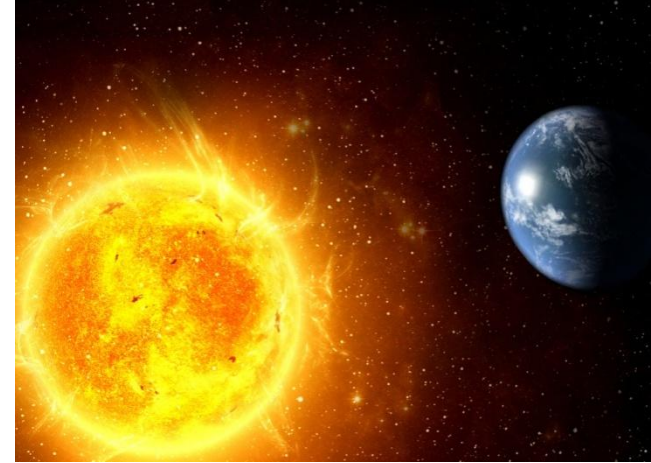


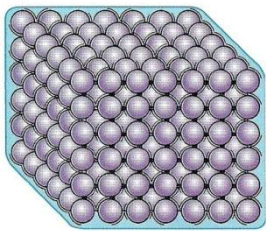


entropy

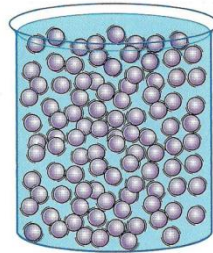
$$E = \frac{1}{2}mv^2$$



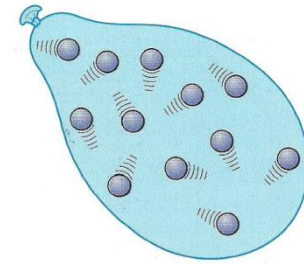
Energy & Heat



Solid Particles vibrate about fixed positions.



Liquid Particles vibrate, but can change positions.



Gas Particles move about freely.

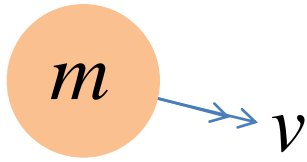
Universe by Numbers: Day 3

July 2016

Dr Andrew French



Kinetic Energy



$$E = \frac{1}{2}mv^2$$

Energy / Joules Mass Speed



$$m = 715,000\text{kg}$$

$$v = 75\text{ms}^{-1}$$

$$E = 2 \times 10^9 \text{ J}$$



seppo.net

$$m = 70\text{kg}$$

$$v = 8\text{ms}^{-1}$$

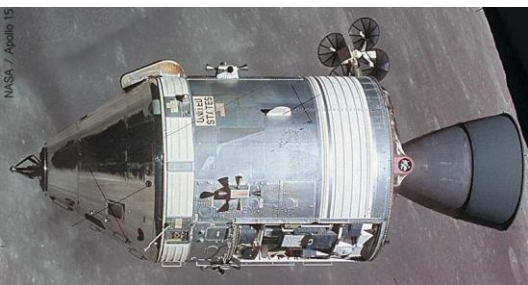
$$E = 2240\text{J}$$



$$m = 70\text{kg}$$

$$v = 31\text{ms}^{-1}$$

$$E = 33,600\text{J}$$

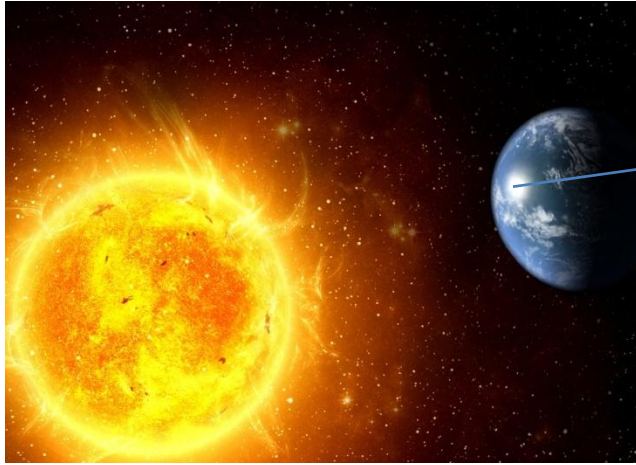


$$m = 14,690\text{kg}$$

$$v = 11,082\text{ms}^{-1}$$

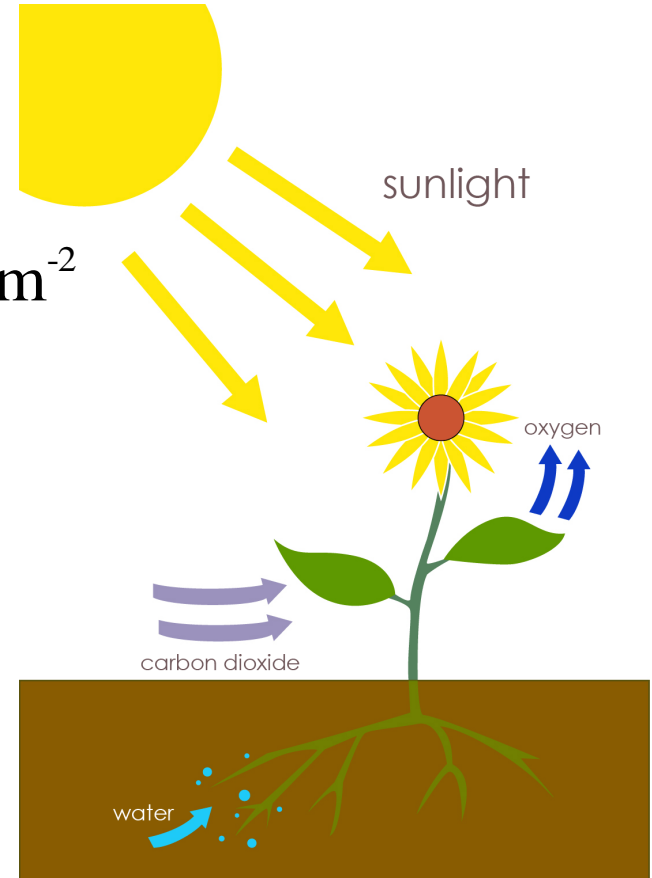
$$E = 9 \times 10^{11} \text{ J}$$

A fundamental law of Physics is that of **conservation of energy**. Energy *cannot* be created or destroyed, merely converted into other forms



$$3.8 \times 10^{26} \text{ Js}^{-1}$$

$$1.4 \times 10^3 \text{ Js}^{-1} \text{ m}^{-2}$$



Photosynthesis



$$1.7 \times 10^7 \text{ Jkg}^{-1}$$

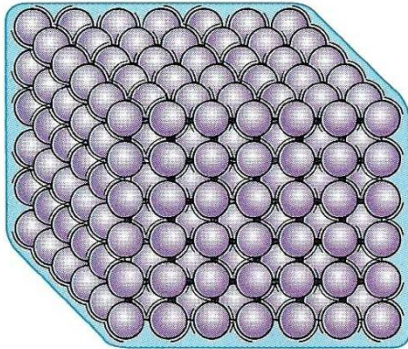
17kJ per g



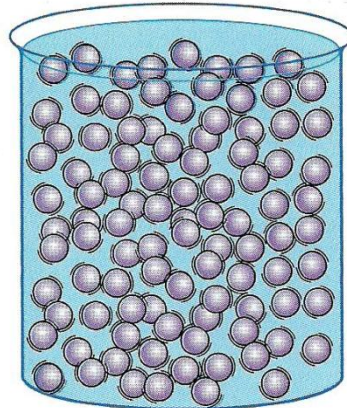
10^7 J/day Suggested calorie intake (Adult Male)

$$1\text{kcal} = 4184\text{J}$$

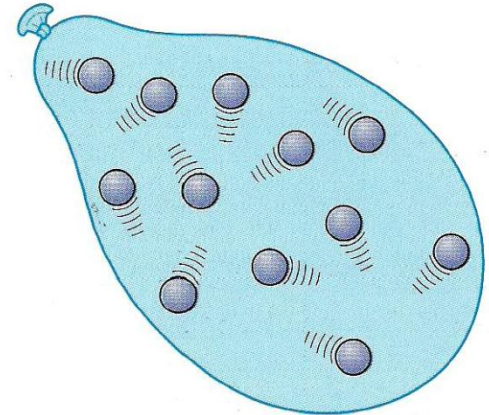
Kinetic theory: matter consists of tiny particles which are constantly in motion. They attract each other when close, but this force weakens if they move further apart. They *repel* if they get *really* close!



Solid Particles vibrate about fixed positions.

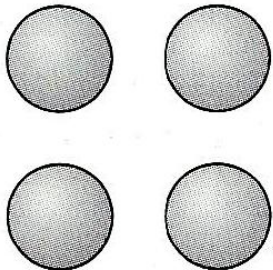


Liquid Particles vibrate, but can change positions.

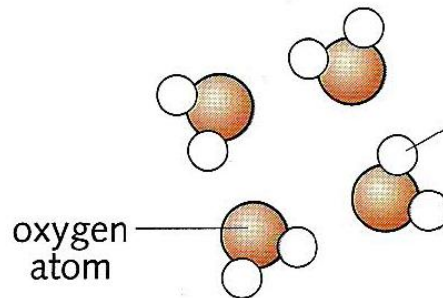


Gas Particles move about freely.

Phases of matter: Solid, Liquid or Gas ← PhET

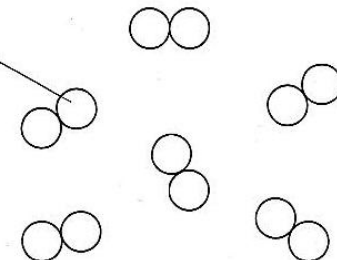


Iron atoms



Water molecules

hydrogen
atom



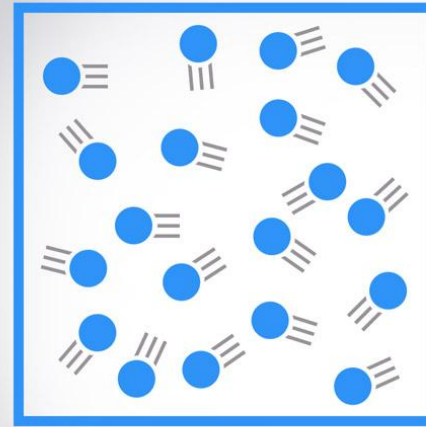
Hydrogen molecules

Heat is the **energy** that we have to put into a material to warm it up.

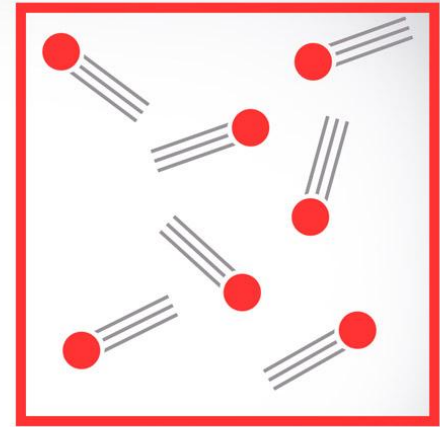
Temperature is a measure of the **average kinetic energy of each molecule**

The total energy (or 'internal energy') of the molecule is the **kinetic energy** plus the **potential energy of bonds** formed with nearby molecules

Inter-molecular bonds will be strongest in a solid, weakest in a gas



Cold Air



Hot Air

One **mole** of gas contains 6×10^{23} molecules!
12g of Carbon is one mole

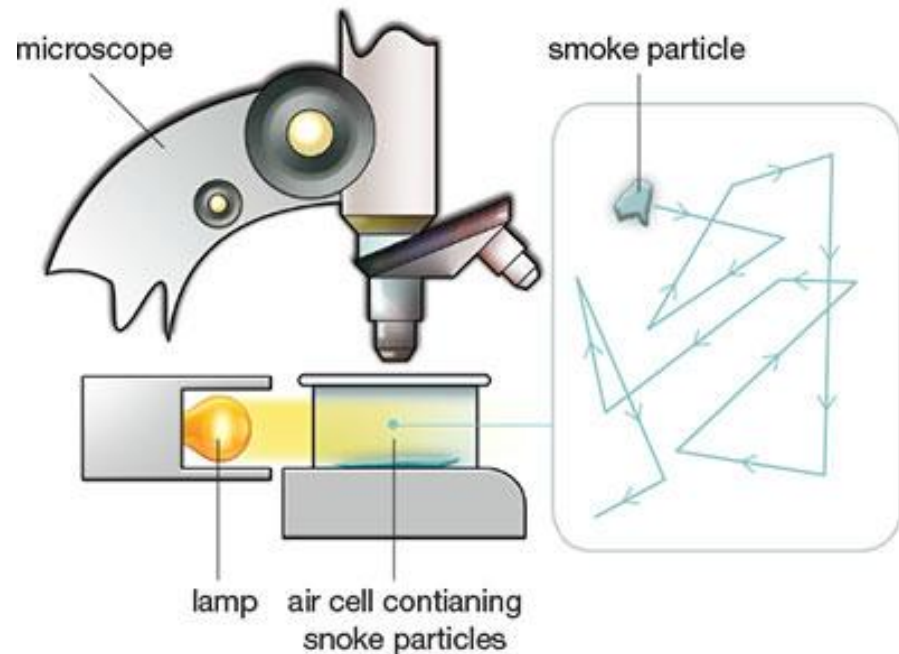
So we talk about *average* energy

How *hot* something feels is more related to *heat that can be transferred from it rather than temperature*. A dense gas will therefore feel hotter than a low density gas at the same temperature, since more molecules will collide with your hand per second (and therefore transfer their energy) when the gas is more dense.

Temperature is not the same as heat. For example, a spoonful of boiling water has exactly the same temperature (100°C) as a saucepanful of boiling water, but you could get far less thermal energy (heat) from it.

Brownian Motion is named after Robert Brown, who investigated the seemingly random motion of pollen grains in water.

The motion is a motivation for the **kinetic theory**, that is the zig-zag movement of the grains (or indeed smoke particles in the experiment described below) is caused by the random jostling of *smaller, unseen, molecules*. e.g. which comprise water, air etc.



http://www.igcsephysics.com/sample/thermal_physics/page86.html

Robert Brown
1773 -1858

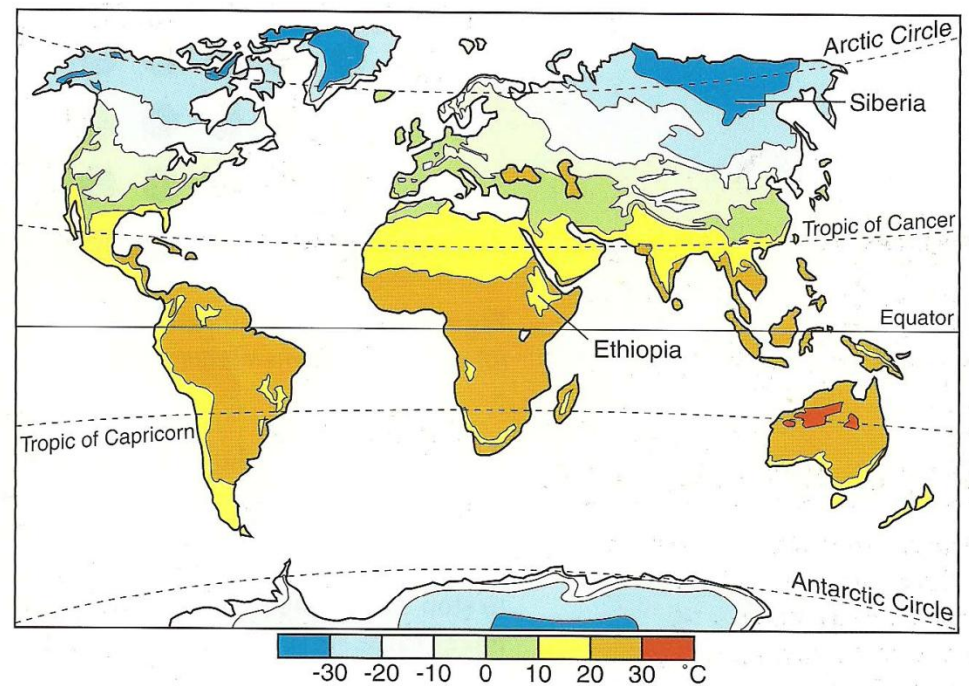


One **mole** of gas contains 6×10^{23} molecules!
The complexity of such a vast number of collisions
is why random motion is observed.

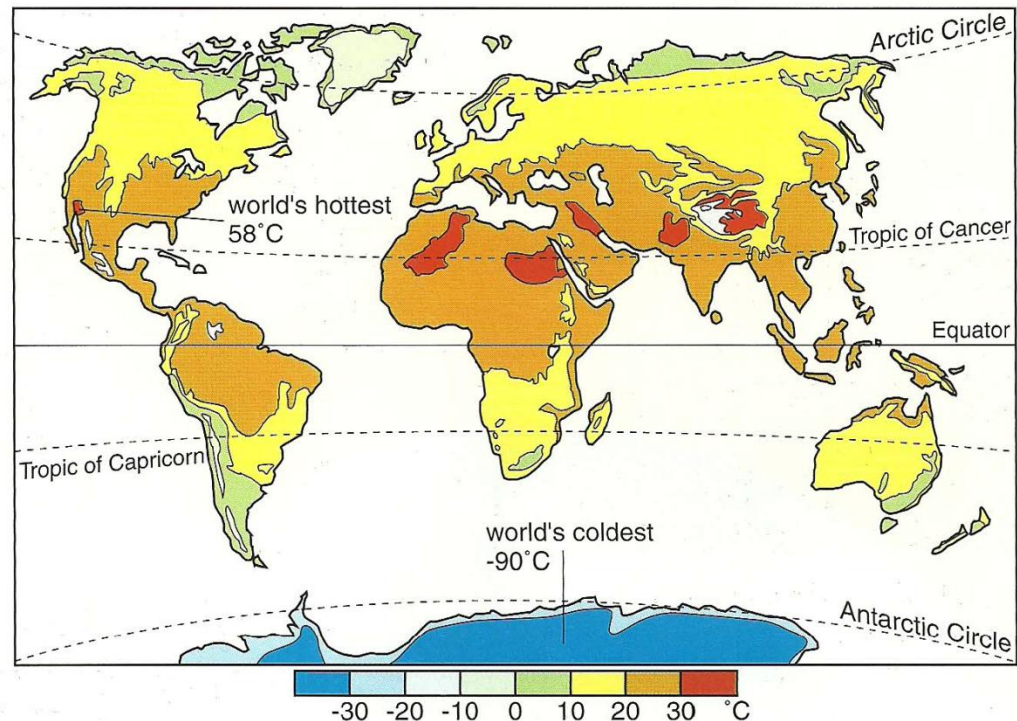
| | |
|-----------------|---------------|
| Sun's centre | 15 000 000 °C |
| Sun's surface | 6000 °C |
| bulb filament | 2500 °C |
| bunsen flame | 1500 °C |
| boiling water | 100 °C |
| human body | 37 °C |
| warm room | 20 °C |
| melting ice | 0 °C |
| food in freezer | -18 °C |
| liquid oxygen | -180 °C |
| absolute zero | -273 °C |

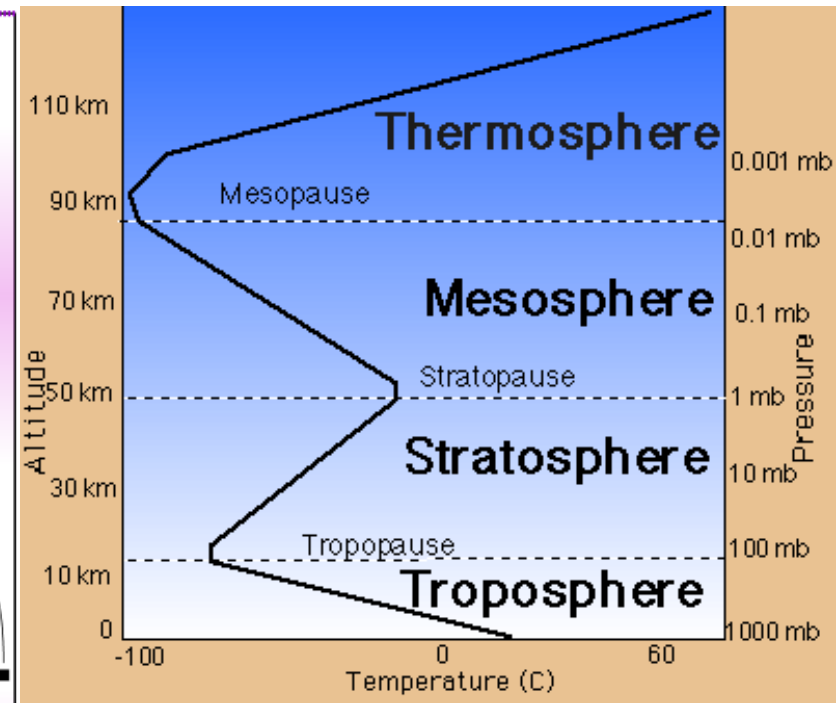
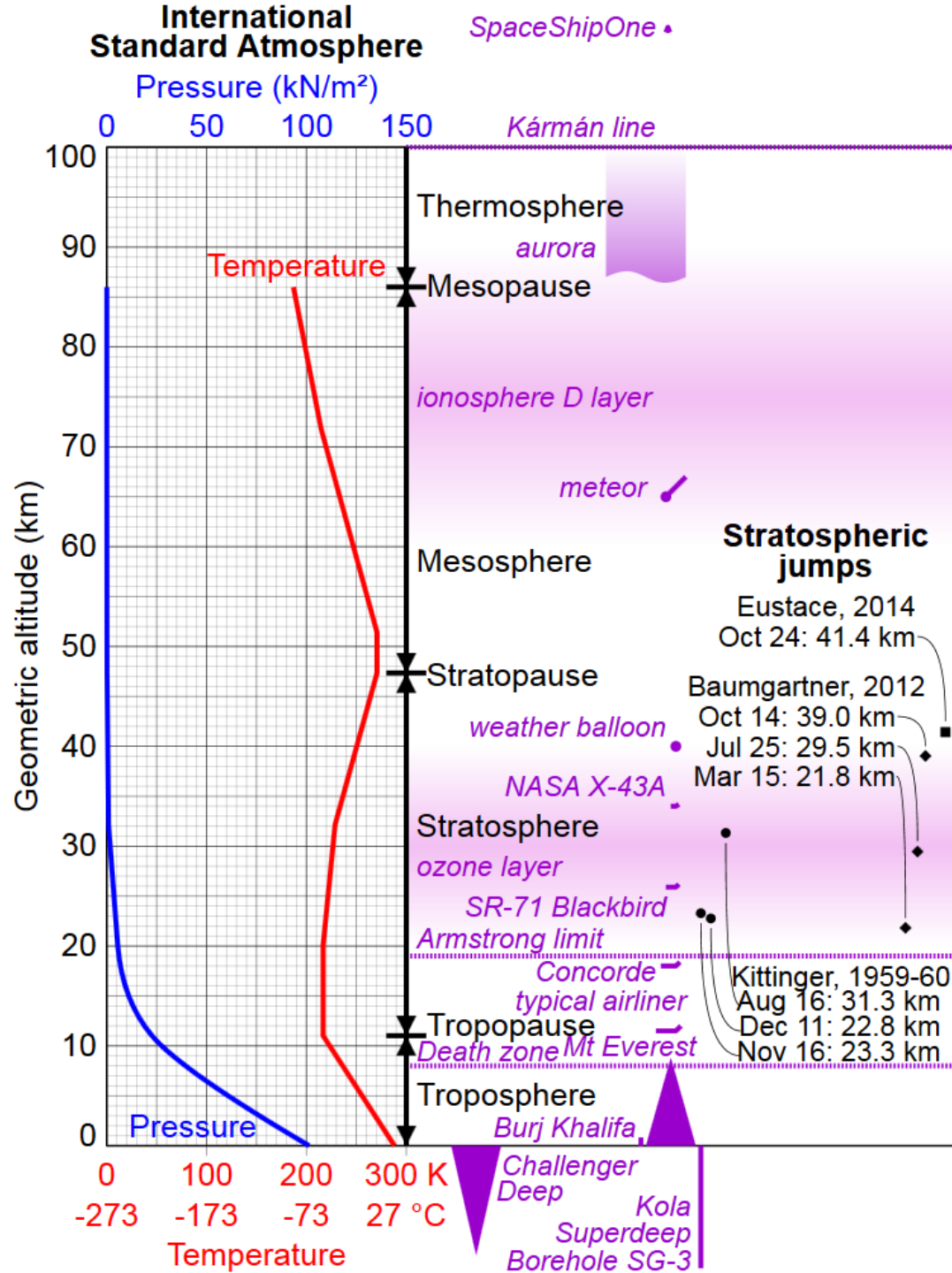
Temperature – the average energy of molecules

Jan



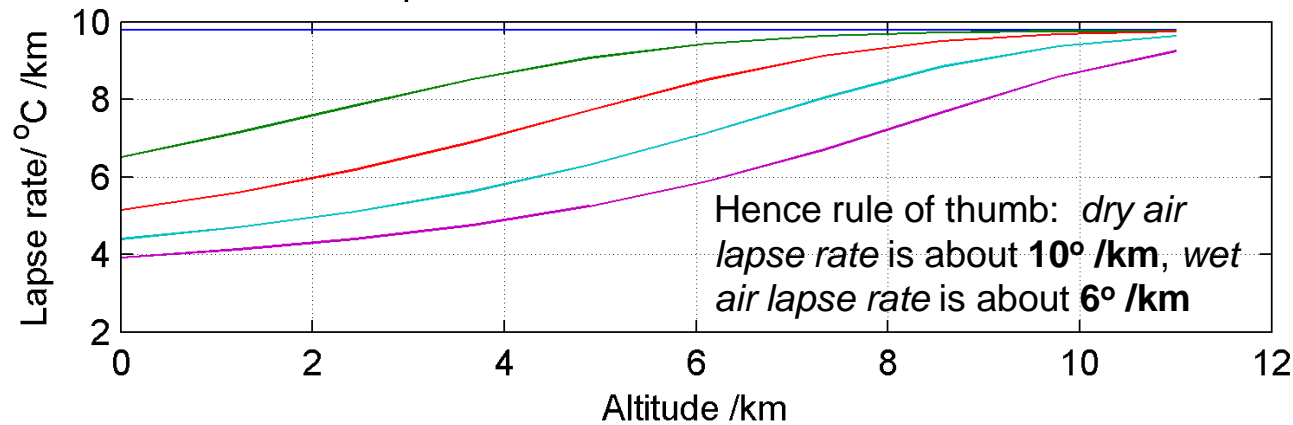
July



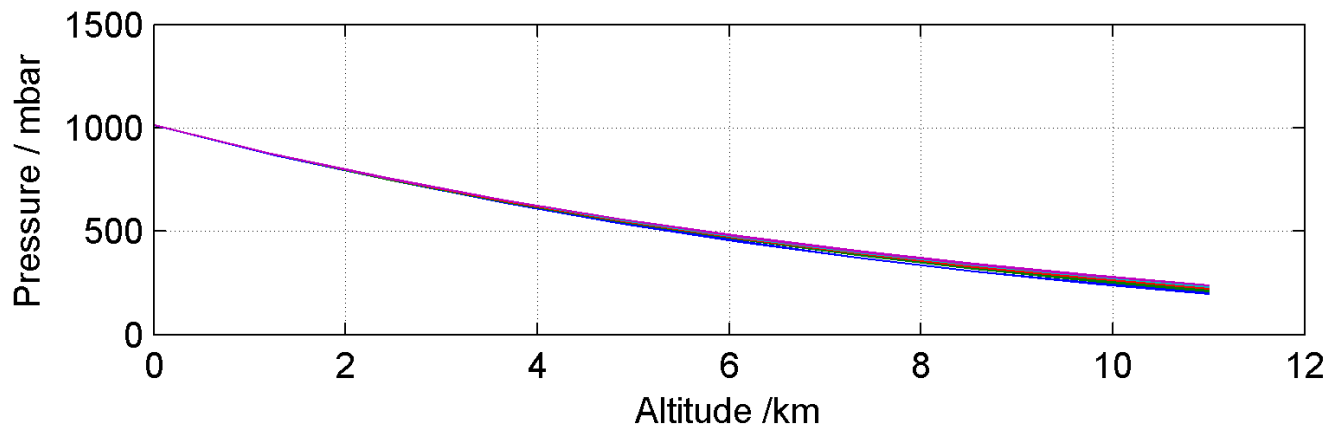
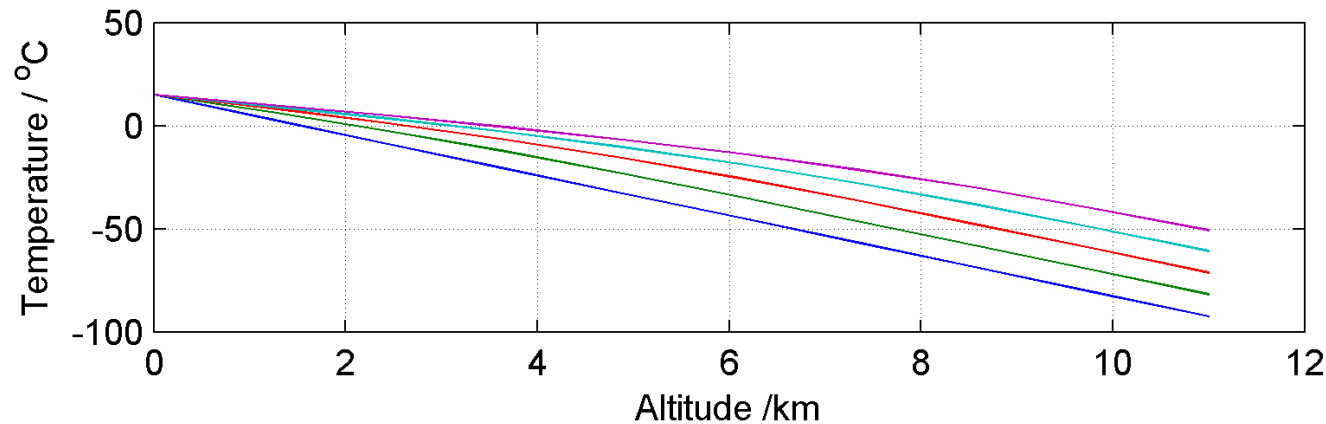


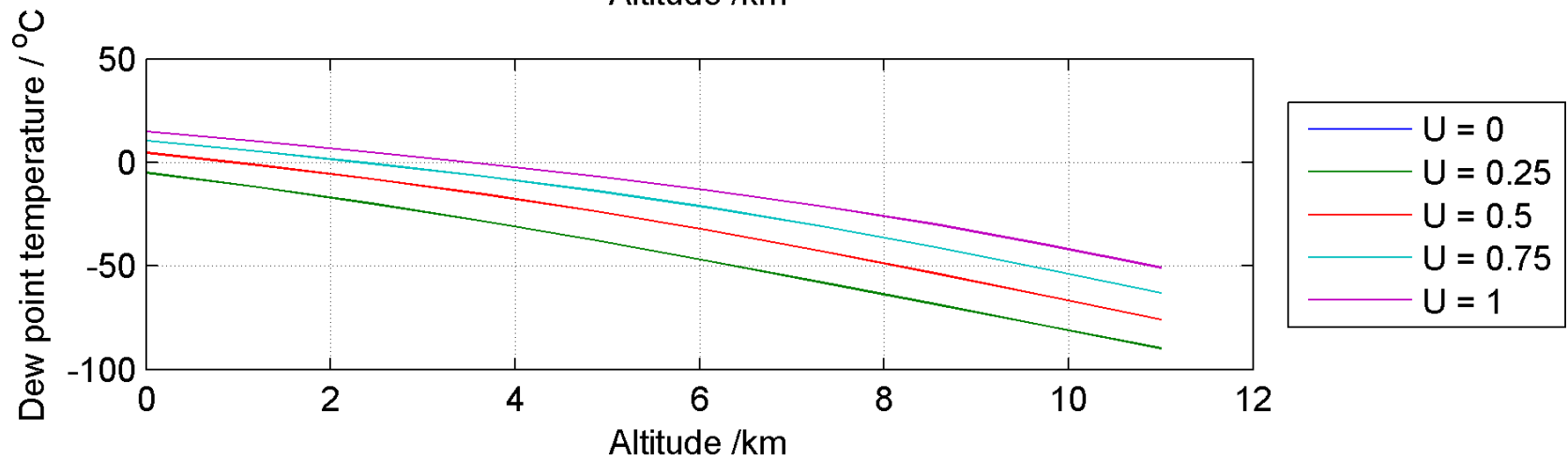
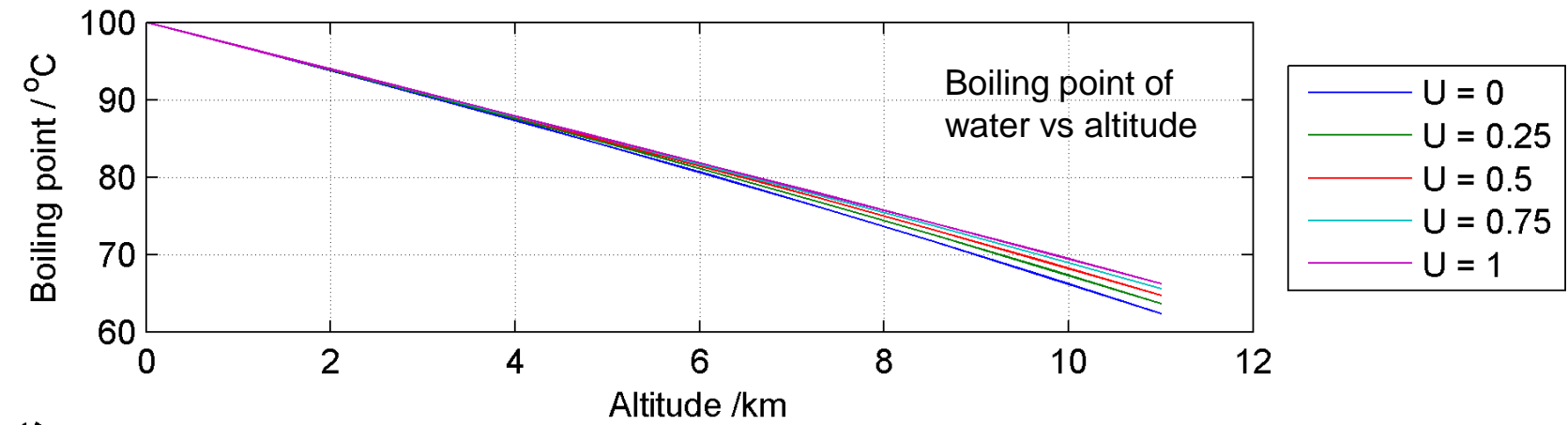
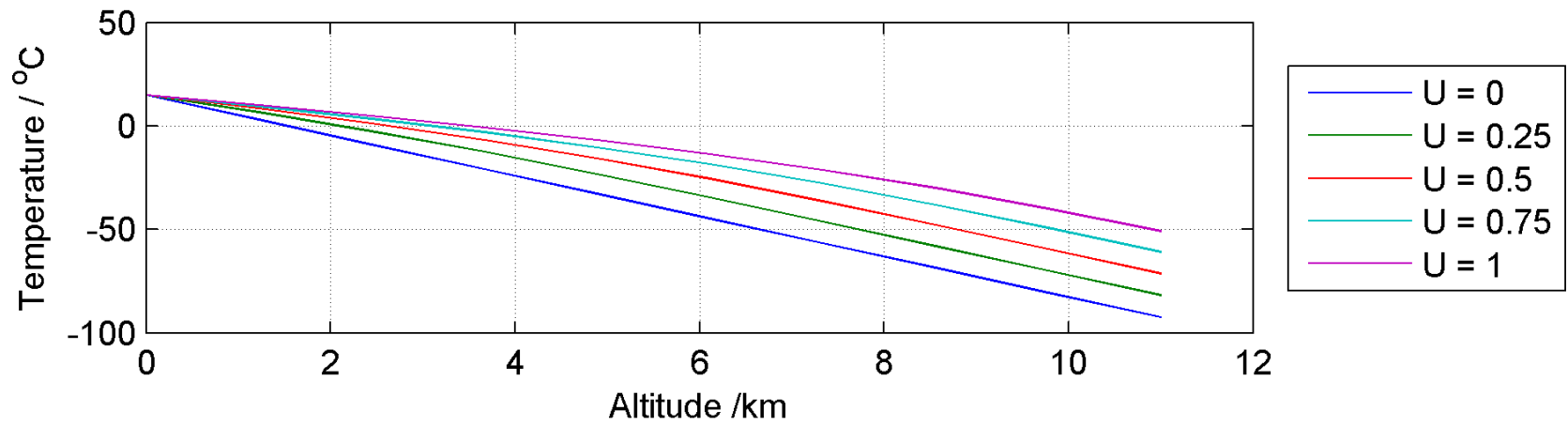
Note in the *thermosphere* the temperature (i.e. mean kinetic energy of air molecules) can rise to *several thousand degrees Celsius*. However, the air density is *extremely low*, which means it would still feel very cold indeed, as very little heat could be extracted from a given volume of air.

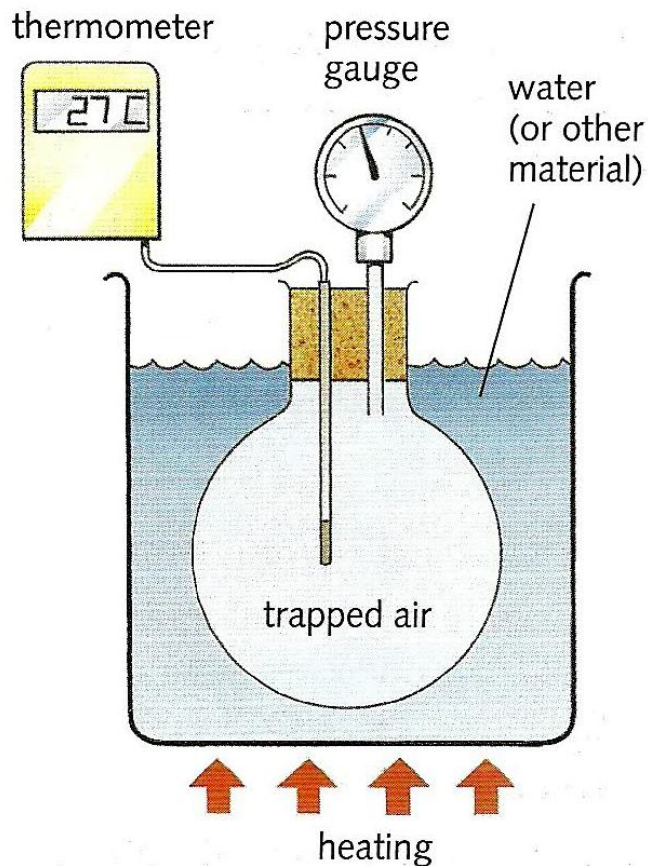
Lapse rates for different relative humidities



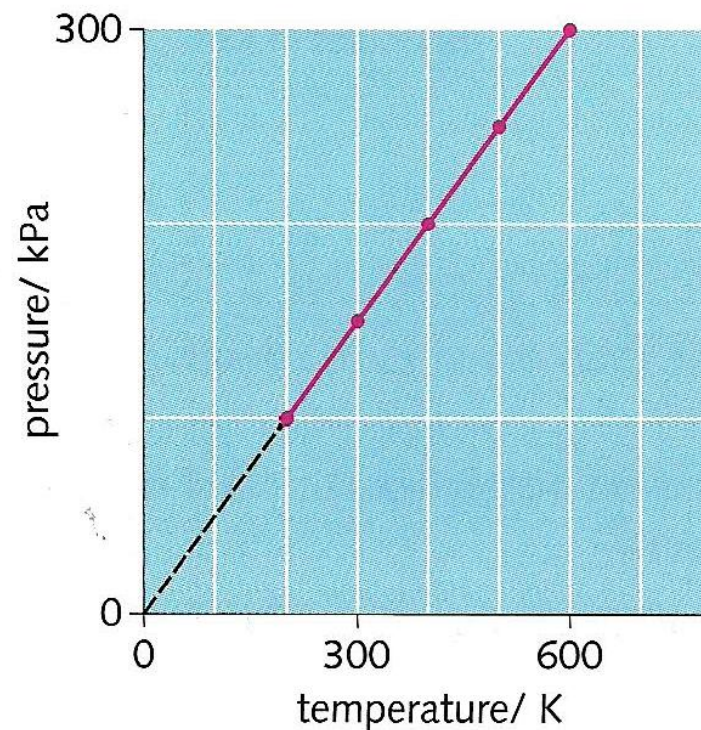
$U = 1$ means 'wet air'
 $U = 0$ means dry air







| pressure | temperature | temperature |
|----------|-------------|-------------|
| kPa | °C | K |
| 100 | −73 | 200 |
| 150 | 27 | 300 |
| 200 | 127 | 400 |
| 250 | 227 | 500 |
| 300 | 327 | 600 |



At **constant volume**, an *ideal* gas obeys the equation

$$P \propto T$$

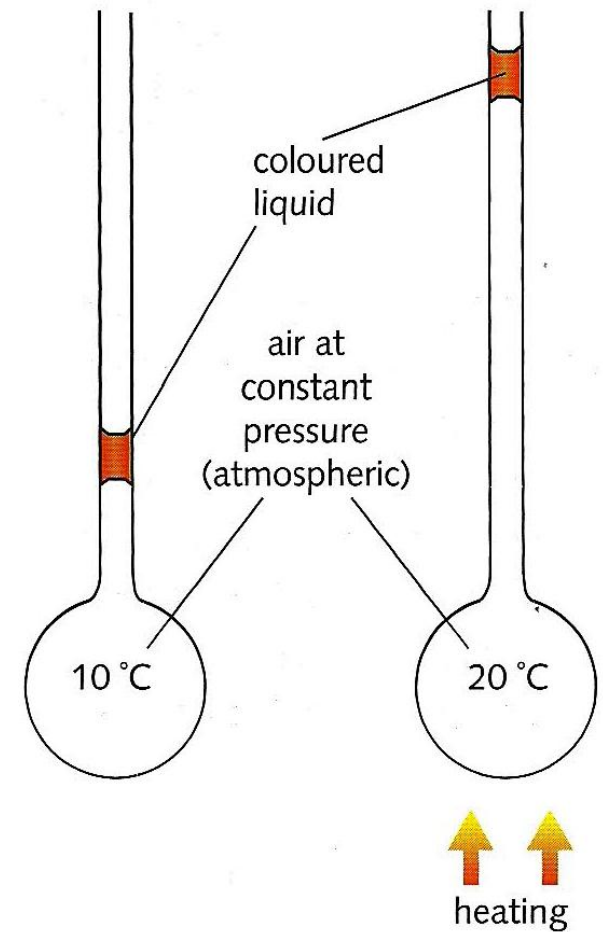
i.e. we don't worry about the gas inter-molecule forces. They collide and transfer energy elastically.

Charles' Law states

At **constant pressure**, an *ideal* gas obeys the equation

$$V \propto T$$

Jacques Charles (1746-1823)

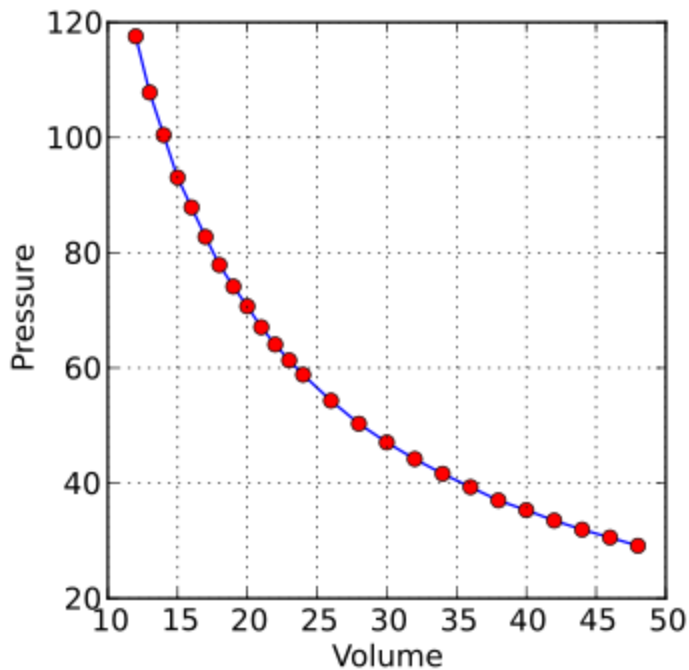


Robert Boyle (1627–91)

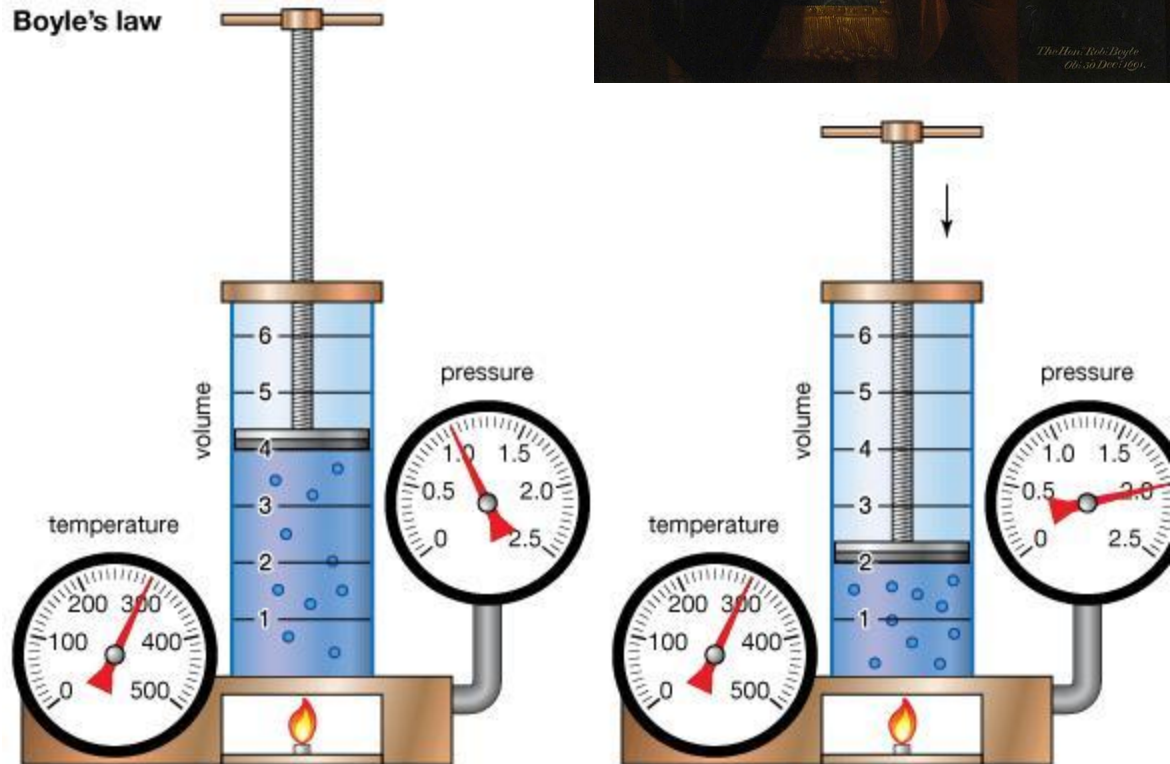
Boyle's Law states

At **constant temperature**, an *ideal* gas obeys the equation

$$P \propto \frac{1}{V}$$



Boyle's law



The Ideal Gas Equation

$$PV = nRT$$

Temperature
in Kelvin



Pressure in
Newtons / square
metre
(Pascals or Pa)

Gas
volume in
cubic
metres

Number of
moles of
gas

Molar Gas constant
 $R = 8.314 \text{ Jmol}^{-1}\text{K}^{-1}$

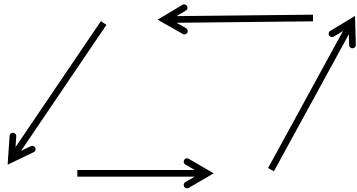
Avogadro's number
 6.02×10^{23} molecules per mole

1 atmosphere = 101,325 Pa
1 bar = 100,000 Pa
1 mbar = 100 Pa

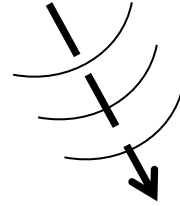
Note real gases are *not* ideal – when molecules get close enough there are **intermolecular forces** which will significantly effect the motion of the molecules. However, the 'ideal gas' approximation is typically a good one in most practical situations.

Conduction

Heat Transfer

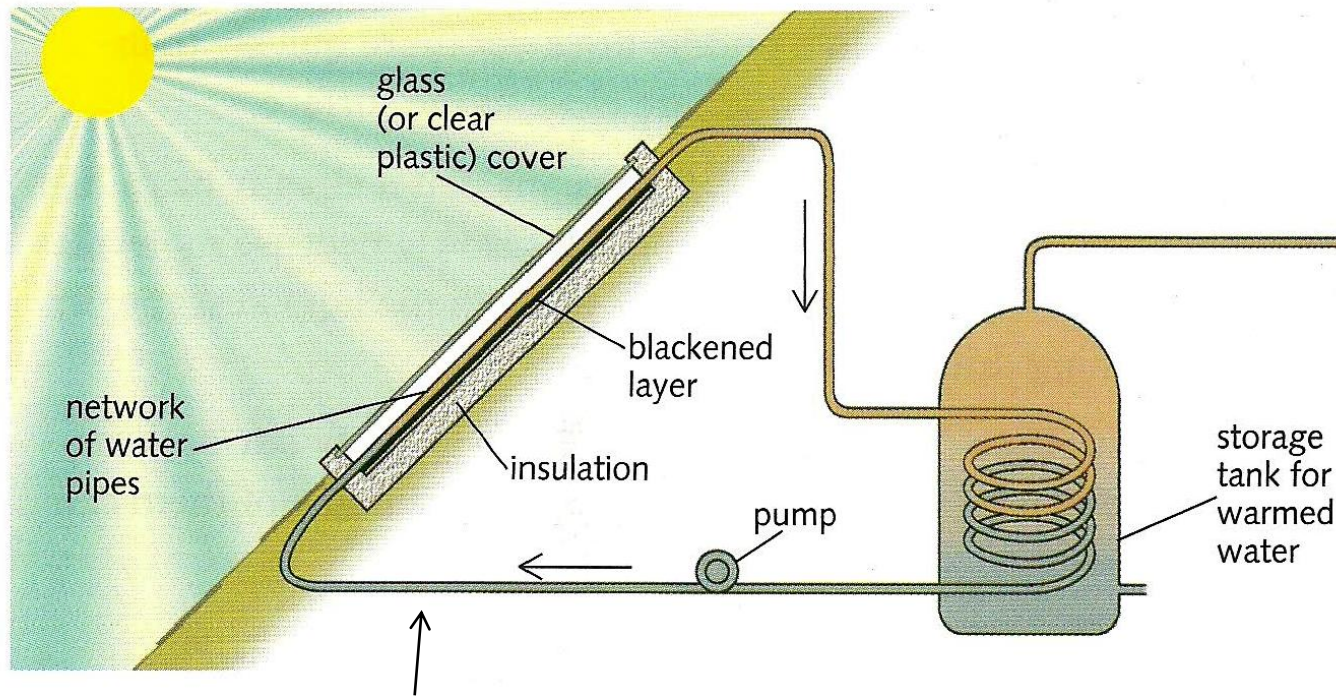


Convection



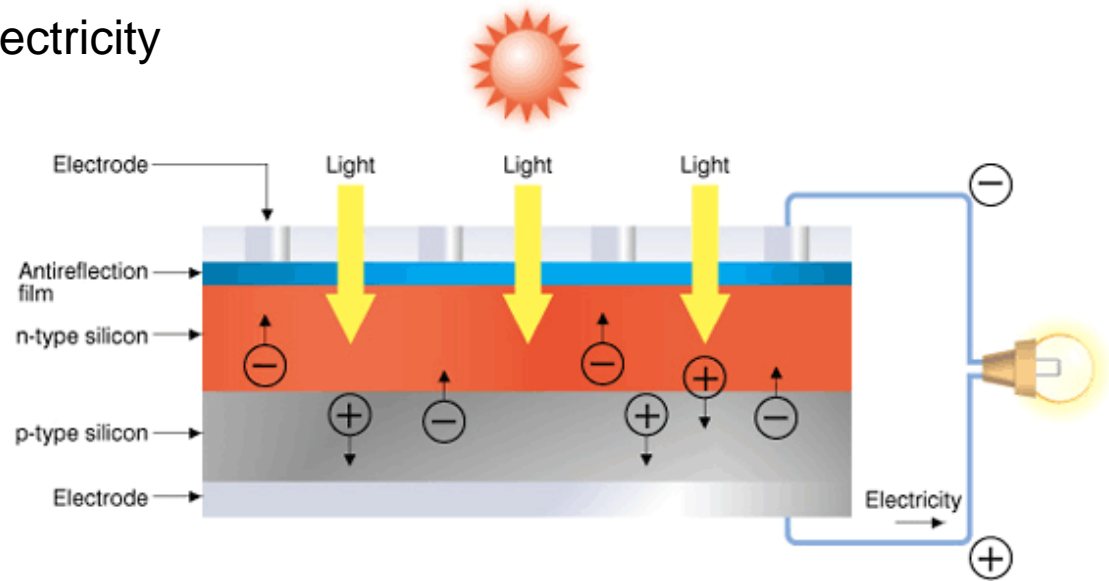
Radiation

The Solar Panel

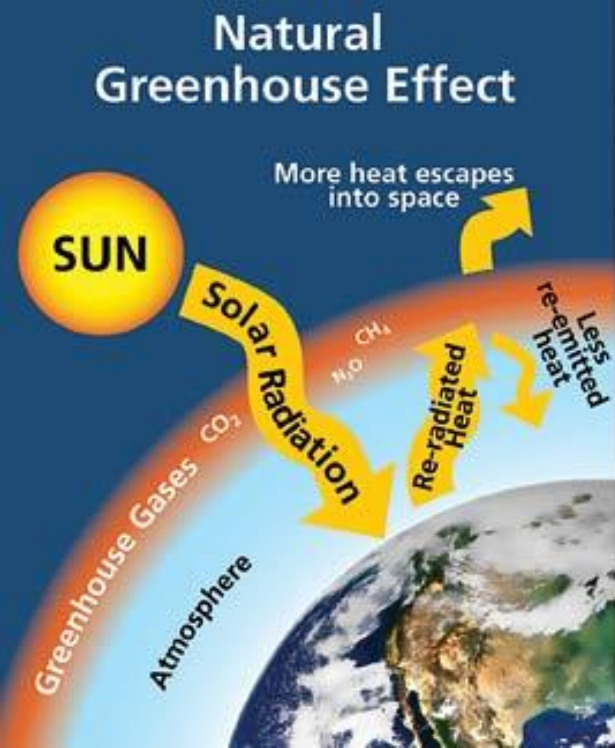


This is a *passive* solar heater.
An *active* solar panel generates electricity
directly using a photovoltaic cell

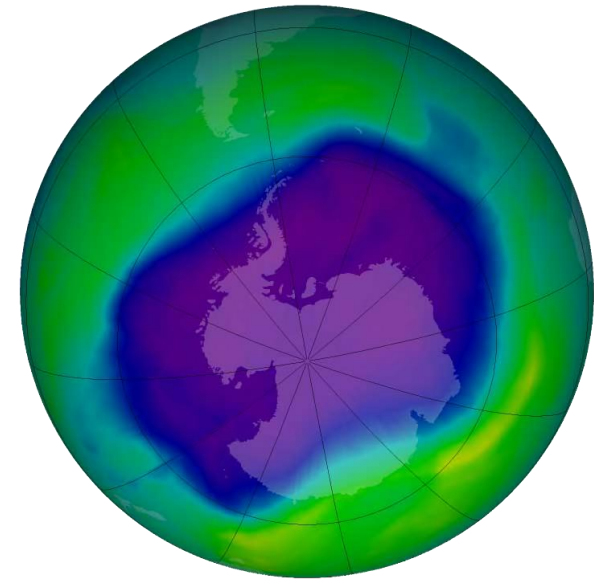
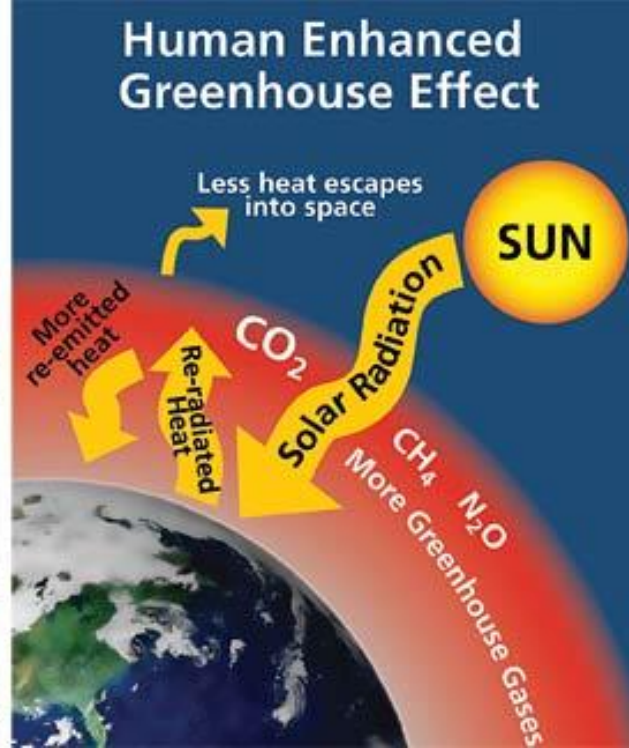
How a Photovoltaic cell generates power



Natural Greenhouse Effect



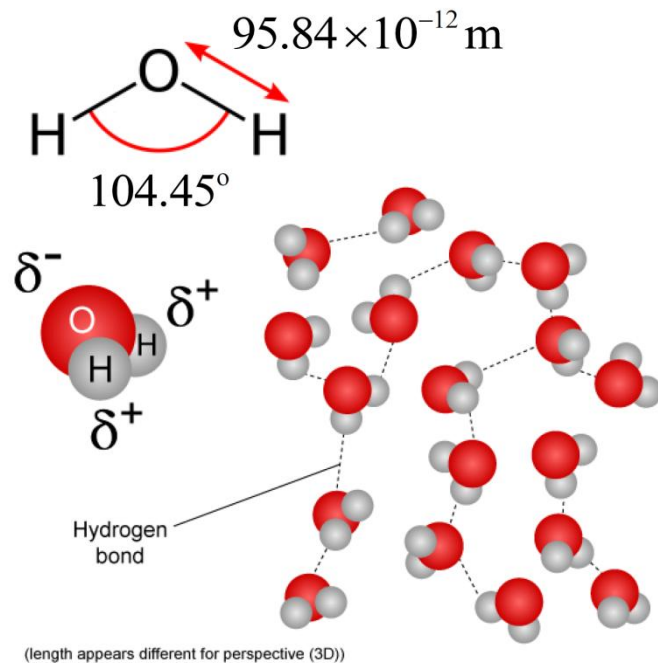
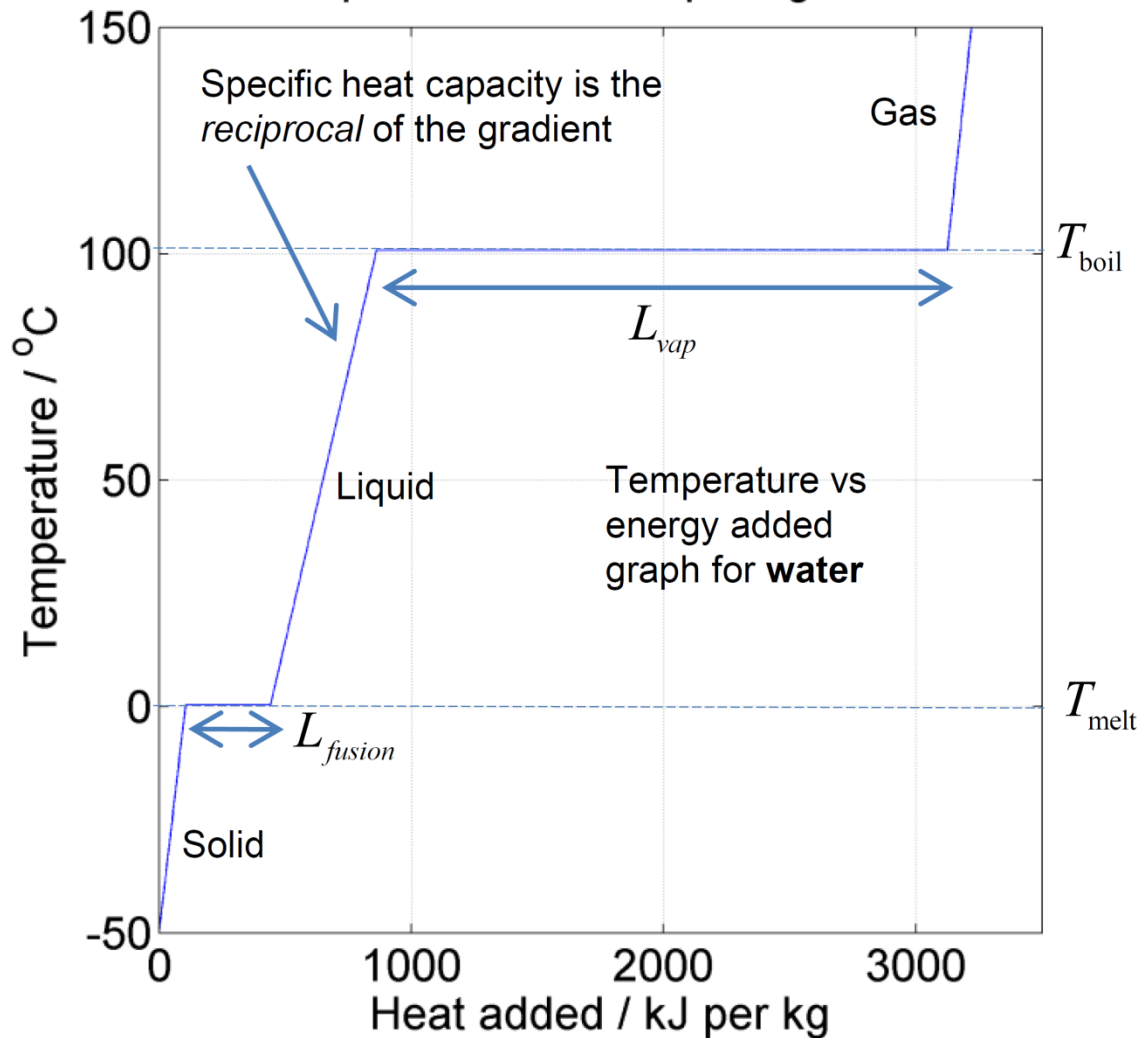
Human Enhanced Greenhouse Effect



[The Eden Project](#)



Temperature vs heat per kg added



| | Melting temp. / °C | Boiling temp. / °C | Latent heat of fusion / $\times 10^4 \text{ Jkg}^{-1}$ | Latent heat of vaporization / $\times 10^5 \text{ Jkg}^{-1}$ | Specific heat capacity (solid) / $\text{Jkg}^{-1}\text{K}^{-1}$ | Specific heat capacity (liquid) / $\text{Jkg}^{-1}\text{K}^{-1}$ | Specific heat capacity (gas) / $\text{Jkg}^{-1}\text{K}^{-1}$ |
|-------|--------------------|--------------------|--|--|---|--|---|
| Water | 0 | 100 | 33.5 | 22.6 | 2,090 | 4,186 | 1,930 |

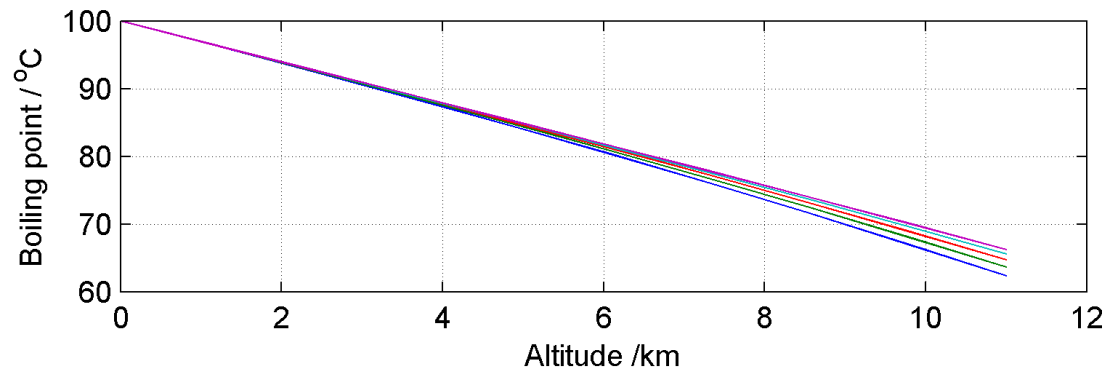
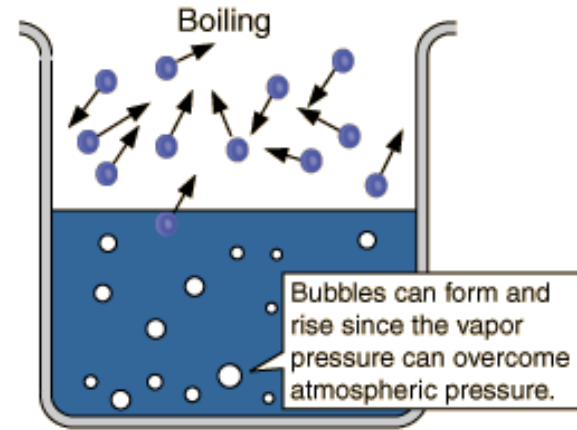
Liquid water will contain bubbles of **vapour**. At 20°C at sea level these are tiny, due to the pressure of the atmosphere and the relatively modest temperature. The bubbles will expand if:

- (i) the atmospheric pressure drops or
- (ii) the temperature rises

Rapid mass bubble expansion, followed by bursting and vapour release is in practical terms what **boiling** is i.e. the **conversion of liquid to gas**.

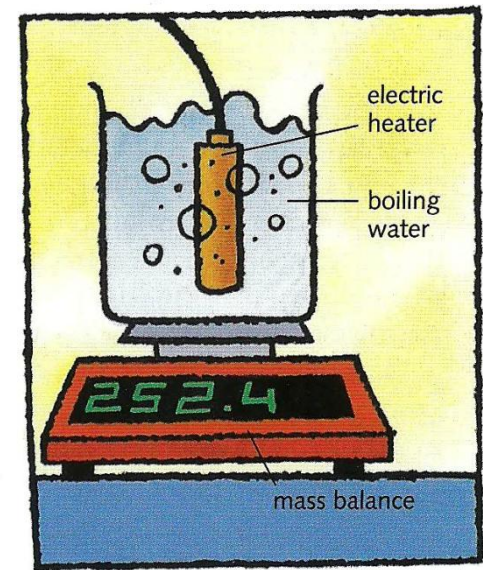
At 100°C at sea level, water will boil. Due to the lowering of atmospheric pressure with altitude, the boiling temperature reduces. At the summit of Mount Everest (8848m) water boils at about 70°C.

Boiling

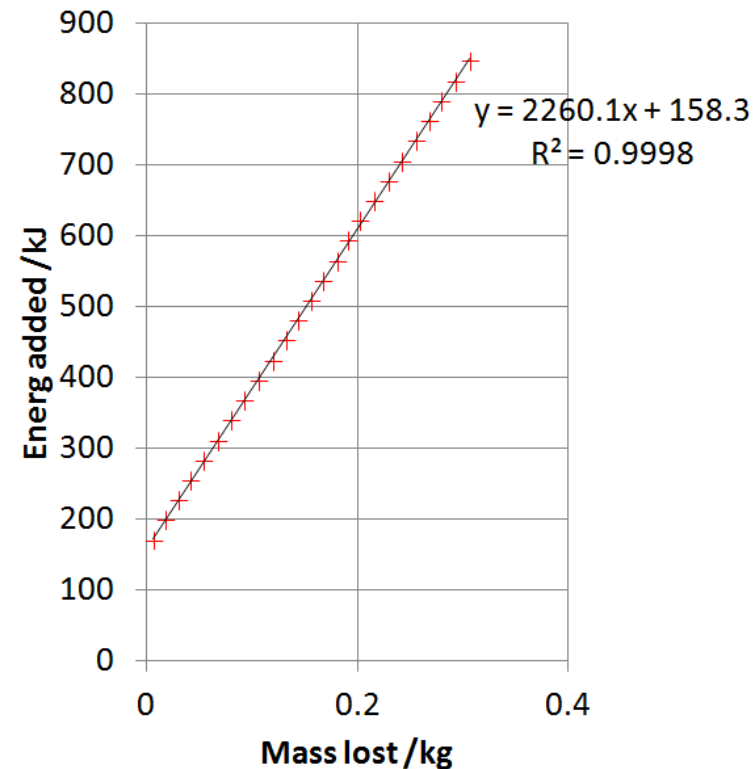


Voltage 235.00
Current 4.00

Power /kW 0.94



Latent heat of vaporization



BOIL

| t/s | Energy added /kJ | mass lost /g | mass lost /kg |
|-----|------------------|--------------|---------------|
| 0 | 0 | 0 | 0 |
| 30 | 28.2 | -0.4 | -0.0004 |
| 60 | 56.4 | 0.3 | 0.0003 |
| 90 | 84.6 | 0.9 | 0.0009 |
| 120 | 112.8 | 0.7 | 0.0007 |
| 150 | 141 | 1.1 | 0.0011 |
| 180 | 169.2 | 5.9 | 0.0059 |
| 210 | 197.4 | 16.8 | 0.0168 |
| 240 | 225.6 | 29 | 0.029 |
| 270 | 253.8 | 41.4 | 0.0414 |
| 300 | 282 | 53.7 | 0.0537 |
| 330 | 310.2 | 67.5 | 0.0675 |
| 360 | 338.4 | 80 | 0.08 |
| 390 | 366.6 | 92.2 | 0.0922 |
| 420 | 394.8 | 105.2 | 0.1052 |
| 450 | 423 | 118.9 | 0.1189 |
| 480 | 451.2 | 132.1 | 0.1321 |
| 510 | 479.4 | 142.9 | 0.1429 |
| 540 | 507.6 | 155.2 | 0.1552 |
| 570 | 535.8 | 166.5 | 0.1665 |
| 600 | 564 | 179.9 | 0.1799 |
| 630 | 592.2 | 190 | 0.19 |
| 660 | 620.4 | 201.8 | 0.2018 |
| 690 | 648.6 | 215.4 | 0.2154 |
| 720 | 676.8 | 228.5 | 0.2285 |
| 750 | 705 | 241.3 | 0.2413 |



$$S = k_B \ln W$$

"Number of
ways of
arranging
energy"

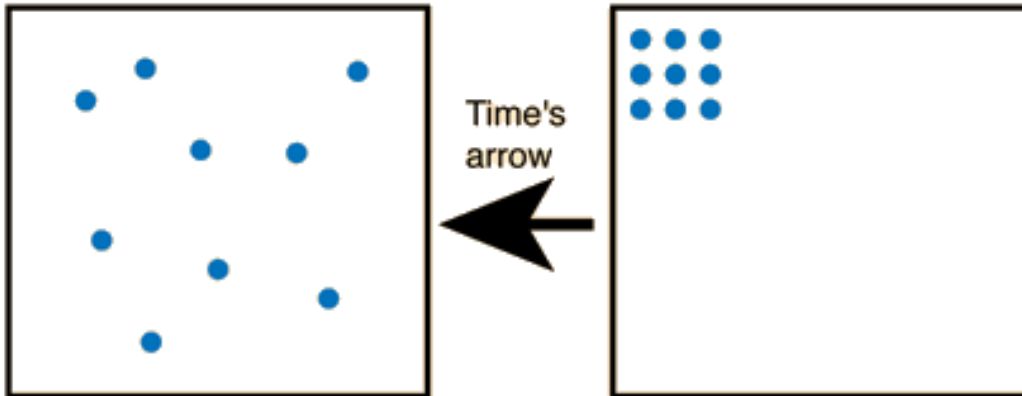


Ludwig
Boltzmann
1844-1906

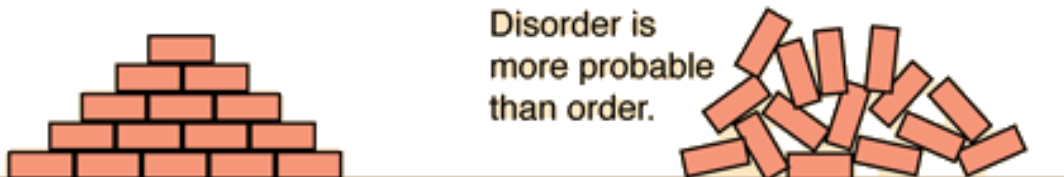


Rudolf Clausius
1822–1888

If the particles represent gas molecules at normal temperatures inside a closed container, which of the illustrated configurations came first?



If you tossed bricks off a truck, which kind of pile of bricks would you more likely produce?



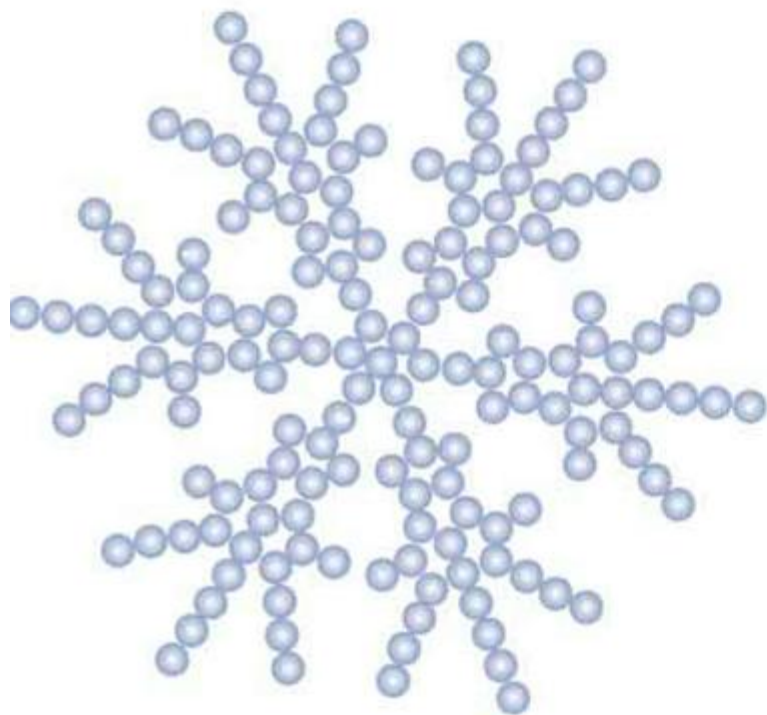
Second Law of Thermodynamics

**For all processes, the overall
entropy of the universe must
increase**

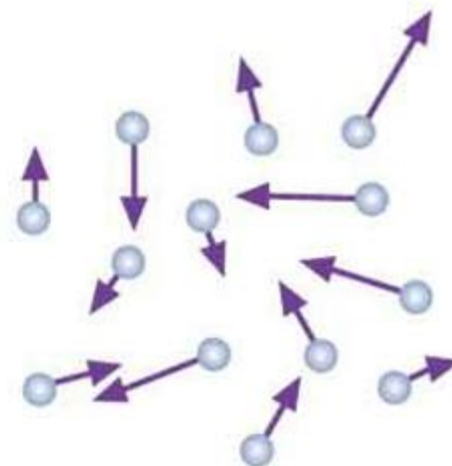
Which is why **tidying up**
requires **work!**

Order

Disorder

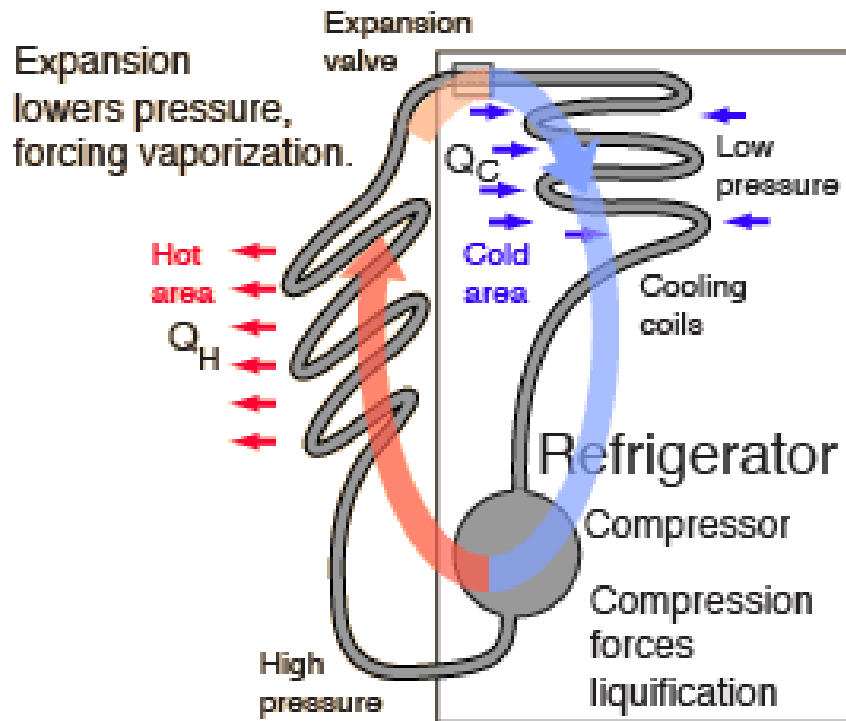


Ice

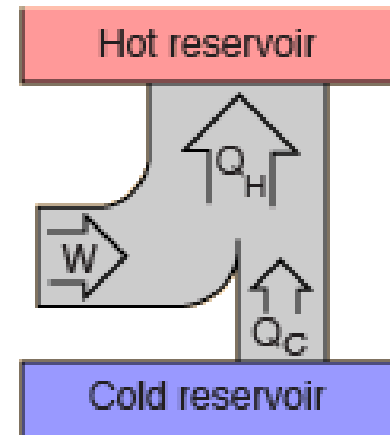


Water

entropy



All real refrigerators require work to get heat to flow from a cold area to a warmer area.



In order to cool a fridge we need to 'pump' heat from a cold region (inside the fridge) to a hotter region (surrounding air).

To satisfy the *Second Law of Thermodynamics* we must **do work** to make this happen

A flow of heat from hot to cold regions will happen *without* a work input!



Topics to reflect on:

Energy. Formula for **kinetic energy**. **Conservation of energy**.

Kinetic theory: The **molecular** model of **heat**. Difference between **heat** (total energy flow) and **temperature** (average kinetic energy per molecule).

States of matter: **Solid**, **Liquid** and **Gas**

Heat transfer: Conduction, convection and radiation

The **atmosphere** and the **Greenhouse effect**

The **Ideal gas: Pressure, Volume, Temperature**

Specific heat and boiling

Entropy – a measure of disorder

Depending on your course, we may not cover all of these. Review the topics you did meet. If you have time to spare, read on!