{1}{2}Mv^2 = \frac{1}{2}Mu^2 + mgh$

$$\begin{aligned}\therefore v &= \sqrt{u^2 + 2gh} \\ &= \sqrt{3.0^2 + 2 \times 9.81 \times 2.0} \quad (\text{m/s}) \\ &= 6.9 \text{ m/s}\end{aligned}$$

(ii) After first bounce, ball has energy (KE) $\eta (\frac{1}{2}mu^2 + mgh)$.

By conservation of energy it will rise to height h_1 , where $mgh_1 = \eta (\frac{1}{2}mu^2 + mgh)$.



After the next bounce, $mgh_2 = \eta mgh_1$,

$$\Rightarrow h_n = \eta^n \left(\frac{u^2}{2g} + h \right)$$

$$\text{So } h_4 = 0.7^4 \left(\frac{3.0^2}{2 \times 9.81} + 2.0 \right) \quad (\text{m})$$

$$= 0.7^4 \times 2.46 \quad (\text{m}) = 0.59 \text{ m}$$

$$[h_1 = 0.7 \times 2.46 \text{ m} = 1.72 \text{ m}]$$

- b) A long-distance cyclist uses a cycle computer that is dual-powered: it has an internal battery and a solar panel. While the cyclist is riding in direct sunlight, the solar panel on the computer provides energy at a rate equal to $\frac{1}{3}$ of the power consumption. If a fully-charged cycle computer lasts 10 hours when riding at night, calculate how long a fully charged computer will run for in direct sunlight.

[2]

* Assume power consumption P is constant, and a fully charged computer constitutes energy E

$$\text{So } E = P t_N$$

where $t_N = 10 \text{ hours}$
(i.e. night time)

* During t_D hours in direct sunlight
it will expend $P t_D$ Joules and gain $\frac{1}{3} P t_D$ Joules.

$$\text{So } \underbrace{P t_D}_{\substack{\text{total} \\ \text{energy spent}}} = \underbrace{P t_N}_{\substack{\text{initial} \\ \text{charge} \\ \text{energy}}} + \underbrace{\frac{1}{3} P t_D}_{\substack{\text{gain from} \\ \text{solar panel}}}$$

$$\therefore \frac{2}{3} t_D = t_N \Rightarrow t_D = \frac{3}{2} t_N$$

$$\Rightarrow t_D = 15 \text{ hours}$$

[If solar panel energy input is ηP Watts $\leftarrow 0 < \eta < 1$

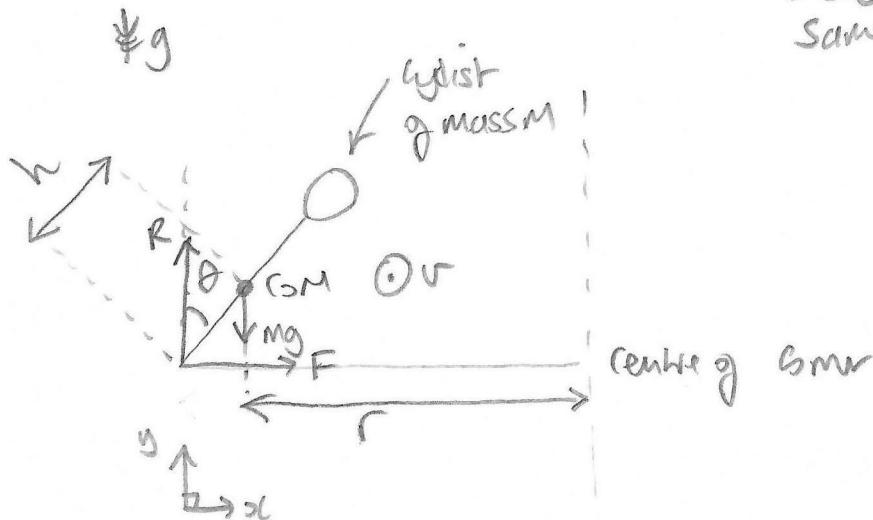
$$\Rightarrow P t_D = P t_N + \eta P t_D \text{ is the energy balance}$$

$$\therefore t_D(1-\eta) = t_N$$

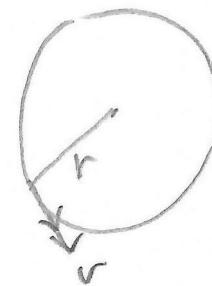
$$t_D = \frac{1}{1-\eta} t_N$$

].

- c) Riding round a corner at 10 km h^{-1} a cyclist leans over at an angle of 12° to the vertical.
At what angle would they lean over if they went round the corner at 15 km h^{-1} ?



θ
Same θ , $r = \text{constant}$ [3]

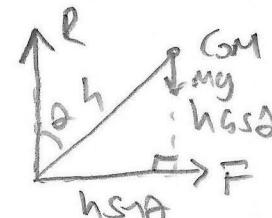


Note this is the circle (sector) made by the cyclist COM.

Circular motion $\Rightarrow N\ddot{x} \parallel x$ direction

$$\frac{mv^2}{r} = F \quad \textcircled{1}$$

If it is accelerating!



and $N\ddot{x} \parallel y$ direction

$$\ddot{\theta} = \frac{R - mg}{m} \quad \textcircled{2}$$

$$\Rightarrow R = mg$$

You can't choose ANY location since $m\ddot{r}$ in eq.

Now one assumes no rotation about the COM, so no net torque $\ddot{\theta} +$

$$\therefore \ddot{\theta} = \frac{R\ddot{h}_{S1F} - F\ddot{h}_{S2F}}{m} \Rightarrow F = R\tan\theta$$

$$\therefore \text{in } \textcircled{1} \quad \frac{mv^2}{r} = mg\tan\theta$$

$$\Rightarrow \tan\theta = \frac{v^2}{rg}$$

$$\tan 12^\circ = \frac{(10 \text{ tan/h})^2}{rg}$$

so if $rg = \text{constant}$

$$\frac{\tan 15^\circ}{\tan 12^\circ} = \left(\frac{15}{12}\right)^2$$

$$\Rightarrow \theta_{15} = \tan^{-1} \left(\tan 12^\circ + 1.5^2 \right)$$

$$= 26^\circ$$

- d) Io and Europa are both moons of Jupiter. Europa takes twice as long as Io to complete an orbit. What is the ratio of the centripetal acceleration of Io and Europa, $\frac{a_{Io}}{a_{Europa}}$?

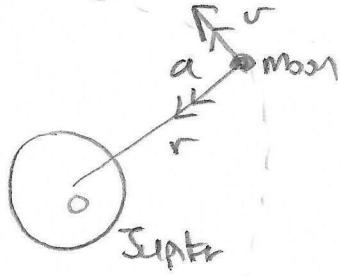
You may use the result that for this gravitational system, $\omega^2 r^3 = \text{constant}$, where ω is the angular velocity and r is the radius of the orbit.

[2] This is

a statement of
Kepler's third law

$$\text{Since } \omega = \frac{2\pi}{T} \quad T \text{ period}$$

and $\therefore [T^2 \propto r^3]$



$$a = r\omega^2$$

$$[v = r\omega]$$

So

$$\frac{a_{Io}}{a_{Europa}} = \frac{r_{Io} \omega_{Io}^2}{r_{Europa} \omega_{Europa}^2}$$

$$\omega = \frac{2\pi}{T}$$

Now

$$\omega_{Io}^2 r_{Io}^3 = \omega_{Europa}^2 r_{Europa}^3$$



So

$$\frac{r_{Io}}{r_{Europa}} = \left(\frac{\omega_{Europa}}{\omega_{Io}} \right)^{2/3} = \left(\frac{T_{Io}}{T_{Europa}} \right)^{2/3}$$

$$\therefore \frac{a_{Io}}{a_{Europa}} = \left(\frac{T_{Io}}{T_{Europa}} \right)^{2/3} \left(\frac{T_{Europa}}{T_{Io}} \right)^2$$

$$= \left(\frac{T_{Europa}}{T_{Io}} \right)^{4/3} = 2^{4/3} = \boxed{2.52}$$

(Since $T_{Europa} = T_{Io} + 2$)

e) A particle of mass m_1 and initial speed u makes an elastic collision with a stationary particle of mass m_2 . The particles move off at speeds v_1 and v_2 respectively, at equal angles θ either side of the initial incident direction of m_1 .

(i) What is the largest ratio of $\frac{m_1}{m_2}$ for which this equal angle condition can occur?

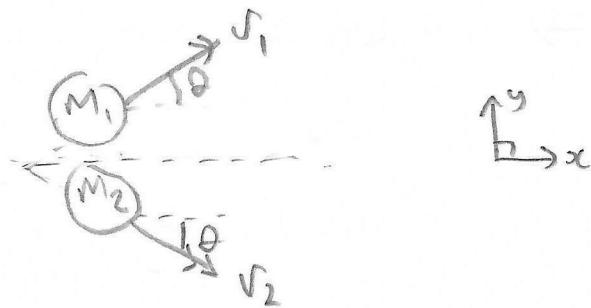
(ii) If $m_1 = m_2$, what is the largest angle of deflection, θ , of particle m_1 for this equal angle condition?

[5]

BEFORE



AFTER



eliminate v_1 and v_2

Conservation of momentum: //x

$$m_1 u = m_1 v_1 \cos \theta + m_2 v_2 \cos \theta \quad (1)$$

$$0 = m_1 v_1 \sin \theta - m_2 v_2 \sin \theta \quad (2)$$

Conservation of energy (since elastic)

$$\frac{1}{2} m_1 u^2 = \frac{1}{2} m_1 v_1^2 + \frac{1}{2} m_2 v_2^2 \quad (3)$$

(2):

$$v_2 = \frac{v_1 m_1}{m_2}$$

$$\text{In (1): } m_1 u = m_1 v_1 \cos \theta + m_2 v_2 \frac{m_1}{m_2} \cos \theta = 2 m_1 v_1 \cos \theta$$

$$\therefore v_1 = \frac{u}{2 \cos \theta}$$

$$\therefore v_2 = \frac{m_1}{m_2} \frac{u}{2 \cos \theta}$$

$$\therefore \text{In (3): } \frac{1}{2} m_1 u^2 = \frac{1}{2} m_1 \frac{u^2}{4 \cos^2 \theta} + \frac{1}{2} m_2 \frac{m_1^2}{m_2^2} \frac{u^2}{4 \cos^2 \theta}$$

$$1 = \frac{1}{4 \cos^2 \theta} + \frac{m_1}{m_2} \frac{1}{4 \cos^2 \theta}$$

$$\Rightarrow 4 \cos^2 \theta = 1 + \frac{m_1}{m_2}$$

↓ PTO

i.) Now $6s^2\theta$ has a maximum value of 1 when $\theta = 0$

So $4\cos^2\theta = 4 \Rightarrow 1 + \frac{M_1}{M_2} = 4$ for the

largest ratio $\frac{M_1}{M_2}$ i.e. $\boxed{\frac{M_1}{M_2} = 3}$ is the largest ratio.

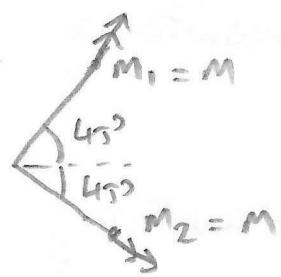
(to satisfy equal angle of split conditions)

iii) If $\frac{M_1}{M_2} = 1 \Rightarrow 4\cos^2\theta = 2$

$$\cos^2\theta = \frac{1}{2}$$

$$\cos\theta = \frac{1}{\sqrt{2}}$$

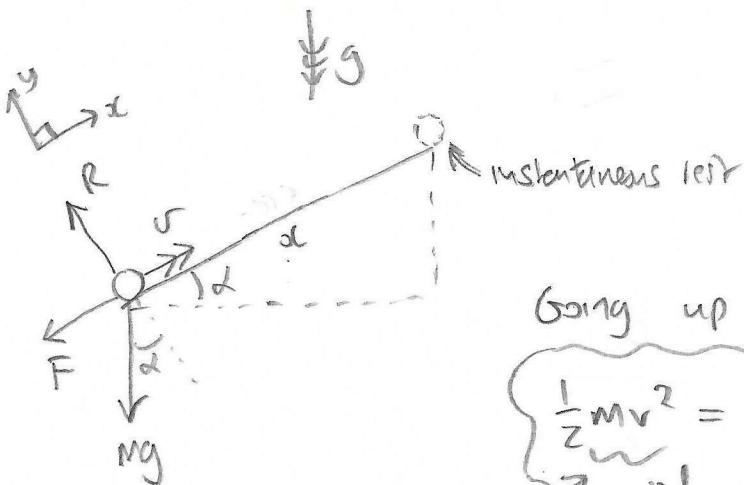
$$\Rightarrow \theta = \boxed{45^\circ}$$



- f) A body is projected with velocity v up a plane inclined at an angle α to the horizontal. When it returns through its starting point it is moving with half the speed with which it was projected.
Determine the coefficient of friction μ , in terms of the angle of the plane.

Hint: The coefficient of friction is given by $F_{\text{friction}} = \mu N$ where N is the normal contact force and F_{friction} is the frictional force on the body.

[4]



$$N \parallel g: 0 = R - mg \cos \alpha$$

$$\therefore R = mg \cos \alpha$$

$$\therefore F = \mu mg \cos \alpha$$

Conservation of energy

Going up the slope a distance x :

$$\frac{1}{2}mv^2 = mgx \sin \alpha + Fx$$

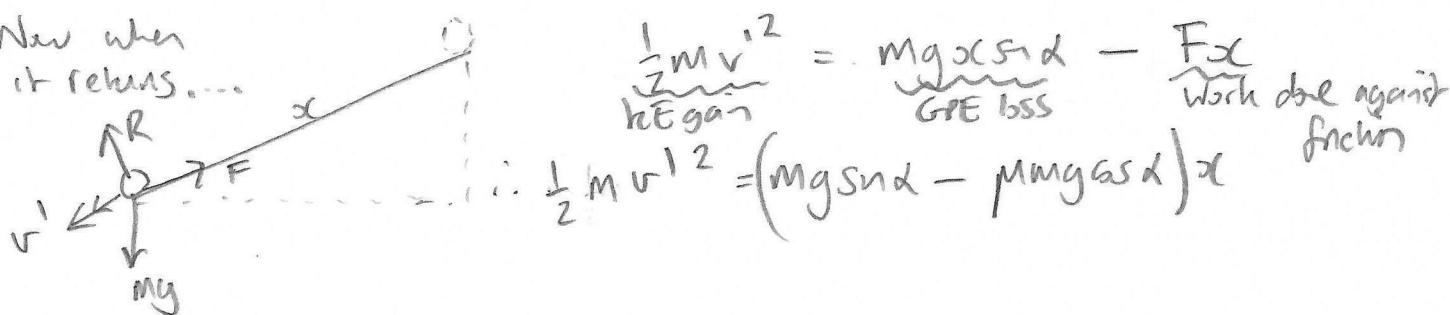
Initial KE

Work done against gravity

Work done against friction.

$$\text{So } x = \frac{\frac{1}{2}mv^2}{F + mg \sin \alpha} = \frac{\frac{1}{2}mv^2}{(mg \sin \alpha + \mu mg \cos \alpha)} = \frac{v^2}{2g / (Mg \sin \alpha + \mu Mg \cos \alpha)}$$

Now when it returns....



$$\frac{1}{2}mv'^2 = mgx \sin \alpha - Fx$$

Initial KE gain

Work done against friction

$$\therefore \frac{1}{2}mv'^2 = (mg \sin \alpha - \mu mg \cos \alpha)x$$

$$\text{So } \frac{v'^2}{2g} = \frac{\sin \alpha - \mu \cos \alpha}{\sin \alpha + \mu \cos \alpha} \frac{v^2}{2g}$$

$$\Rightarrow \left(\frac{v'}{v}\right)^2 (\sin \alpha + \mu \cos \alpha) = \sin \alpha - \mu \cos \alpha$$

↓ PO

$$M \left\{ \sin \alpha + \left(\frac{v'}{r} \right)^2 \sin \alpha \right\} = \sin \alpha - \sin \alpha \left(\frac{v'}{r} \right)^2$$

$$\therefore M = \tan \alpha \left(\frac{1 - \left(\frac{v'}{r} \right)^2}{1 + \left(\frac{v'}{r} \right)^2} \right)$$

$$\text{so } \frac{v'}{r} = \frac{1}{2}$$

$$\Rightarrow M = \tan \alpha \times \frac{\frac{3}{4}}{\frac{5}{4}}$$

$$\Rightarrow M = \frac{3}{5} \tan \alpha$$

} Can also solve this by NII // shape and
 XUVAT, but I think energy conservation
 gives you more insight}.

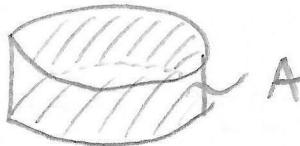
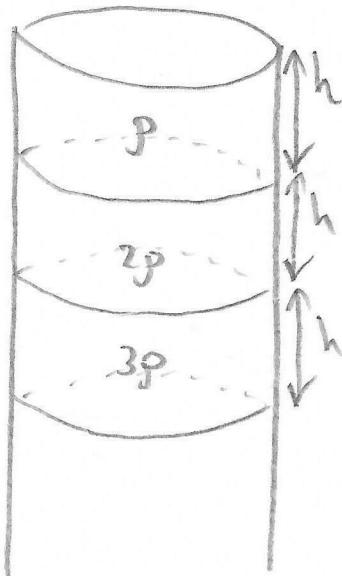
g) A cylindrical container is filled with equal volumes of n different liquids which do not mix, so that they form horizontal layers each of height h . The densities of the liquids are $\rho, 2\rho, 3\rho, \dots$, with the lowest liquid of density $n\rho$. The curved surface area of the cylinder enclosing each liquid is A .

- Give an expression for the force F_1 on the area A surrounding the top liquid in terms of ρ, g, h, A .
- What is the force F_2 on the surface A surrounding the second liquid down from the top, in terms of F_1 ?
- Give an expression for the force F_n on the area A surrounding the n^{th} liquid at the bottom of the cylinder in term of F_1 ?

Hint : $1 + 2 + 3 + \dots + n = \sum_{i=1}^n i = \frac{n}{2}(n+1)$

[4]

kg



i) Average fluid pressure in top liquid is
 $\frac{1}{2} \rho gh$

(i.e zero relative to atmospheric pressure at bp and ρgh at bottom of first layer, and linearly increasing over this range)

$$\therefore F_1 = \frac{1}{2} \rho gh A$$

i)



$$(ii) F_2 = (pgh + \frac{1}{2} (2\rho)gh) A$$

pressure due
to top layer

↑
average
pressure of
second layer

$$\therefore F_2 = 2 \rho gh A$$

↓ PTO

(iii) So force on exterior walls in layer n is

$$F_n = A \left(\rho g h + 2\rho g h + \dots + (n-1)\rho g h + \frac{1}{2} n \rho g h \right)$$

$$F_n = A \rho g h \left(1 + 2 + 3 + \dots + n-1 + \frac{n}{2} \right)$$

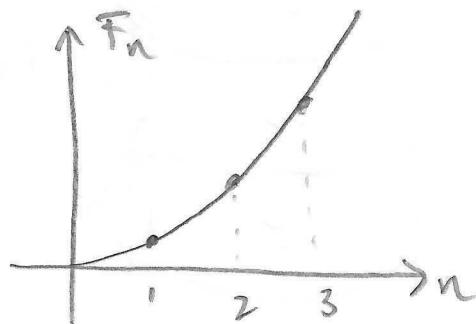
Now

$$1 + 2 + \dots + n-1 = \frac{1}{2} n(n+1)$$

$$\text{So } F_n = A \rho g h \left(\frac{1}{2} n(n+1) - \frac{n}{2} \right)$$

$$= \frac{1}{2} A \rho g h (n^2 + n - n)$$

$$\therefore F_n = \frac{1}{2} A \rho g h n^2$$



$$[S_n = 1 + 2 + 3 + \dots + n]$$

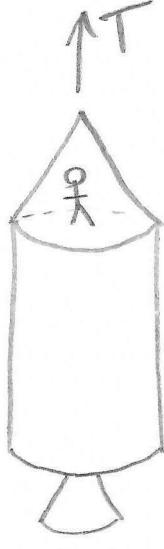
$$S_n = n + n-1 + n-2 + \dots + 1 \quad \text{reverse order}$$

$$\therefore 2S_n = n(n+1) \Rightarrow [S_n = \frac{1}{2} n(n+1)]$$

- h) A rocket of mass $m_r = 5000 \text{ kg}$ contains a further mass $m_0 = 5000 \text{ kg}$ of fuel. Once the fuel is ignited, 50 kg per second of hot gas is expelled downwards at a speed of 2000 m s^{-1} .

 - (i) Calculate the thrust, T , applied to the rocket,
 - (ii) Find an expression for the acceleration of the rocket, a , in terms of its total mass m , T and g ,
 - (iii) Find an expression for the acceleration of the rocket as a function of time, t , in terms of T , g , m_r , m_0 and t_0 , the total time for which the thrust acts.
 - (iv) Calculate the time after launch at which the weight of an astronaut on board will have appeared to double.

[4]



{ Ignore air resistance }

{ ignore that g changes w/ altitude
 \downarrow as rocket moves further from Earth }

(c) $\boxed{\overline{F} = mc}$ is rate of change of momentum

$$= 55 \times 2000 \quad (\text{N})$$

$$= \boxed{10^5} \text{ N}$$



$$(ii) \text{ NIF } \uparrow : (m_r + m_b - \mu t) a = T - mg$$

$$\therefore a = \frac{T}{m} - g$$

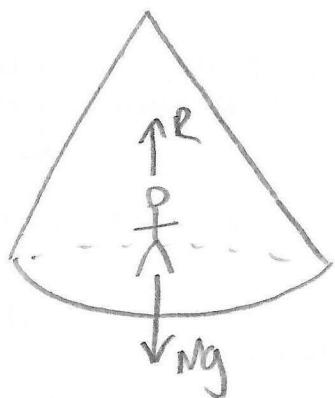
During stronger phase.

$$(iii) \quad a = \frac{T}{M_r + M_0 - \mu t} - g \quad \mu = \frac{M_0}{t_0}$$

$$a = \frac{T}{m_r + m_0 - \frac{M_0 t}{t_0}} - g$$

PTD

(iii) In astronaut's frame



$$\uparrow a$$

$$N \cancel{F} \quad ma = R - mg$$

$$So \quad R = ma + mg$$

is normal contact force on astronaut.

"weight appears to double" when $R = 2mg$

$$\therefore 2g = a + g \Rightarrow \boxed{a = g}$$

$$\text{when } a = g$$

$$g = \frac{T}{m_r + m_0 - \frac{tm_0}{t_0}} - g$$

$$\Rightarrow m_r + m_0 - \frac{tm_0}{t_0} = \frac{T}{2g}$$

$$\therefore t = \left(m_r + m_0 - \frac{T}{2g} \right) \frac{t_0}{m_0}$$

$$\therefore t = \left(\frac{m_r}{m_0} + 1 - \frac{T}{2m_0 g} \right) t_0$$

$$= \left(\frac{5000}{5000} + 1 - \frac{10^5 \times 2}{2 \times 5000 \times 9.81} \right) \left(\frac{100}{30} \right) \quad (51)$$

$$= \boxed{98s}$$

$$[\text{if } \frac{m_0}{t_0} = 50 \text{ kg/s}$$

$$\therefore t_0 = \frac{5000 \text{ kg}}{50 \text{ kg/s}}$$

$$= 100s]$$

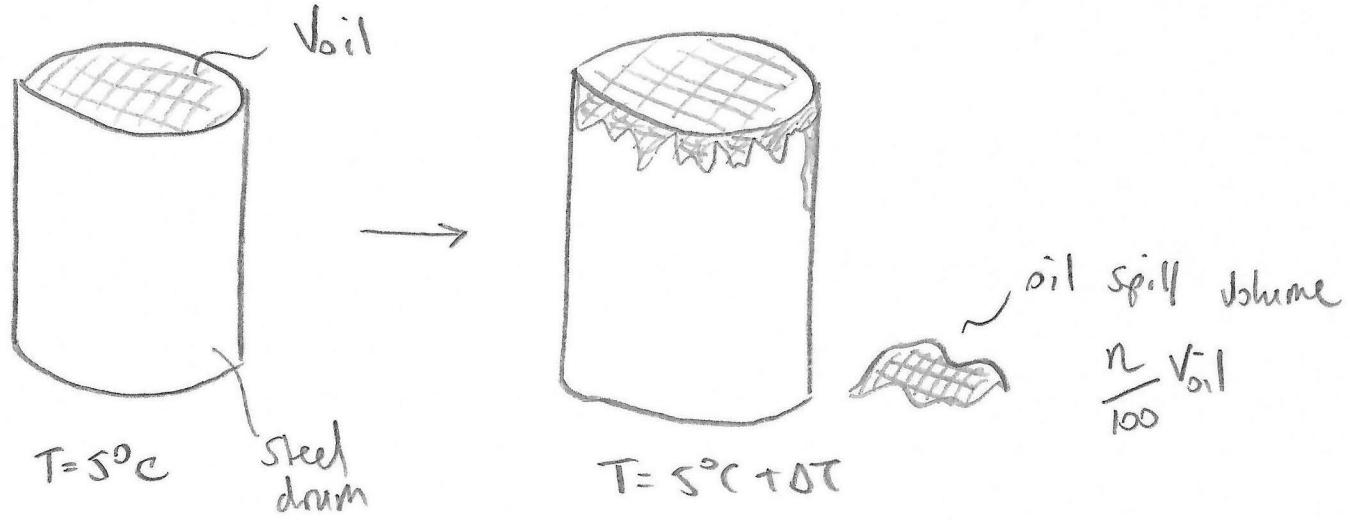
- i) An open topped steel drum is completely filled with oil on a day when the temperature is 5.0°C . On a warm day the temperature rises, and 2.4% of the oil spills out.
What is the temperature reached on that day?

$$5^{\circ}\text{C} + \Delta T$$

The volume coefficient of expansion of oil is $7.0 \times 10^{-4} \text{ }^{\circ}\text{C}^{-1}$
The linear coefficient of expansion of steel is $1.2 \times 10^{-5} \text{ }^{\circ}\text{C}^{-1}$

$$\begin{matrix} k_{\text{oil}} \\ k_{\text{steel}} \end{matrix}$$

[3]



oil leakage volume = Extra oil volume due -
to thermal expansion Extra steel drum
[internal volume] due
to thermal expansion.

$$\frac{n}{100} V_{\text{oil}} = k_{\text{oil}} V_{\text{oil}} \Delta T - 3 k_{\text{steel}} V_{\text{oil}} \Delta T$$

Initially
 V_{oil}

The drum expands (isotropically)
→ 3 dimensions.

$$\boxed{\Delta T = \frac{\frac{n}{100}}{k_{\text{oil}} - 3k_{\text{steel}}}}$$

$$= \frac{2.4/100}{7.0 \times 10^{-4} - 3 \times 1.2 \times 10^{-5}} \quad {}^{\circ}\text{C} = 36.1^{\circ}\text{C}$$

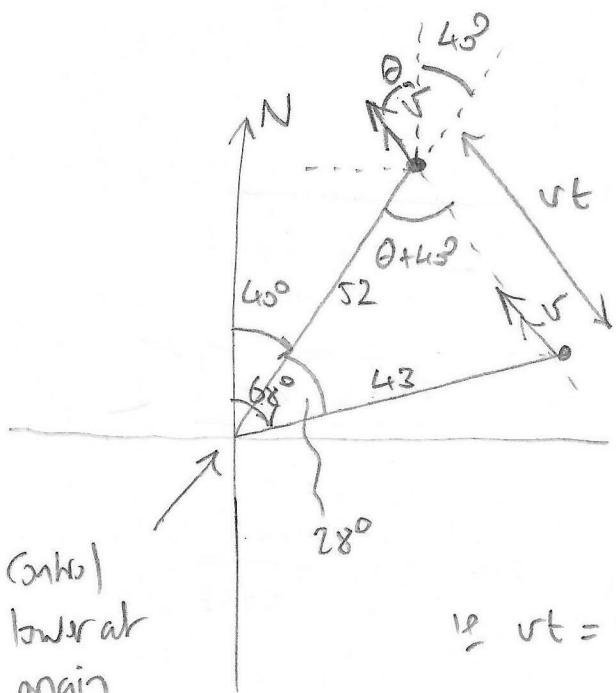
so temp was 41.1°C

* Assume the aircraft is flying at constant velocity! *

- j) Observed from an air traffic control tower, an aeroplane has a bearing of 068° and a range of 43 km. Five minutes later the bearing of the aircraft is 040° with a range of 52 km. Determine $t = 5 \times 60s$

- (i) The speed of the aircraft in $m s^{-1}$.
(ii) Its bearing and range 10 minutes after the second sighting.

[4]



(i) Cosine rule:

$$v^2 t^2 = 52^2 + 43^2 - 2(52)(43)\cos 28^\circ \\ = 604.46... (\text{km})^2$$

$$\Rightarrow v = \frac{\sqrt{604.46...}}{5 \times 60} \text{ km/s}$$

$$\therefore vt = 24.6 \text{ km}$$

$$= \boxed{82 \text{ m/s}} \rightarrow \text{Calc memory}$$

$$(ii) \text{ Sine rule: } \frac{\sin(\theta + 40^\circ)}{43} = \frac{\sin 28^\circ}{vt}$$

$$\therefore \theta = \sin^{-1} \left(\frac{43 \sin 28^\circ}{vt} \right) - 40^\circ \\ = \boxed{15.19^\circ}$$

$$\therefore \underline{v} = \begin{pmatrix} -\sin \theta \\ \cos \theta \end{pmatrix} \times 82 \text{ m/s} = \begin{pmatrix} -21.5 \\ 79 \end{pmatrix} \text{ m/s}$$

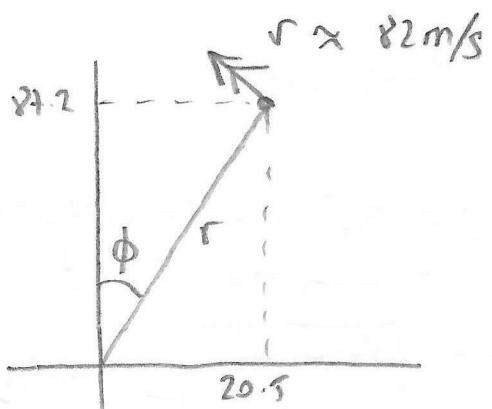
$600s = t$ after second sighting, position vector is

$$\underline{r} = \begin{pmatrix} 52 \sin 40^\circ \\ 52 \cos 40^\circ \end{pmatrix} + \begin{pmatrix} -21.5 \\ 79 \end{pmatrix} \frac{t}{1000} \text{ km.}$$

$$\therefore \underline{l} = \begin{pmatrix} 52 \sin 45^\circ \\ 52 \cos 45^\circ \end{pmatrix} + \begin{pmatrix} -21.5 \\ 79 \end{pmatrix} \times \frac{600}{1000} \text{ km}$$

$$= \begin{pmatrix} 20.5 \\ 87.2 \end{pmatrix} \text{ km}$$

MS says
89.7 km



$$\therefore \text{range } r = \sqrt{20.5^2 + 87.2^2}$$

$$= \boxed{89.6 \text{ km}}$$

$$\text{bearing } \phi = \tan^{-1} \left(\frac{20.5}{87.2} \right)$$

$$= \boxed{13.2^\circ}$$

{ Ideally we would not bend mid-way
 ↳ this is a good problem to solve algebraically and evaluate via a spreadsheet or programming script}.

[Alternative method is to continue drawing triangles and use sine and cosine rules, but I think it is simpler to use $\underline{l} = \underline{l}_0 + \underline{l}_t$ if vector geometry. This is how you would plot it on a computer anyway!]

This is actually useful!

- k) A gas is found to obey the equation relating p, V, n, R, T

$$p(V - b) = nRT \exp\left(\frac{-a}{nRTV}\right)$$

where p is the gas pressure

V is its volume

R is the molar gas constant

n is the number of moles

a and b are constants.

- (i) Determine the SI base units (m, kg, s) in which a and b are expressed.

- (ii) If $b \ll V$ and $a \ll nRTV$, show that this expression approximates to the ideal gas equation relating p, V, n, R, T at a particular temperature T_c . Determine T_c in terms of a, b, n and R .

Hint: For $x \ll 1$ $\exp(x) = e^x \approx 1 + x$

This is all you need.

[4]

(i) $[b]$ must be m^3 since a volume

$\left[\frac{-a}{nRTV}\right]$ must be dimensionless since an exponent

$$\therefore e \approx 2.7183\dots$$

$$\begin{aligned} \therefore [a] &= [nRTV] = \text{mol} \times \text{J} \cdot \text{mol}^{-1} \cdot \text{K}^{-1} \cdot \text{m}^3 \\ &= \text{kg} \cdot \text{m}^2 \cdot \text{s}^{-2} \cdot \text{m}^3 \\ &= \boxed{\text{kg} \cdot \text{m}^5 \cdot \text{s}^{-2}} \end{aligned}$$

$$\begin{aligned} [S] &= \underbrace{\text{kg} \cdot \text{m}^2}_{\text{SI unit of force}} + \underbrace{\text{m}}_{\text{SI unit of distance}} \end{aligned}$$

$$e^{-\frac{a}{nRTV}} \approx 1 - \frac{a}{nRTV}$$

$$\frac{a}{nRTV} \ll 1$$

If $b \ll V \Rightarrow V-b \approx V$

$$\therefore pV \propto nRT \left(1 - \frac{a}{nRTV}\right)$$

Don't! \downarrow PRO. Actually don't approx $V-b \approx V$ yet!

$$\rightarrow pV - pb \approx nRT - \frac{a}{V} \quad \leftarrow \text{only } \frac{a}{nRTV} \ll 1$$

$$pV \approx nRT + pb - \frac{a}{V}$$

so $pV \approx nRT$ if $pb - \frac{a}{V} = 0$
 (ideal gas eqn)

$$\Rightarrow pV = \frac{a}{b} . \quad \text{let this occur at } T=T_C$$

$$\text{so } nRT_C = \frac{a}{b}$$

$$T_C = \frac{a}{b n R}$$



↳ no forces at surface ↳ doesn't move!

- 1) A thin rod is balanced in a smooth hemispherical bowl fixed to a table, touching both the interior and the rim as shown in Fig. 1. The rim of the bowl remains horizontal. Expressed in the simplest form, determine the radius of the bowl r in terms of the length of the rod l , and the angle θ to the horizontal.

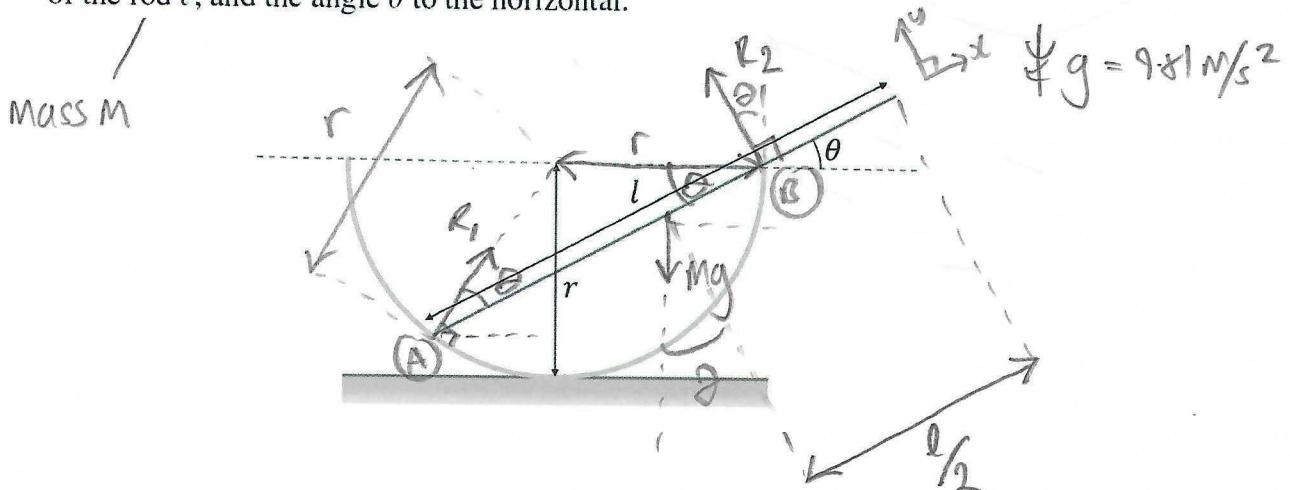


Figure 1: A rod in a smooth bowl.

[6]

NIL // rod:

$$R_1 \cos \theta - mg \sin \theta = 0 \Rightarrow R_1 = mg \tan \theta$$

↳ in E2

Mass \times acc = vector sum of force = 0

NIF \perp rod:

$$R_1 \sin \theta + R_2 - mg \cos \theta = 0$$

$$\Rightarrow R_2 = mg \cos \theta - R_1 \sin \theta$$

$$R_2 = mg \cos \theta - mg \frac{\sin^2 \theta}{\cos \theta}$$

2+ moments (note in E2) about B

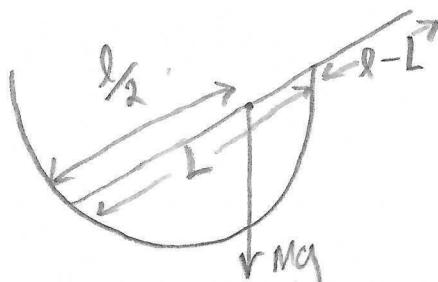
More complicated expression, so let's not use it in moments calc.

$$-R_1 \sin \theta L + mg \cos \theta (L - l/2) = 0$$

$$\therefore L (mg \cos \theta - mg \tan \theta \sin \theta) = \frac{l}{2} mg \cos \theta$$

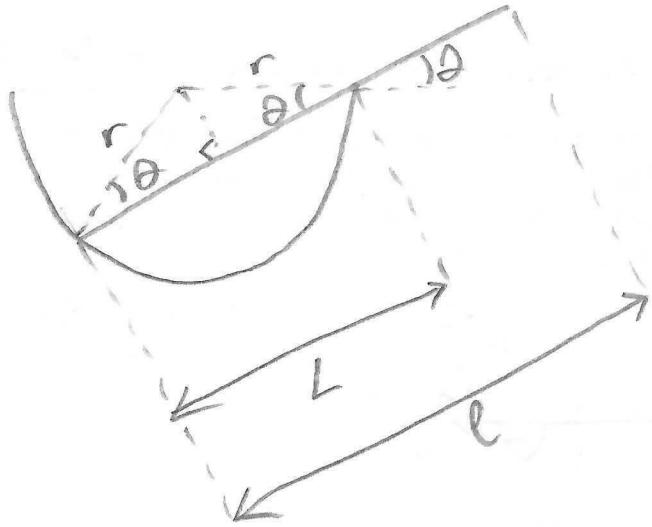
$$\Rightarrow L = \frac{l}{2} \frac{\cos \theta}{\cos \theta - \tan \theta \sin \theta}$$

↓ PTO



$$2r \cos\theta = L$$

from geometry



$$\therefore 2r \cos\theta = \frac{l}{2} \frac{\sin\theta}{\sin\theta - \tan\theta \sin\theta}$$

$$r = \frac{l}{4} \frac{1}{\sin\theta - \frac{\sin^2\theta}{\sin\theta}}$$

$$r = \frac{l}{4} \frac{\sin\theta}{\cos^2\theta - \sin^2\theta}$$

$$r = \frac{l}{4} \frac{\sin\theta}{\cos 2\theta}$$

$$[\sin(\theta + \phi) = \sin\theta \cos\phi + \cos\theta \sin\phi]$$



$$\sin(A \pm B) = \sin A \cos B \mp \cos A \sin(\pm B)$$

$$= \sin A \cos B \mp \cos A \sin B]$$

- m) In the circuit shown in **Fig. 2**, the three cells each supply an emf of 5.0 V and have an internal resistance of 5.0Ω . The external resistors also each have a resistance of 5.0Ω .

What arrangement of switches gives

- the maximum current,
- the minimum non-zero current.
- Determine the current in each case.

[3]

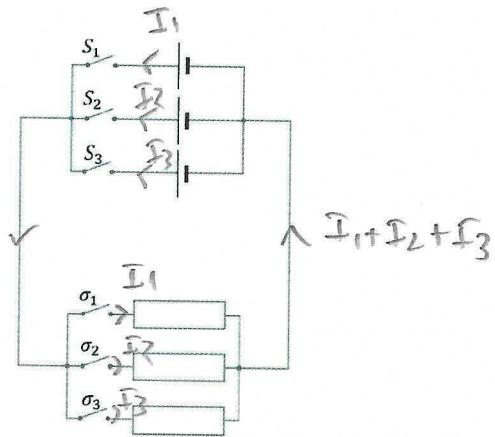


Figure 2: Three cells and switches in parallel.

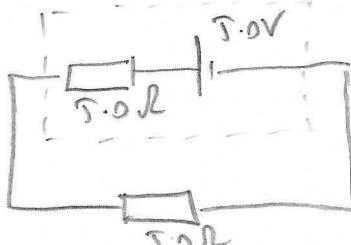
(i)

- * Closing all switches S_1, S_2, S_3 doesn't change the EMF
(KTF: can't have a closed loop including lower resistors which involves more than one 5.0V EMF)
- * If all switches $\sigma_1, \sigma_2, \sigma_3$ are closed AND S_1, S_2, S_3

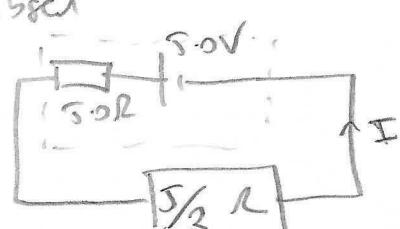
$$I_{1,2,3} = \frac{5.0V}{10\Omega} = 0.5A$$

∴ total current is $3 \times 0.5A = 1.5A$ MAX

- * One $S_{1,2,3}$ is closed and one $\sigma_{1,2,3}$ other
- Total current is $I_{1,2,3} = 0.5A$ MIN



- Note if one $S_{1,2,3}$ is closed and all $\sigma_{1,2,3}$ are closed



$$\therefore I = \frac{5.0}{5 + \frac{5}{3}} A = 0.75A$$

- n) The circuit of Fig. 3 consists of four resistors and a switch S. When the S is open, the current flowing through the $5\text{ k}\Omega$ resistor is I_0 . When S is closed, the current flowing through the same resistor is I_c . What is the ratio $\frac{I_c}{I_0}$, giving your answer as the ratio of two integers.

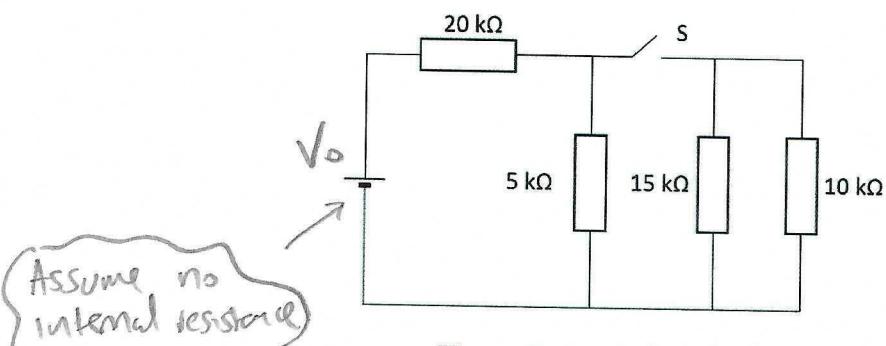
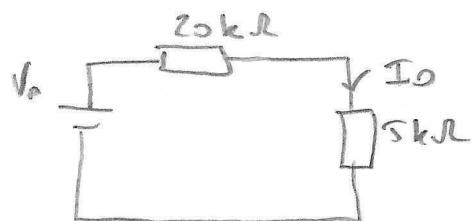
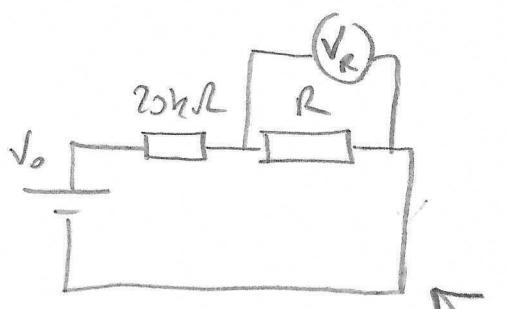


Figure 3: A switched circuit.

[4]



$$I_0 = \frac{V_o}{25\text{ k}\Omega}$$



S closed: Equivalent circuit is

$$R = \frac{1}{\frac{1}{5} + \frac{1}{15} + \frac{1}{10}} \text{ k}\Omega$$

$$= 2.72 \text{ k}\Omega \quad (\boxed{\frac{30}{11} \text{ k}\Omega} \text{ as an exact fraction})$$

$$\text{So } \frac{V_R}{V_o} = \frac{R}{R+20\text{k}\Omega} = \frac{\frac{30}{11} \text{ k}\Omega}{\frac{30}{11} \text{ k}\Omega + 20} = \boxed{\frac{3}{25}} \quad \text{Potential divider}$$

$$\therefore \text{Since } V_R = I_c \times 5\text{k}\Omega \Rightarrow I_c = \frac{\frac{3}{25} V_o}{5\text{k}\Omega}$$

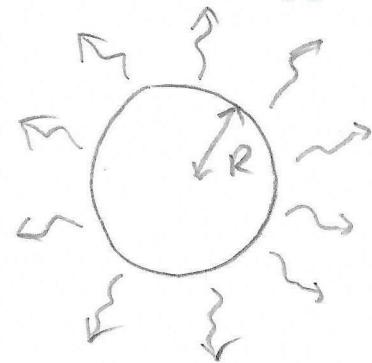
$$\therefore \frac{I_c}{I_0} = \frac{3}{125} \times 25 = \boxed{\frac{3}{5}}$$

- o) The Stefan-Boltzmann law says that the emitted power of a spherical "black body", P , is related to the radius, R , and absolute surface temperature, T , as $P \propto R^2 T^4$. Wien's "displacement law" says that the wavelength λ_{max} corresponding to the peak value of the emitted power of this spectrum is inversely proportional to the absolute surface temperature. The Sun currently has its peak wavelength as 500 nm. What will be the new peak wavelength when it becomes a red giant, given its radius will be 200 times larger and its power output 4000 times larger?

[3]

* $P = k R^2 T^4$ Stefan-Boltzmann

* $\lambda_{\text{max}} = \frac{b}{T}$ Wien



So

$$\lambda_{\text{max}} = \frac{b}{(P/kR^2)^{1/4}}$$

$$\lambda_{\text{max}, \odot} = b k^{1/4} \frac{R_{\odot}^{1/2}}{P_{\odot}^{1/4}}$$

$$\lambda_{\text{max}, RS} = b k^{1/4} \frac{R_{RS}^{1/2}}{P_{RS}^{1/4}}$$

So

$$\frac{\lambda_{\text{max}, RS}}{\lambda_{\text{max}, \odot}} = \left(\frac{R_{RS}}{R_{\odot}} \right)^{1/2} \left(\frac{P_{\odot}}{P_{RS}} \right)^{1/4}$$

$$= \frac{200^{1/2}}{4000^{1/4}} = 1.778\dots$$

$$\begin{aligned} \lambda_{\text{max}, RS} &= 1.778\dots \times 500 \text{ nm} \\ &= 889 \text{ nm} \end{aligned}$$



A nice "paradigm example".

E

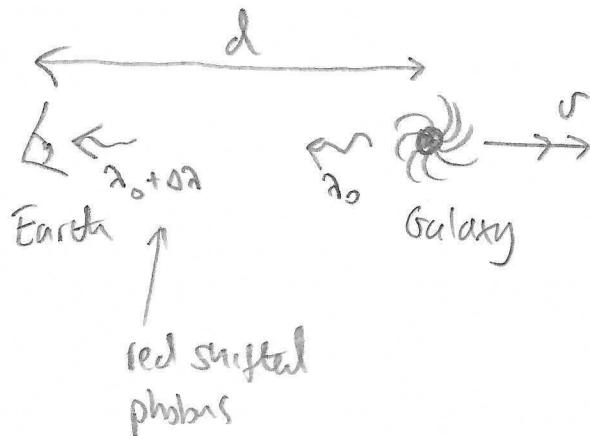
- p) One spectral line in hydrogen is caused by photons with an energy of 2.55 eV. The same line is redshifted in the spectrum of a distant galaxy by 5.4 nm. Calculate

- (i) the wavelength of the photon,
- (ii) the speed of recession of the galaxy,
- (iii) the distance to the galaxy.

How far away is the galaxy? Give your answer in megaparsecs (Mpc).

The Hubble constant, $H_0 = 70 \text{ km s}^{-1} \text{ Mpc}^{-1}$.

[3]



$$(i) E = \frac{hc}{\lambda_0}$$

$$\therefore \lambda_0 = \frac{hc}{E}$$

$$\therefore \lambda_0 = \frac{6.63 \times 10^{-34} \times 2.998 \times 10^8}{2.55 + 1.602 \times 10^{-12}} \text{ m}$$

(ii) Doppler effect (z₁ axis)

$$= 4.87 \times 10^{-7} \text{ m}$$

$$= 487 \text{ nm}$$

$$\boxed{\frac{\Delta \lambda}{\lambda_0} = \frac{v}{c}}$$

$$\therefore v = \frac{5.4}{4.87} \times 3.00 \times 10^8 \text{ m/s}$$

$$= 3.32 \times 10^6 \text{ m/s}$$

(iii) Hubble's law

$$\boxed{v = H_0 d}$$

$$\therefore d = \frac{v}{H_0}$$

$$v = \frac{\Delta \lambda}{\lambda_0} c$$

$$\therefore d = \frac{5.4}{4.87} \times 3.00 \times 10^8 \text{ (km)} / \text{s} \quad \frac{70 \text{ km/s}}{70 \text{ km/s/Mpc}}$$

$$= 47.5 \text{ Mpc}$$

For less compounded rounding errors use

$$\therefore d = \frac{\Delta \lambda}{\lambda_0} \frac{c}{H_0}$$

$$\Rightarrow d =$$

$$= 47.5 \text{ Mpc}$$

- q) A car drives along a road that has small depressions regularly spaced about 8.0 m apart. When four 80 kg passengers enter the 800 kg car, it sinks down by 1.8 cm. At approximately what speed might travelling in the vehicle become very uncomfortable?

[3]

* This is a resonance problem. When the time between bumps \approx period of oscillation of the car's suspension \Rightarrow maximum amplitude of vibrations

* Model the car suspension as a mass-spring system.
 Assume 'sink' of 1.8cm is related to the unoccupied car. \therefore spring constant $k = \frac{4 \times 80 + 981}{1.8 \times 10^{-2}}$
 $= 1.74 \times 10^5 \text{ N/m}$

N/m

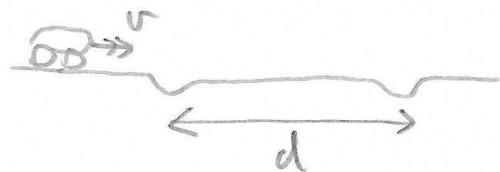
* Period of S.H.M. in a mass-spring system is

$$T = 2\pi \sqrt{\frac{m}{k}}$$

This is now the total mass
& car + passengers

* Time between bumps

$$t = \frac{d}{v}$$



$$\text{So if } t=T \Rightarrow \frac{d}{v} = 2\pi \sqrt{\frac{m}{k}}$$

$$\therefore v = \frac{1}{2\pi} d \sqrt{\frac{k}{m}}$$

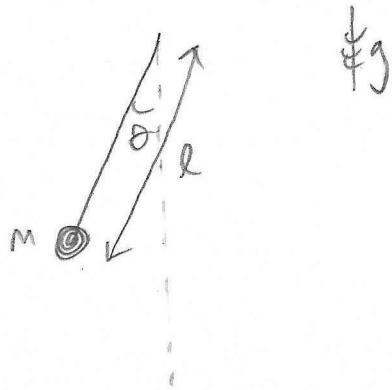
$$= \frac{1}{2\pi} \times 8 \sqrt{\frac{1.74 \times 10^5}{4 \times 80 + 800}} \text{ m/s}$$

$$= 15.9 \text{ m/s}$$

↑ don't forget the
car mass

- r) A pendulum clock is controlled by the swing of a simple pendulum (a mass on the end of a light rod) and is intended to have a period of 1.00 seconds. However, the clock runs slow by ten minutes each day. What percentage change should be made in the length of the pendulum?

[3]



* Period of pendulum is

$$T = 2\pi \sqrt{\frac{l}{g}}$$

(for small oscillations
s.t. $\sin \theta \approx \theta$
in radians)

T is too long by ΔT . $T_0 = 1.00\text{s}$ is desired
let $l = l_0 + \delta l$ where δl is the change which
reduces T to T_0

$$\text{So } T_0 + \Delta T = 2\pi \sqrt{\frac{l_0}{g}}$$

$$T_0 = 2\pi \sqrt{\frac{l_0 + \delta l}{g}}$$

MS anticipates a binomial expansion of $\left(\sqrt{1 + \frac{\delta l}{l_0}}\right)^{-1}$

to $1 - \frac{1}{2} \frac{\delta l}{l_0}$, but we don't need to!

$$\Rightarrow \frac{\delta l}{l_0} \approx -\frac{2\Delta T}{T_0}$$

$$= -\frac{1}{72}$$

$$\approx -1.4\%$$

$$\text{Now } \frac{\Delta T}{T_0} = \frac{10 \times 60}{24 \times 3600} = \frac{1}{144}.$$

$$1 + \frac{\Delta T}{T_0} = \frac{\frac{2\pi}{\sqrt{g}} \sqrt{l_0}}{\frac{2\pi}{\sqrt{g}} \sqrt{l_0 + \delta l}}$$

$$\frac{\delta l}{l_0} = \left(1 + \frac{\Delta T}{T_0}\right)^{-2} - 1$$

Exact.

$$= \left(1 + \frac{1}{144}\right)^{-2} - 1$$

$$= -0.014$$

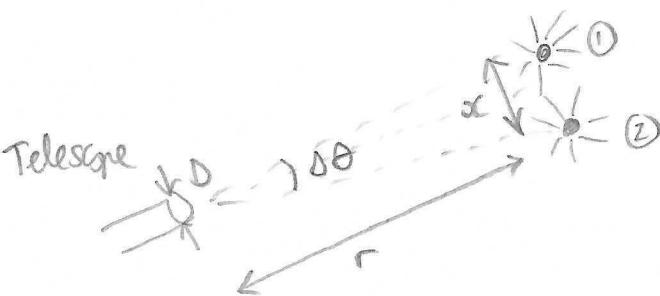
to 2sf
Same as approx.

$$1 + \frac{\Delta T}{T_0} = \frac{1}{\sqrt{1 + \frac{\delta l}{l_0}}}$$

so shorten l_0 by 1.4%

- s) A binary star system is 2140 light years away and consists of two stars like the Sun. The average separation between the stars is 0.00593 light years. Determine the diameter of the telescope needed to resolve them if using a visible wavelength of 550 nm.

[2]



Since $r \gg x$

$$r\theta \approx x$$

$$\theta \approx \frac{x}{r}$$

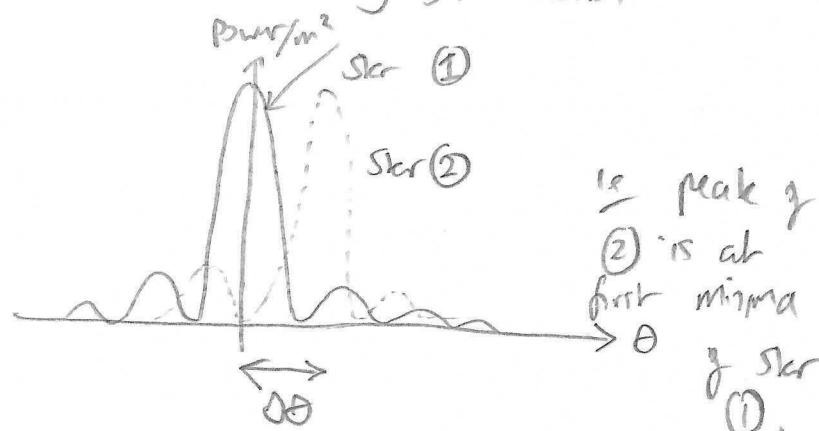
$$\approx \frac{0.00593}{2140} \text{ radians}$$

$$\approx [2.77 \times 10^{-6} \text{ radians}]$$

To 'just' resolve the stars, assume the Rayleigh criterion
is a diffraction limit

$$\theta \approx 1.22 \frac{\lambda}{D} \quad \text{for}$$

a spherical aperture of diameter D.



$$[\text{If thin slits, } \theta \approx \frac{\lambda}{D}]$$

$$\text{So to resolve the stars, } D > \frac{\lambda}{\theta} \times 1.22$$

$$D > \frac{550 \times 10^{-9} \text{ m}}{2.77 \times 10^{-6}} \times 1.22$$

$$D > 0.24 \text{ m}$$

which sounds like something an amateur astronomer could achieve with a 15cm telescope.

t) A glass prism of refracting angle 75.0° is shown in Fig. 4 has a refractive index of $n = 1.40$.

- (i) For what range of incident angles will light from air that is incident on face AB emerge from face AC?
- (ii) Show your result on a diagram.

[3]

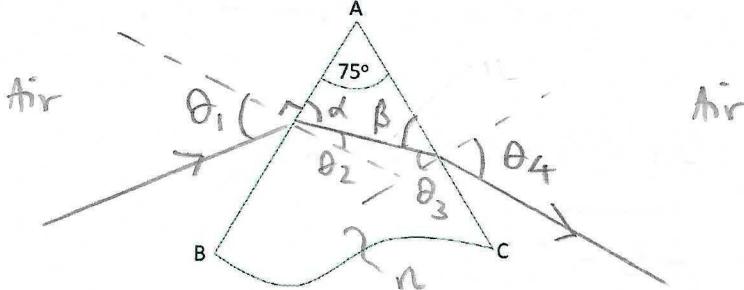


Figure 4: Glass prism with an apex angle of 75° .

Normals:

$$90^\circ = \alpha + \theta_2 \quad ① \qquad 75^\circ + \alpha + \beta = 180^\circ \quad ③$$
$$90^\circ = \beta + \theta_3 \quad ②$$



Snell:

$$n \sin \theta_1 = \sin \theta_2 \quad ④$$
$$\sin \theta_4 = n \sin \theta_3 \quad ⑤$$

So $180^\circ = \alpha + \beta + \theta_2 + \theta_3 \quad ① + ②$

~~$75^\circ + \alpha + \beta = \alpha + \theta_2 + \theta_3$~~ using ③

$$\therefore 75^\circ = \theta_2 + \theta_3 \quad \therefore \theta_3 = 75^\circ - \theta_2$$

$\therefore n \sin \theta_4 = n \sin (75^\circ - \theta_2)$

$$\boxed{\theta_4 = \sin^{-1} \left(n \sin \left(75^\circ - \sin^{-1} \left(\frac{1}{n} \sin \theta_1 \right) \right) \right)}$$

$\frac{1}{n} \sin \theta_1$ will < 1 for $0 \leq \theta_1 \leq 90^\circ$ so expect rays to emerge from AC for all angles of incidence. ↓ PCO

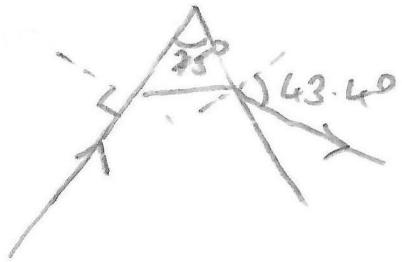
* largest θ_4 is clearly when θ_1 is largest, so

set $\theta_1 = 90^\circ \therefore \sin\theta_1 = 1$

$\therefore \sin^{-1}\left(\frac{1}{n} \sin\theta_1\right) = \sin^{-1}\left(\frac{1}{1.4}\right) = \boxed{45.6^\circ}$

$\therefore \theta_4 = \sin^{-1}\left(1.4 \sin\left(75^\circ - 45.6^\circ\right)\right)$

$$\begin{aligned} &= \sin^{-1}(1.4 \sin(29.4^\circ)) \\ &= \boxed{43.4^\circ} \end{aligned}$$



So expect mat θ_4 to be 43.4°

[when $\theta_1 = 0$, $\theta_4 = \sin^{-1}(1.4 \sin 75^\circ)$

$$= \sin^{-1}(1.35) \quad \text{is no}$$

real solution

so range of θ_1 that give rise to real θ_4 is

when $n \sin(75^\circ - \sin^{-1}\left(\frac{1}{n} \sin\theta_1\right)) \leq 1$

$$75^\circ - \sin^{-1}\left(\frac{\sin\theta_1}{n}\right) \leq \sin^{-1}\left(\frac{1}{n}\right)$$

$$75^\circ - \sin^{-1}\left(\frac{1}{n}\right) \leq \sin^{-1}\left(\frac{\sin\theta_1}{n}\right)$$

$$\theta_1 \geq \sin^{-1}\left(n \sin\left(75^\circ - \sin^{-1}\left(\frac{1}{n}\right)\right)\right)$$

$$\geq \sin^{-1}\left(1.4 \sin\left(75^\circ - \sin^{-1}\left(\frac{1}{1.4}\right)\right)\right)$$

$$\geq \boxed{43.4^\circ} \quad \text{A nice result!}$$

]

*Trig fest! *

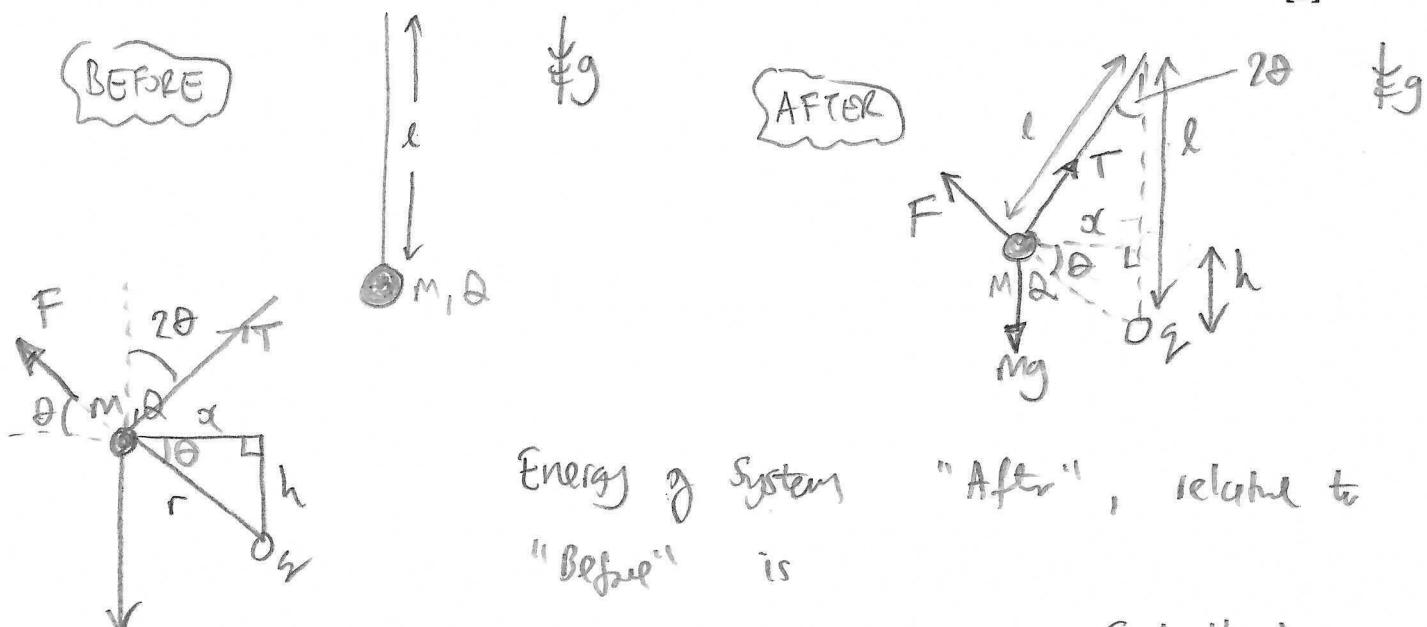
A somewhat epic question...

MUCH More structure needed - may algebraic pitfalls
... But need a classic!!

- u) A particle **A**, of mass m carrying a charge of Q is suspended by an insulating thread of length ℓ . Another particle **B**, of negligible mass but of positive charge $+q$ is brought towards **A**, which is repelled. When **B** arrives at the point previously occupied by **A**, the system is in (neutral) equilibrium.

Calculate the work done in terms of m, g, ℓ, q, Q , and k , where $k = \frac{1}{4\pi\epsilon_0}$.

[6]



Energy of System "Aft" relative to "Bef" is

$$E = mgh + \frac{2Q}{4\pi\epsilon_0 r^2}$$

Gauss's law:

$$F = \frac{2Q}{4\pi\epsilon_0 r^2}$$

This is what we want. We need to find h, r

$$\begin{aligned} l^2 &= 180^2 - 2(90^2 - \theta) \\ l &= \sqrt{180^2 - 180^2 + 2\theta} \end{aligned}$$

Isosceles triangle

NII on mass m

$$\uparrow: 0 = T_{S2\theta} + F_{S\theta} - mg \quad (1)$$

$$\rightarrow: 0 = T_{S12\theta} - F_{S\theta} \quad (2)$$

$$\therefore T_{S12\theta} = F_{S\theta}$$

$$T_{S12\theta} = mg - F_{S\theta}$$

$$\therefore \tan 2\theta = \frac{F_{S\theta}}{mg - F_{S\theta}}$$

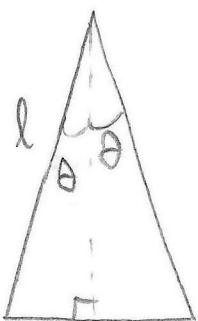
$$\text{Now } T_{S1\theta} = h$$

$$\text{and also } h = l - l \cos 2\theta$$

! 6

↓ PTD

GOAL * want to eliminate θ * $\rightarrow r = \dots, h = \dots$



$$\text{So } \frac{l}{2} = l \sin \theta \quad \therefore \quad \boxed{\sin \theta = \frac{r}{2l}} \quad *$$

Also from geometry $h = l(1 - \cos^2 \theta)$

$$\therefore \boxed{\cos^2 \theta = 1 - \frac{h}{l}}$$

* useful!

$$\text{Now } \tan 2\theta = \frac{2 \tan \theta}{1 - \tan^2 \theta}$$

$$\frac{2 \tan \theta}{1 - \tan^2 \theta} = \frac{1}{\frac{mg}{F_{ss\theta}} - \tan \theta}$$

$$1 - \tan^2 \theta = \left(\frac{mg}{F_{ss\theta}} - \tan \theta \right) \times 2 \tan \theta$$

$$1 - \tan^2 \theta = \frac{2mg \tan \theta}{F_{ss\theta}} - 2 \tan^2 \theta$$

$$1 + \tan^2 \theta = \frac{2mg \sin \theta}{F_{ss\theta}^2}$$

$$\therefore \boxed{r = \left(\frac{2\theta l}{4\pi \epsilon_0 Mg} \right)^{1/3}}$$

Since $h = rs\theta$ and $\sin \theta = \frac{r}{2l}$

$$\Rightarrow \boxed{h = \frac{r^2}{2l}}$$

useful result
e.g. in Newton's rings, "weapon of doom"...

$$\therefore \boxed{h = \frac{1}{2l} \left(\frac{2\theta l}{4\pi \epsilon_0 Mg} \right)^{2/3}}$$

$$\therefore E = \frac{Mg}{2l} \left(\frac{2\theta l}{4\pi \epsilon_0 Mg} \right)^{2/3} + \frac{2\theta l}{4\pi \epsilon_0} \left(\frac{4\pi \epsilon_0 Mg}{2\theta l} \right)^{1/3}$$

↓ PTO

$$E = \left(\frac{q^2 \alpha}{2\epsilon_0} \times \frac{1}{4\pi\epsilon_0} \right)^{2/3} \left(\frac{1}{2} + 1 \right) \left(\frac{mg}{\ell} \right)^{1/3}$$

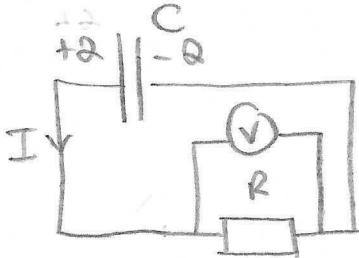
$$E = \frac{3}{2} \left(\frac{q^2 \alpha^2}{16\pi^2 \epsilon_0^2} \frac{mg}{\ell} \right)^{1/3}$$

- v) A capacitor of value 1.0 F discharges through a device whose resistance R varies linearly with applied potential difference, V , so that $R = AV + B$, where A and B are constants. The resistance of the device has a value of 10.0Ω when $V = 6.0$ V, and 4.0Ω when $V = 0.06$ V.

The capacitor is initially charged to a potential of 6.0 V. ✓

Determine how long it takes for the capacitor to discharge to 1% of this initial value.

[6]



$$V = IR$$

$$V = I(AV + B)$$

$$\therefore I = \frac{V}{AV + B} \quad \text{(1)}$$

$$\boxed{I = \frac{dQ}{dt}}$$

$$Q = CV$$

where Q is capacitor charge (separation)

$$\text{So } I = -C \frac{dV}{dt} \quad \text{(2)} \quad \therefore \text{ equating (1) and (2)}$$

$$-C \frac{dV}{dt} = \frac{V}{AV + B}$$

$$\int_{V_0}^{\sqrt{AV+B}} \frac{dV}{\sqrt{AV+B}} = -\frac{1}{C} \int_0^t dt$$

$$\int_{V_0}^{\sqrt{V}} \left(A + \frac{B}{\sqrt{V}} \right) dV = -\frac{t}{C}$$

$$A(\sqrt{V} - \sqrt{V_0}) + B \ln\left(\frac{\sqrt{V}}{\sqrt{V_0}}\right) = -\frac{t}{C}$$

$$A\sqrt{V_0}\left(\frac{\sqrt{V}}{\sqrt{V_0}} - 1\right) + B \ln\left(\frac{\sqrt{V}}{\sqrt{V_0}}\right) = -\frac{t}{C}$$

$$\therefore t = ACV_0 \left(1 - \frac{\sqrt{V}}{\sqrt{V_0}} \right) + BC \ln\left(\frac{\sqrt{V}}{\sqrt{V_0}}\right)$$

$$10 = 6A + B$$

$$4 = 0.06A + B$$

$$\therefore 6 = 5.94A$$

$$\boxed{\frac{6}{5.94} = A} \quad (\Omega V^{-1})$$

$$\therefore B = 10 - \frac{36}{5.94}$$

$$= \boxed{\frac{130}{33}} \Omega$$

↓ PTO

when $\frac{Q}{Q_0} = 0.01 \Rightarrow \frac{V}{V_0} = 0.01$ since $Q = CV$.

$$\therefore t = \frac{100}{99} \times 1.0 \times 6.0 (1 - 0.01) + \frac{130}{33} \times 1.0 \ln 100 \quad (s)$$

$$\frac{6}{5.94}$$



$$t = ACV_b \left(1 - \frac{V}{V_0}\right) + B \ln\left(\frac{V_0}{V}\right)$$

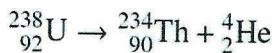
$$= 6 + 18.14$$

$$= \boxed{24.14}$$

In frame of stationary :
U nucleus

w) The $^{238}_{92}\text{U}$ decays according to

p is momentum



Determine the kinetic energy of the emitted α -particle in MeV.

(U)

BEFORE

(Th)

p_T

AFTER

(He)

p_{He}

$$KE = \frac{p^2}{2m}$$

Energy &
momentum must
be conserved

$$c = 2.998 \times 10^8 \text{ m s}^{-1}$$

Mass of the $^{238}_{92}\text{U}$ nucleus is $3.85395 \times 10^{-25} \text{ kg}$

Mass of the $^{234}_{90}\text{Th}$ nucleus is $3.78737 \times 10^{-25} \text{ kg}$

Mass of the α -particle is $6.64807 \times 10^{-27} \text{ kg}$

So Th must have
some KE.

so classical physics [5]

{Energy conservation:}

{Assume rel. speeds $\ll c$ }

$$M_U c^2 = M_T c^2 + M_{\text{He}} c^2 + \frac{p_T^2}{2M_T} + \frac{p_{\text{He}}^2}{2M_{\text{He}}}$$

{Momentum conservation}

$$p_T = p_{\text{He}} \leftarrow \text{Since in zero momentum frame.}$$

This is
what we
want.

$$\therefore \frac{p_{\text{He}}^2}{2} \left(\frac{1}{M_{\text{He}}} + \frac{1}{M_T} \right) = (M_U - M_T - M_{\text{He}}) c^2$$

$$\therefore \frac{p_{\text{He}}^2}{2M_{\text{He}}} = \frac{(M_U - M_T - M_{\text{He}}) c^2}{1 + \frac{M_{\text{He}}}{M_T}}$$

KE of
 α particle

$$= \frac{\left(3.85395 - 3.78737 - \frac{6.64807}{100}\right) \times 10^{-25} \times (3.00 \times 10^8)^2}{1 + \frac{6.64807}{100 \times 3.78737}}$$

$$= 8.77 \times 10^{-13} \text{ J} = 5.48 \text{ MeV}$$

$$[1 \text{ MeV} = 1.692 \times 10^{-13} \text{ J}]$$