

$S$  Lorentz transform relates distance  $x$  and time  $t$  between 'static' frame  $S$  and moving frame  $S'$ , which moves  $\parallel x$  at velocity  $V$

$$t = \gamma \left( t' + \frac{Vx'}{c^2} \right)$$

$$\gamma = \left( 1 - \frac{v^2}{c^2} \right)^{-\frac{1}{2}}$$

Now if clock is stationary in  $S'$ ,  $x' = 0$

$$\therefore t = \gamma t'$$

let  $t' = \frac{1}{2}t \quad \therefore t = \gamma \frac{1}{2}t \Rightarrow \boxed{\gamma = 2}$

i.e. "moving clock runs at half the rate of an equivalent clock in  $S$ "

$$\Rightarrow 1 - \frac{v^2}{c^2} = \frac{1}{4} \Rightarrow \frac{v^2}{c^2} = \frac{3}{4}$$

$$\Rightarrow \boxed{v = \frac{\sqrt{3}}{2}c} \quad \approx \boxed{0.866c}$$

2,

A metre stick travels at speed  $V$  relative to frame  $S$   $\parallel$  to the  $x$  direction. In a frame  $S'$  where the metre stick is stationary

$$\boxed{x' = \gamma(x - vt)}$$
 from the Lorentz Transform

Now  $x' = 1$  metre and let stick length be measured in  $S$  when  $t = 0$ .

Here  $1 = \gamma x \quad \therefore x = \frac{1}{\gamma}$  i.e. 'lengths contract'

Now if  $x = 0.9$  m

$$\gamma = \frac{10}{9}$$

$$\left( 1 - \frac{v^2}{c^2} \right)^{-\frac{1}{2}} = \frac{10}{9}$$

$$1 - \frac{v^2}{c^2} = \frac{81}{100}$$

$$\frac{v^2}{c^2} = 1 - \frac{81}{100}$$

$$\frac{v^2}{c^2} = \frac{19}{100}$$

$$\boxed{v = \frac{\sqrt{19}}{10}c}$$

$$\approx \boxed{0.436c}$$

①

3/ let  $t' = 26 \text{ ns}$  be lifetime of the  $\pi$  meson in its rest frame.  
 Hence taking the Earth to be a frame  $S$  and the meson to be at rest in a frame  $S'$  moving at speed  $v = 0.95c$  relative to  $S$  in the  $x$  direction

Lorentz transform  $\Rightarrow t = \gamma(t' + vx'/c^2)$   $x' = 0$

so  $t = (1 - 0.95^2)^{-\frac{1}{2}} \times 26 \text{ ns}$

$t = 3.20 \times 26 \text{ ns}$

$t = 83.3 \text{ ns}$

so meson will travel  $\approx 0.95c \times 83.3 \text{ ns} \approx 23.7 \text{ m}$

before it decays.

4/ Atomic clock measures time  $t' = 3600 \text{ s}$  on-board a jet moving at  $400 \text{ m s}^{-1}$

on the ground, the equivalent time interval is  $t = \gamma t'$

(Note  $x' = 0$  as it is assumed the clock is at rest relative to the  $S'$  frame co-moving with the aircraft)

$t = \left(1 - \left(\frac{400}{2.9979 \times 10^8}\right)^2\right)^{-\frac{1}{2}} \times 3600$

Now calculator precision problems working this out directly

So let  $t = 3600 + \delta t$ , and since  $\frac{v^2}{c^2} \ll 1$

use a single term Binomial expansion

$t \approx t' \left(1 + (-\frac{1}{2})(-\frac{v^2}{c^2}) + \frac{(-\frac{1}{2})(-\frac{3}{2})(-\frac{v^2}{c^2})^2}{2} + \dots\right)$

$\approx t' \left(1 + \frac{v^2}{2c^2}\right)$

$\delta t \approx \frac{t' v^2}{2c^2}$

$\delta t = \frac{3600 \times 400^2}{2 \times (2.9979 \times 10^8)^2} \approx 3.2 \text{ ns}$

②

5/ observed lifetime of muons is  $\tau = \gamma \tau'$   
 where  $\tau'$  is lifetime in rest frame of the muon

if the muon travels at speed  $0.95c$

$$\tau = (1 - 0.95^2)^{-\frac{1}{2}} \tau' \quad \tau' = 2.2 \mu\text{s}$$

$$\therefore \tau = 7.105 \mu\text{s}$$

# of muons after time  $t$  is  $N = N_0 e^{-t/\tau}$

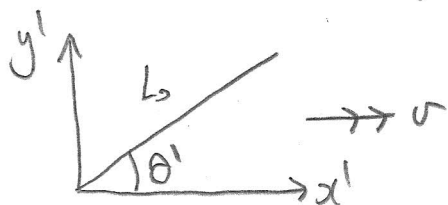
so if muons travel distance  $x$ ,  $t = x/0.95c$

$$\therefore \frac{N}{N_0} = \exp\left(\frac{-x}{0.95c \times 7.105 \times 10^{-6}}\right)$$

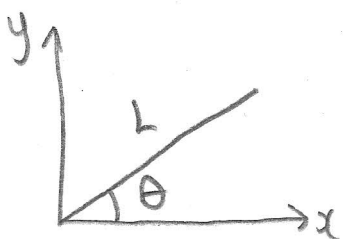
let  $x = 3000 \text{ m}$   $\therefore \frac{N}{N_0} = \exp\left(\frac{-3000}{0.95 \times 2.998 \times 10^8 \times 7.105 \times 10^{-6}}\right)$   
 $= 0.224$

ie only about  $22\%$  left

6/ In rest frame of rod ( $S'$ ), which moves at speed  $v$  with respect to a frame  $S$ .



The rod is photographed in  $S$  and looks like



using Lorentz transform  $\begin{cases} y' = y \\ x' = \gamma(x - vt) \end{cases}$

let photograph occur when  $t = 0$

so  $L_0 \sin \theta' = L \sin \theta$  (1) (y)

$L_0 \cos \theta' = \gamma L \cos \theta$  (2) (x)

It is important that this is for the  $S$  frame

so (1)/(2) :

$$\tan \theta = \gamma \tan \theta'$$

since  $\gamma \geq 1$  and  $0 < \theta < 90^\circ$

$$\frac{d}{d\theta} \tan \theta > 0 \Rightarrow \theta > \theta'$$

ie rod is tilted upwards in  $S$ .

$$①^2 + ②^2 : L^2 (\sin^2 \theta + \gamma^2 \cos^2 \theta) = L_0^2 (\sin^2 \theta' + \cos^2 \theta') = L_0^2$$

$$\therefore L = \frac{L_0}{\sqrt{\sin^2 \theta + \gamma^2 \cos^2 \theta}}$$

Now it is not clear whether  $L > L_0$  or  $L < L_0$   
 here since  $\sin^2 \theta$  and  $\cos^2 \theta < 1$

$$\therefore \text{Instead use } L_0^2 \cos^2 \theta' = \gamma^2 L^2 \cos^2 \theta$$

$$\therefore L_0^2 \sec^2 \theta = \gamma^2 L^2 \sec^2 \theta'$$

$$L_0^2 (1 + \tan^2 \theta) = \gamma^2 L^2 (1 + \tan^2 \theta')$$

$$\text{Now } \tan \theta = \gamma \tan \theta' \quad \text{so}$$

$$L_0^2 (1 + \gamma^2 \tan^2 \theta') = \gamma^2 L^2 (1 + \tan^2 \theta')$$

$$L = L_0 \sqrt{\frac{(1 + \gamma^2 \tan^2 \theta')}{\gamma^2 (1 + \tan^2 \theta')}}}$$

$$L = L_0 \sqrt{\frac{\frac{1}{\gamma^2} + \tan^2 \theta'}{1 + \tan^2 \theta'}}$$

$$\frac{1}{\gamma^2} = 1 - \frac{v^2}{c^2}$$

$$\text{So } L = L_0 \sqrt{\frac{1 + \tan^2 \theta' - \frac{v^2}{c^2}}{1 + \tan^2 \theta'}}$$

$$L = L_0 \sqrt{1 - \frac{v^2}{c^2} \cos^2 \theta'}$$

$$\leftarrow 1 + \tan^2 \theta' = \frac{1}{\cos^2 \theta'}$$

$$\text{Since } \frac{v^2 \cos^2 \theta'}{c^2} > 0$$

$$\Rightarrow \boxed{L < L_0} \text{ contracts}$$

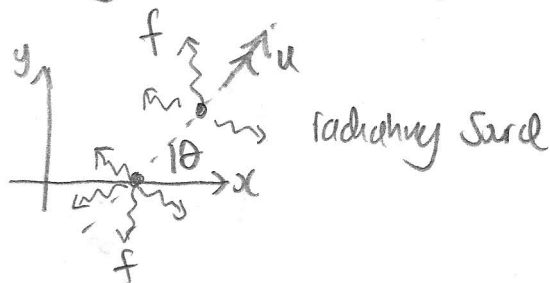
④

# 7, Relativistic Doppler Shift formula

$$f = \frac{f'}{\gamma \left(1 + \frac{u \cos \theta}{c}\right)}$$

$w =$  wave speed

$f' =$  frequency in rest frame of source.



Let  $w = c \quad \therefore \lambda f = c$   
 $\lambda' f' = c$

So  $\frac{c}{\lambda} = \frac{c}{\lambda' \gamma \left(1 + \frac{u \cos \theta}{c}\right)}$

$$\therefore \lambda = \lambda' \gamma \left(1 + \frac{u \cos \theta}{c}\right)$$

Now when  $\theta = 180^\circ$  (i.e. source moving towards observer at  $x=0, y=0$ )  $\lambda'_G = 500 \text{ nm}$ .

when  $\theta = 0^\circ$ ,  $\lambda'_R = 700 \text{ nm}$

For source (the relativistic basis),  $u = \frac{c}{10} \quad \therefore \gamma = (1 - 0.1^2)^{-\frac{1}{2}}$

So  $\lambda'_G = \lambda' \gamma (1 - u/c)$

$\lambda'_R = \lambda' \gamma (1 + u/c)$

$$\frac{\lambda'_R}{\lambda'_G} = \frac{1 + u/c}{1 - u/c} \Rightarrow (1 - u/c) \frac{\lambda'_R}{\lambda'_G} = 1 + u/c$$

$$\frac{u}{c} \left(1 + \frac{\lambda'_R}{\lambda'_G}\right) = \frac{\lambda'_R}{\lambda'_G} - 1 \quad \therefore u = c \left( \frac{\frac{\lambda'_R}{\lambda'_G} - 1}{\frac{\lambda'_R}{\lambda'_G} + 1} \right)$$

$$u = \frac{c (\lambda'_R - \lambda'_G)}{\lambda'_R + \lambda'_G}$$

$$\therefore u = \frac{200}{1200} c = \frac{c}{6}$$

$$u \approx 0.17c$$

$$\therefore \lambda' = \frac{\lambda_0}{\gamma(1 - v/c)}$$

$$u = \frac{1}{6}c$$

$$\gamma = \left(1 - \frac{1}{36}\right)^{-\frac{1}{2}} = \frac{6}{\sqrt{35}}$$

$$\therefore \lambda' = \frac{\lambda_0}{\frac{6}{\sqrt{35}} + \frac{5}{6}}$$

$$\lambda' = \lambda_0 \times \frac{\sqrt{35}}{5}$$

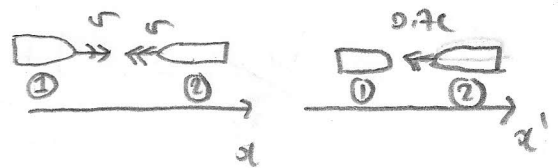
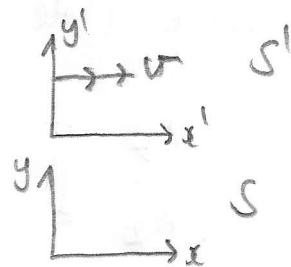
$$\lambda' = 500 \text{ nm} \times 1.183$$

$$= \boxed{592 \text{ nm}}$$

2/ let  $S'$  be rest frame of spaceship 1  
 $S$  be the Earth frame

Lorentz transform  
of velocity

$$u_x' = \frac{u_x - v}{1 - u_x v/c^2}$$



For brevity let all velocities be  $\parallel$   $x$  and  $x'$  directions

$$u_1' = 0 \quad u_2' = -0.7c$$

$$u_1 = v \quad u_2 = -v$$

$$-0.7c = \frac{-2v}{1 + v^2/c^2}$$

$$\left(1 + \frac{v^2}{c^2}\right) 0.7c = 2v$$

$$0.7 \frac{v^2}{c} - 2v + 0.7c = 0$$

$$0.7 \left(\frac{v}{c}\right)^2 - 2\left(\frac{v}{c}\right) + 0.7 = 0$$

$$v/c = \frac{2 \pm \sqrt{4 - 4(0.7)^2}}{1.4}$$

$$v/c = \frac{2 \pm 2\sqrt{1 - 0.7^2}}{1.4}$$

$$v/c = \frac{1}{0.7} (1 \pm \sqrt{1 - 0.7^2})$$

$$v/c = 0.408, 2.45$$

Now  $v \leq c$

So only solution is

$$v = 0.408c$$

in S frame

$$\begin{aligned} u_1 &= 0.408c \\ u_2 &= -0.408c \end{aligned}$$

9/

$$u_A = 0.5c$$

$$u_B = 0.8c$$

Define S' frame to be moving towards the Earth at  $v = 0.8c$ , its rest frame of B. Let  $x$  point towards the Earth.

$$\text{So } u_A' = \frac{u_A - v}{1 - u_A v / c^2} = \frac{0.5c - 0.8c}{1 - 0.5 \times 0.8 c^2 / c^2}$$

$$\begin{array}{c} \square \rightarrow u_A = 0.5c \\ A \end{array}$$

$$\begin{array}{c} \square \rightarrow u_B = 0.8c \\ B \end{array}$$

$$\rightarrow x$$

$$u_A' = \frac{-0.3c}{1 - 0.4}$$

$$u_A' = \frac{-0.3c}{0.6}$$

$$u_A' = \boxed{-c/2}$$

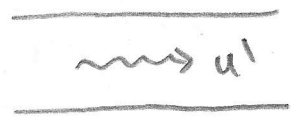
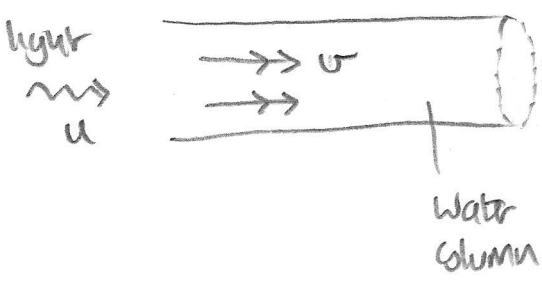
away from Earth relative to B.

$$\begin{array}{c} c/2 \\ \leftarrow \square A \end{array}$$

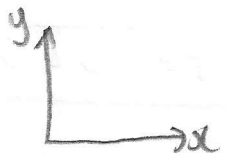
$$\square B$$

$$\rightarrow x'$$

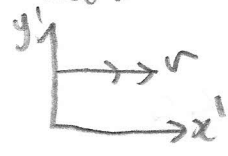
10/



lab frame S



S' frame moving with water



velocity transformations

$$u = \frac{u' + v}{1 + \frac{u'v}{c^2}}$$

$u' = \frac{c}{n}$  ( $n = \text{refractive index of water}$ )

$$\therefore u = \left( v + \frac{c}{n} \right) \left( 1 + \frac{vc}{nc^2} \right)^{-1}$$

$$u = \left( v + \frac{c}{n} \right) \left( 1 + \frac{v}{nc} \right)^{-1} \quad \text{Now } \frac{v}{nc} \ll 1$$

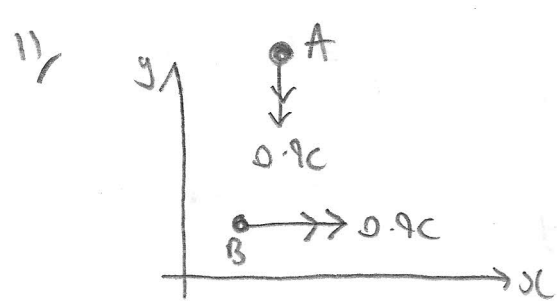
$$\therefore u = \left( v + \frac{c}{n} \right) \left( 1 - \frac{v}{nc} + \frac{(-1)(-2)}{2} \left( \frac{v}{nc} \right)^2 + \dots \right)$$

$$u = v + \frac{c}{n} - \frac{v^2}{nc} - \frac{v}{n^2} + \dots$$

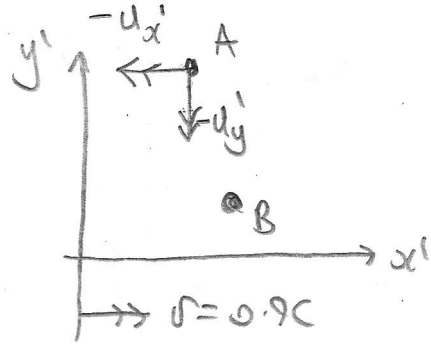
$$u = \frac{c}{n} + v \left( 1 - \frac{1}{n^2} \right) + O\left( \frac{v^2}{c} \right)$$

$\Rightarrow$   $u \approx \frac{c}{n} + v \left( 1 - \frac{1}{n^2} \right)$  as required.





S frame



S' frame  
(B is stationary)

velocity transforms

$$u_x' = \frac{u_x - v}{1 - \frac{u_x v}{c^2}}$$

$$u_y' = \frac{u_y}{\gamma \left(1 - \frac{u_x v}{c^2}\right)} \quad \gamma = \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}}$$

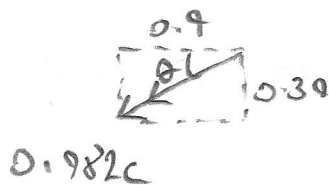
Now  $u_x = 0$      $u_y = -0.9c$      $v = 0.9c$

So  $-u_x' = -0.9c \Rightarrow \boxed{u_x' = 0.9c}$

$-u_y' = \frac{-0.9c}{(1 - 0.9^2)^{-\frac{1}{2}}} \Rightarrow \boxed{u_y' = 0.39c}$

So from B, A moves at velocity  $\sqrt{(u_x')^2 + (u_y')^2} = 0.982c$

at angle



$$\theta = \tan^{-1}\left(\frac{0.39}{0.9}\right)$$

$$= \boxed{23.6^\circ} \text{ from horizontal}$$

12

Relativistic momentum

$$\underline{p} = \gamma m \underline{u}$$

$$\gamma = \left(1 - \frac{u^2}{c^2}\right)^{-\frac{1}{2}}$$

$$u = |\underline{u}|$$

velocity

$$\underline{f} = \frac{d}{dt} (\gamma m \underline{u})$$

(Newton II: Force = rate of change of momentum)

$$\therefore \underline{f} = m \gamma \underline{a} + m \underline{u} \frac{d\gamma}{dt} \quad \text{if } m = \text{constant}$$

↑  
mass

$\underline{a}$  is acceleration  $\underline{a} = \frac{d\underline{u}}{dt}$

Now

$$\begin{aligned} \frac{d\gamma}{dt} &= \frac{d}{dt} \left( \left(1 - \frac{\underline{u} \cdot \underline{u}}{c^2}\right)^{-\frac{1}{2}} \right) \\ &= -\frac{1}{2} \left(1 - \frac{\underline{u} \cdot \underline{u}}{c^2}\right)^{-\frac{3}{2}} \left(-\frac{1}{c^2}\right) \frac{d}{dt} (\underline{u} \cdot \underline{u}) \\ &= \frac{1}{2c^2} \gamma^3 \left( \underline{u} \cdot \frac{d\underline{u}}{dt} + \frac{d\underline{u}}{dt} \cdot \underline{u} \right) \\ &= \frac{\gamma^3}{c^2} (\underline{u} \cdot \underline{a}) \end{aligned}$$

$$\underline{f} = m \gamma \underline{a} + m \frac{\gamma^3}{c^2} (\underline{u} \cdot \underline{a}) \underline{u} \quad (*)$$

Now when  $\underline{f} \parallel \underline{u}$ , it makes sense that  $\underline{a} \parallel \underline{u}$

So write (\*) if  $\underline{f} = F \hat{u}$

$$F = m \gamma \frac{du}{dt} + m \frac{\gamma^3}{c^2} u^2 \frac{du}{dt}$$

$$F = m \gamma^3 \frac{du}{dt} \left( \frac{1}{\gamma^2} + \frac{u^2}{c^2} \right)$$

(10)

$$\text{Now } \frac{1}{\gamma^2} = 1 - \frac{u^2}{c^2}$$

$$\therefore \frac{1}{\gamma^2} + \frac{u^2}{c^2} = 1$$

$$\therefore F = m \gamma^3 \frac{du}{dt}$$

$$\text{or } \boxed{F = m \left(1 - \frac{u^2}{c^2}\right)^{-3/2} \frac{du}{dt}}$$

If  $F$  is a constant and  $u=0$  when  $t=0$

$$\frac{Ft}{m} = \int_0^u \left(1 - \frac{u^2}{c^2}\right)^{-3/2} du = \int_0^u \gamma^3 du$$

$$\text{Now } \frac{d\gamma}{du} = -\frac{1}{2} \left(1 - \frac{u^2}{c^2}\right)^{-3/2} \left(-\frac{2u}{c^2}\right) = \gamma^3 \frac{u}{c^2}$$

$$\text{So } du = \frac{c^2}{u \gamma^3} d\gamma$$

$$\therefore \gamma^3 du = \frac{c^2}{u} d\gamma$$

$$\text{Now } u = c \sqrt{1 - \frac{1}{\gamma^2}} \quad (\text{+ve version})$$

$$u = \frac{c \sqrt{\gamma^2 - 1}}{\gamma}$$

$$\frac{c^2}{u} = \frac{\gamma c}{\sqrt{\gamma^2 - 1}}$$

$$\text{So } \frac{Ft}{m} = c \int_1^\gamma \frac{\gamma}{\sqrt{\gamma^2 - 1}} d\gamma$$

↑ when  $u=0$ ,  $\gamma=1$

$$\text{Now } \frac{d}{dt} \sqrt{\gamma^2 - 1} = \frac{1}{2} (\gamma^2 - 1)^{-\frac{1}{2}} \times 2\dot{\gamma}$$

$$= \frac{\dot{\gamma}}{\sqrt{\gamma^2 - 1}}$$

$$\text{So } \int_1^{\gamma} \frac{\dot{\gamma}}{\sqrt{\gamma^2 - 1}} dt = \left[ \sqrt{\gamma^2 - 1} \right]_1^{\gamma}$$

$$= \sqrt{\gamma^2 - 1}$$

$$\therefore \frac{Ft}{m} = c \sqrt{\gamma^2 - 1}$$

$$\therefore \left( \frac{Ft}{mc} \right)^2 = \gamma^2 - 1$$

$$\gamma^2 = 1 + \left( \frac{Ft}{mc} \right)^2$$

( $\gamma > 1$ )  
(so no -ve)

$$\left( 1 - \frac{u^2}{c^2} \right)^{-1} = \frac{m^2 c^2 + F^2 t^2}{m^2 c^2}$$

$$1 - \frac{u^2}{c^2} = \frac{m^2 c^2}{m^2 c^2 + F^2 t^2}$$

$$1 - \frac{m^2 c^2}{m^2 c^2 + F^2 t^2} = \frac{u^2}{c^2}$$

$$\sqrt{\frac{m^2 c^2 + F^2 t^2 - m^2 c^2}{m^2 c^2 + F^2 t^2}} \quad c = u$$

$$\therefore u = c \sqrt{\frac{F^2 t^2}{m^2 c^2 + F^2 t^2}} \Rightarrow \frac{u}{c} = \frac{Ft}{\sqrt{m^2 c^2 + F^2 t^2}}$$

$$\frac{u}{c} = \sqrt{\frac{1}{\left(\frac{mc}{Ft}\right)^2 + 1}}$$

$$\frac{u}{c} = \left(1 + \left(\frac{mc}{Ft}\right)^2\right)^{-\frac{1}{2}}$$

In classical limit  $u \ll c$   $F = m \frac{du}{dt}$

So  $\int_0^t \frac{F}{m} dt = \int_0^u du \Rightarrow \boxed{u = \frac{Ft}{m}}$

Now  $u \ll c \Rightarrow \frac{u}{c} \ll 1$  which means

$1 + \left(\frac{mc}{Ft}\right)^2 \gg 1$  So  $\frac{mc}{Ft} \gg 1$

$\therefore$  in classical limit  $\frac{u}{c} \approx \frac{Ft}{mc}$

$\Rightarrow \boxed{u \approx \frac{Ft}{m}}$

B/ If an electron gains kinetic energy  $k = 50 \text{ kV} \times 1.6 \times 10^{-19}$

Relativistic	$k = (\gamma - 1)mc^2$
Classical	$k = \frac{1}{2}mv^2$

$$\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}}$$

$m =$  mass of an electron  $= 9.109 \times 10^{-31} \text{ kg}$

(i) So  $\gamma - 1 = \frac{k}{mc^2}$   $\therefore \gamma = 1 + \frac{k}{mc^2}$

$$\left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}} = 1 + \frac{k}{mc^2} \Rightarrow 1 - \frac{v^2}{c^2} = \left(1 + \frac{k}{mc^2}\right)^{-2}$$

$$\frac{v^2}{c^2} = 1 - \left(1 + \frac{k}{mc^2}\right)^{-2}$$

$$v = c \sqrt{1 - \left(1 + \frac{k}{mc^2}\right)^{-2}} \quad (*)$$

Now in classical limit  $v \ll c$

$$\begin{aligned} \text{So } \gamma &= \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}} = 1 + \frac{v^2}{2c^2} + \dots \\ &\approx 1 + \frac{v^2}{2c^2} \end{aligned}$$

$$\therefore (\gamma - 1)mc^2 \approx \frac{v^2}{2c^2} mc^2 = \boxed{\frac{1}{2}mv^2} \quad \checkmark$$

$$\text{So if } k = \frac{1}{2}mv^2 \Rightarrow v = \sqrt{\frac{2k}{m}}$$

$$\therefore \frac{v}{c} = \sqrt{\frac{2k}{mc^2}}$$

$$\frac{v}{c} = \sqrt{\frac{2 \times 50 \times 10^{-3} \times 1.602 \times 10^{-19}}{9.109 \times 10^{-31} \times (2.998 \times 10^8)^2}} \approx \boxed{0.4423}$$

(classical expression)

Equivalent relativistic expression (\*) yields

$$\boxed{\frac{v}{c} = 0.4127}$$

ie the classical expression yields a speed which is about

$$\frac{0.4423 - 0.4127}{0.4127} \times 100 \approx \boxed{7\% \text{ higher}}$$

If the maximum distance travelled by an electron in the TV set  $\approx 0.5 \text{ m}$ . Difference in travel times  $- t_{\text{rel}} - t_{\text{classical}} = \Delta t$

$$\approx \Delta t \approx \frac{0.5}{c} \left( \frac{1}{0.4127} - \frac{1}{0.4423} \right)$$

$$\approx \boxed{0.27 \text{ ns}}$$

$$(1 \text{ ns} = 10^{-9} \text{ s})$$

If a TV scans at 50Hz and has the resolution of HD digital resolution i.e. 1920 x 1080 pixels

$$\text{Time between pixels is } \approx \Delta t = \frac{1/50}{1920 + 1080} = 9.6 \times 10^{-9} \text{ s}$$

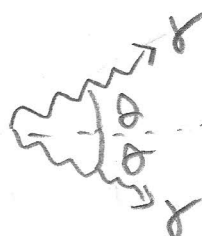
$$\text{i.e. } \approx \boxed{10 \text{ ns}}$$

So on this timescale the difference between the classical and relativistic travel times is less than an 'inter-pixel' distance by  $\uparrow \frac{1}{36}$  i.e. it is not likely to have a significant effect.

14

$$e^- \xrightarrow{u} \text{ (1 MeV) } \uparrow k$$

$$e^+ \longrightarrow$$



Each photon has energy  $E_\gamma$  and momentum  $p_\gamma$

Let both electron and positron have mass  $m = 9.109 \times 10^{-31} \text{ kg}$

$$\text{By conservation of energy } k + 2mc^2 = 2E_\gamma$$

$$\text{So } E_\gamma = \frac{k}{2} + mc^2 = 0.5 \text{ MeV} +$$

$$\frac{9.109 \times 10^{-31} + (2.998 \times 10^8)^2}{10^6 + 1.609 \times 10^{-19}}$$

$$= \boxed{1.0088 \text{ MeV}}$$

$$\uparrow 0.5088 \text{ MeV}$$

(15)

Now using Energy - Momentum

$$E^2 - p^2 c^2 = m^2 c^4$$

for the electron, we can find the total momentum before the collision (since the positron is at rest).

$$\text{So } p = \frac{\sqrt{E^2 - m^2 c^4}}{c}$$

Now  $E = mc^2 + K$  (total energy of electron)

$$\text{So } E^2 = m^2 c^4 + 2Kmc^2 + K^2$$

$$\therefore E^2 - m^2 c^4 = K(K + 2mc^2)$$

$$\therefore p = \frac{\sqrt{K} \sqrt{K + 2mc^2}}{c}$$

Now suitable units are MeV for energy and  $\frac{\text{MeV}}{c}$  for momentum

$$K = 1 \text{ MeV}$$

$$mc^2 = 0.5088 \text{ MeV}$$

$$\therefore p = \frac{\sqrt{1 + 0.5088 \times 2} \text{ MeV}}{c} = \boxed{1.420 \frac{\text{MeV}}{c}}$$

Now by conservation of momentum

$$p = 2 p_f \cos \theta$$

$$\text{Now } E_f = p_f c \quad \text{So}$$

$$\theta = \cos^{-1} \left( \frac{p}{2 p_f} \right)$$

$$\theta = \cos^{-1} \left( \frac{pc}{2E_f} \right)$$

$$\theta = \cos^{-1} \left( \frac{1.420}{1.0088} \right) = \boxed{45.3^\circ}$$

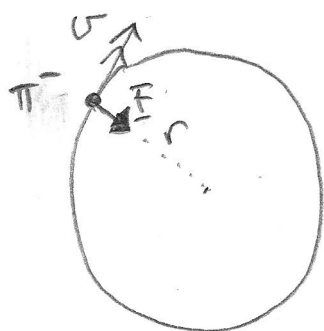


$$15/ \quad k^0 \rightarrow \pi^+ + \pi^-$$

$$m_\pi = 140 \text{ MeV}/c^2$$

$k^0$  is initially at rest.

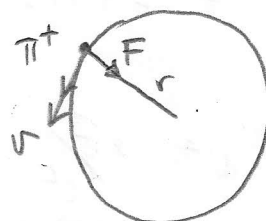
By conservation of energy  $m_k c^2 = 2m_\pi c^2 + 2E = 2E$   
 where  $E$  is the kinetic energy of each pion. (Assume the same!)



Force on pions due to magnetic field is

$$\underline{F} = z_\pi \underline{v} \times \underline{B}$$

$$\text{i.e. } |\underline{F}| = \boxed{F = z_\pi v B}$$



(X)  $\underline{B}$

Now 
$$\underline{F} = m\gamma \underline{a} + m\gamma^3 \left( \frac{\underline{a} \cdot \underline{v}}{c^2} \right) \underline{v} \quad (\text{Newton II})$$

For circular motion  $\underline{a} \cdot \underline{v} = 0$  so 
$$\underline{F} = m\gamma \underline{a}$$

Now 
$$|\underline{a}| = \frac{v^2}{r}$$

so 
$$z_\pi v B = m_\pi \gamma \frac{v^2}{r}$$

$$z_\pi r B = \gamma m_\pi v$$

$$\boxed{z_\pi r B = p_\pi}$$

Note  $z_{\pi^+} = e$

$z_{\pi^-} = -e$

$e = 1.609 \times 10^{-19} \text{ C}$

By Energy-momentum invariant for the pion

$$E^2 = m_\pi^2 c^4 + p_\pi^2 c^2$$

$$E = \sqrt{m_\pi^2 c^4 + e^2 r^2 B^2 c^2}$$

(17)

$$m_{\pi}^2 c^4 = (140 \text{ MeV})^2$$

$$e r B c = \frac{r B c}{10^6} \text{ MeV}$$

$$= \frac{0.344 \times 2 + 2.998 + 6^2}{10^6} \text{ MeV}$$

$$= 206.3 \text{ MeV}$$

$$\text{So } E = \sqrt{140^2 + 206.3^2} \text{ MeV}$$

$$= 249.129 \text{ MeV}$$

$$\therefore m_k = \frac{2E}{c^2}$$

$$= \boxed{498.6 \frac{\text{MeV}}{c^2}}$$

$$\text{From above } p_{\pi} = e r B = \boxed{206.3 \frac{\text{MeV}}{c}}$$

Now

$$p = \gamma m v$$

$$p^2 = \gamma^2 m^2 v^2$$

$$p^2 = \left(1 - \frac{v^2}{c^2}\right)^{-1} m^2 v^2$$

$$p^2 = \frac{c^2}{c^2 - v^2} m^2 v^2$$

$$p^2 (c^2 - v^2) = m^2 c^2 v^2$$

$$p^2 c^2 = v^2 (m^2 c^2 + p^2)$$

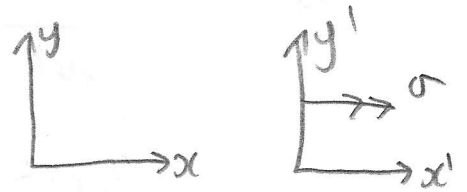
$$\therefore v^2 = \frac{p^2 c^2}{p^2 + m^2 c^2}$$

$$\boxed{v = \frac{p c}{\sqrt{p^2 + m^2 c^2}}}$$

$$So \quad v_{\pi} = \pm \frac{206.3 c}{\sqrt{206.3^2 + 140^2}} = \boxed{0.827c}$$

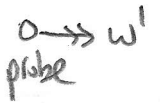
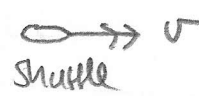
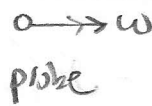
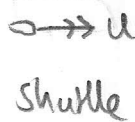
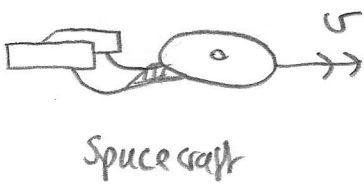
16/ Velocity transformation formula between S and S' frames

$$\boxed{dx = \frac{dx' + v}{1 + v dx'/c^2}}$$



Earth frame

Spacecraft frame

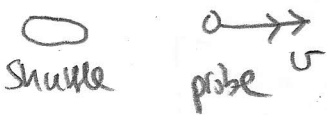


$$\boxed{u = \frac{2v}{1 + v^2/c^2}}$$

∴ speed of shuttle relative to Earth.

Shuttle frame

$$\boxed{w' = \frac{2v}{1 + v^2/c^2}}$$



So  $w = \frac{w' + v}{1 + w'v/c^2}$  [speed of probe relative to Earth]

$$\frac{w'v}{c^2} = \frac{2v^2}{c^2 + v^2} \quad \therefore 1 + \frac{w'v}{c^2} = \frac{c^2 + v^2 + 2v^2}{c^2 + v^2}$$

$$w = \left( \frac{2v}{1 + v^2/c^2} + v \right) \frac{c^2 + v^2}{3v^2 + c^2} = \frac{3v^2 + c^2}{c^2 + v^2}$$

$$\omega = \left( \frac{2vc^2}{c^2+v^2} + \frac{v(c^2+v^2)}{c^2+v^2} \right) \left( \frac{c^2+v^2}{3v^2+c^2} \right)$$

$$\omega = \frac{2vc^2 + vc^2 + v^3}{3v^2 + c^2}$$

$$\omega = \frac{3vc^2 + v^3}{3v^2 + c^2}$$

$$\omega = \frac{3v + \frac{v^3}{c^2}}{3\frac{v^2}{c^2} + 1}$$

$$\omega = v \left( \frac{3 + \frac{v^2}{c^2}}{3\frac{v^2}{c^2} + 1} \right)$$

Now in classical limit  $v \ll c$

$$\Rightarrow \boxed{\omega \rightarrow 3v}$$

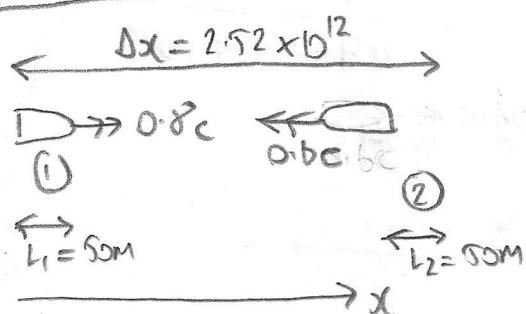
in Relativistic limit  $v \rightarrow c$ ,  $\boxed{\omega \rightarrow v}$

$$\left[ \text{Since } \frac{v^2}{c^2} \rightarrow 1, \omega \rightarrow v \left( \frac{3+1}{3+1} \right) = v \right]$$

$$\text{ie, } \boxed{\omega \rightarrow c}$$

17

Liz (on Earth)



- (i) Proper length of rocket 1 is  $(1 - 0.8^2)^{-\frac{1}{2}} \times 50 = \boxed{83.3m}$   
 (ie  $\delta L_1$ )  
 Proper length of rocket 2 is  $(1 - 0.6^2)^{-\frac{1}{2}} \times 50 = \boxed{62.5m}$

- (ii) If we take a reference frame where rocket ① is stationary, the observed velocity of rocket ② is

$$u_x' = \frac{u_x - v}{1 - \frac{u_x v}{c^2}}$$

$$u_x = -0.6c$$

$$v = 0.8c$$

ie  $u_x' = \frac{-0.6c - 0.8c}{1 + 0.6 \times 0.8}$

$$= \boxed{-0.946c}$$

- $\therefore$  rocket 2 will have length  $\frac{62.5}{(1 - 0.946^2)^{-\frac{1}{2}}} = \boxed{20.3m}$

If we take a reference frame where rocket ② is stationary, the observed velocity of rocket ① is  
 [using  $v = -0.6c$ ,  $u_x = 0.8c$ ]

$$u_x' = \frac{0.8c + 0.6c}{1 + 0.8 \times 0.6} = \boxed{0.946c}$$

- $\therefore$  rocket 1 will have length  $\frac{83.3}{(1 - 0.946^2)^{-\frac{1}{2}}} = \boxed{27.0 \text{ metres}}$

(21)

(iii) From Liz's perspective, relative velocity is  $1.4c$   
 $\therefore$  time to collision is  $t = \frac{2.52 \times 10^{12}}{1.4 + 2.998 \times 10^8}$

$$= 60045$$
$$\approx \boxed{100 \text{ mins}}$$

Since they are approaching each other

(iv) From Rocket 1's perspective, 100 mins in the Earth frame is  $\frac{100}{(1 - 0.8^2)^{-\frac{1}{2}}} = \boxed{60 \text{ mins}}$

From Rocket 2's perspective, 100 mins in the Earth frame is  $\frac{100}{(1 - 0.6^2)^{-\frac{1}{2}}} = \boxed{80 \text{ mins}}$

(v) If crew evacuation time is 50 min (in the rest frame of the rocket) this means an evacuation should be possible before the rockets collide.