

Neither Chalk and Talk or Death by PowerPoint

A case study of how to sensitively augment traditional Mathematics and Physics teaching with computer programming and online educational resources



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Winchester College, UK











I teach at Winchester College

• A UK Independent Boarding School

Niton

- Founded 1382
- 700 boys : 100 teachers

http://www.winchestercollege.org

3



Meads

Bethesda

Chapel & School

Flint court



B2, Flint Court. Max students = 22

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'Faces of Mathematics' →







In the classroom

Hardware & software

Desktop Windows PC linked to College network 🗸

- Network drives for students (Y) and staff (Z)
- Internet access
- SIMS

Double-width white board

Pull down projector screen Overhead projector linked to computer



What are my **requirements** for IT equipment? – **To enable me to teach more effectively** But what do I do in the classroom most often?

• Model what the students do, which is almost always using pen and paper Hence emphasis on a large whiteboard and not a tiny, hard-to-read 'Smartboard'

But 'inspiratia' (hard to draw figures, accurately plotted graphs, simulations etc) are *difficult (i.e. time consuming) to create without the use of IT*

Software

In the Winchester Mathematics department we typically use **Microsoft Word** (with **Mathtype** for equations) to prepare worksheets and exams

FxGraph and FxDraw are used for graphs and illustrations

We also use Autograph and Microsoft Excel for analysis and graph plotting

I tend to use **PowerPoint** for handouts as well as presentations

I also use **MATLAB** for simulations and **Xara Designer** for complex illustrations and website design

Autograph

Occasionally I will use Scientific Word (LaTeX) for mathematical documents

PD

I use **Adobe Acrobat** for PDF display and **IrfanView** for graphics display













Most of my IT usage is *not* 'cutting edge.' My students are typically **most engaged when they experience Mathematics being done** *live*, with a pen or pencil. i.e. *just like how I expect them to do it.*

There is *infinite creative freedom* with a pen, and no 'boot up time' !





However, there *are* many occasions when an IT based solution *is* preferable. **Examples** to follow

My Golden Rule for IT usage

Does it offer tangible **benefits** over simple, traditional writeon-a board teaching techniques, compared to the **cost** in terms of boot-up-time, reliance on network infrastructure etc?

Example 1: Projection of homework PDFs to facilitate feedback

"Middle Part" (Year 10, i.e. 14-15 year olds) Mathematics:

5 lessons + 3 x 30min homework per week

Homework problems, regularly bespoke and always varied, form the backbone of the Winchester Mathematics experience.

Homework feedback is therefore of paramount importance. To make this process *efficient*, I typically **display** PDF copies of my *handwritten* solutions using the projector.

Benefits

- I can face the class and discuss the problems with them
- Students annotate their work via direct visual comparison. It is what their homework ideally should look like.

Question 1

Consider the tangram Some angles 90-2 (1,2)Prove that triangles OAB and OED are similar S Smaller (i) (2.2) (ii) Find coordinates (a,b) (0,1)line OB IS (iii) Calculate the areas of: 4= 20((a) triangle OAB. line ED is (b) triangle OED $y = -\frac{1}{2}x + 1$ (c) quadrilateral EBCD They integed at $2x = -\frac{1}{2}x+1 \Rightarrow \frac{1}{2}x=1 \Rightarrow x=\frac{2}{5}$ $(a, 5) = |(\frac{2}{5}, \frac{4}{5})| \qquad AB = \sqrt{2^{2} + 1^{2}} = \sqrt{5} \quad OO = 2$ So Scale factor OED → OAB is 15 i area factor is If. Now creat of OAB is 2(1)(2) = 1 So area of OED is $\frac{4}{5} = \boxed{0.8}$ i. Area of EBCD = $\frac{4-1-0.8}{5}$ Question 2 Find the coordinates of stationary points of $y = 3x^2 + \frac{48}{x^2}$. Are they minima or maxima? $y = 3x^2 + 48x^{-2}$ Based upon your calculations, try to sketch the curve! $\frac{dy}{dx} = 6x - \frac{96}{3} \xrightarrow{dy} = 0 \Rightarrow \frac{96}{3} = 6x \Rightarrow \frac{96}{6} = x^{4}$ when x= t2 1 y= 3 (22) + 42 = 24 $16 = x^4 \Rightarrow \boxed{x=12}$ So Minima at (12, 24) y=48/22 24 -7 2

 Essence of lesson stored for future use. Minimum subsequent 10 preparation.

Example 2: Use PowerPoint when you really need to give a proper lecture!





VBk Practical Mathematics and Microsoft Excel Course





Mathmā



AF, CNB, MZ, APM

The ratio of Moon and Earth radii is approximately



$$\frac{R_{\oplus}}{R_{M}} \approx \frac{6353 \text{km}}{1737 \text{km}} = 3.66$$

By the construction on the left:

 $R_{M} = R_{\oplus} - \frac{1}{2}a$ $\frac{1}{2}a = R_{\oplus} \tan 36^{\circ}$ $\therefore R_{M} = R_{\oplus} - R_{\oplus} \tan 36^{\circ}$ $\frac{R_{\oplus}}{R_{M}} = \frac{1}{1 - \tan 36^{\circ}} \approx 3.66$

× 3.66





This amazing coincidence also shows that 'the Moon and the Earth square the circle'

 R_{M} $2R_{\oplus}$

The circumference of the Earth plus the circumference of the Moon is given by:

$$C = 2\pi R_{\oplus} + 2\pi R_M$$

$$C = 2\pi R_M \left(\frac{11}{3} + 1\right)$$

$$C = 2\pi R_M \left(\frac{11}{3} + \frac{3}{3}\right) = 2\pi R_M \times \frac{14}{3}$$

$$C = \frac{4 \times 7}{3} \pi R_M$$

The perimeter of a square bounding the Earth is $P = 4 \times 2R_{\oplus} = 4 \times \frac{22}{3} R_M$

$$\frac{P}{C} = \frac{4 \times \frac{22}{3} R_M}{\frac{4 \times 7}{3} \pi R_M} = \frac{22}{7} \times \frac{1}{\pi} = 1.000402...$$

i.e. *P* = *C* to a very good approximation!

Note this is perhaps where the popular approximation $\pi \approx \frac{22}{7}$ might have originated....

Example 3: What if? analysis using automated calculation



It is really hard to describe a 'mathematical landscape' if each elevation has to be painstakingly constructed! Once the basic technique is accepted, automation can enable exploration. In this case, a *bespoke* 'app' is often needed to *make your point clearly* Example 3: What if? analysis using automated calculation







Forced Simple Harmonic Equation solution explorer Andy French. November 2013.





harmonograph

- The Harmonograph was a Victorian curiosity attributed to Professor Blackburn in 1844
- Use two or three pendulums to create strange and beautiful patterns



Example of a *lateral* harmonograph



Photo from The Science Museum

Represent musical harmonies visually with the harmonograph!



F=2.01, D=0.7, A=1, phi=0

> Note the *difference* a small change in F makes....

> > 19

Gravity simulator using Excel

Binary system gravity simulation.		Initial condi	itions has	ed unon Kenleri	an orbits						Planet						
Di Andrew French, June 2013.		initial cond		cu upon kepien					ve	ars	AU	ms^-1	ms^-2	AU	ms^-1	ms^-2	
Mass of star1 (solar masses)	6	Planet							t		x0	vx0	ax0	y0	vy0	ay0	
Mass of star2 (solar masses)	2		AU	AU	AU	ms^-1	ms^-1		0		1.94845	0	-0.013175634	0	50057.51	0	
Mass of planet (Earth masses)	1	theta	r	x0	y0	vx0	vy0		0.0	000072	1.9484498	-23.9134	-0.007887959	0.00076	50057.5	-4.44018E-06	
		0	2.123	1.94845	0	0	50057.50806		0.0	000144	1.9484493	-41.8238	-0.007887961	0.00152	50057.49	-8.88036E-06	
Semi-major axis of initial star orbit /AU	1								0.0	000216	1.9484485	-59.7341	-0.007887963	0.002279	50057.46	-1.33205E-05	
Semi-major axis of initial planet																	
orbit /AU about star1	2.12345	Star 1							0.0	000288	1.9484475	-77.6444	-0.007887967	0.003039	50057.43	-1.77607E-05	
			AU	AU	AU	ms^-1	ms^-1	ms^-1	0.0	00036	1.9484461	-95.5548	-0.007887971	0.003799	50057.38	-2.22009E-05	
Orbital eccentricity of initial star																	
orbit	0.3	theta	r	x1	y1	vx1	vy1	v1	0.0	000432	1.9484446	-113.465	-0.007887975	0.004559	50057.33	-2.66411E-05	
Orbital eccentricity of initial planet																	
orbit about star 1	0	0	0.7	-0.175	0	0	-28696.05328	28696.05328	0.0	000504	1.9484427	-131.376	-0.007887981	0.005318	50057.26	-3.10813E-05	
_									0.0	000576	1.9484406	-149.286	-0.007887987	0.006078	50057.19	-3.55215E-05	
Initial polar angle /degrees of initial star orbit	0	Star 2							0.0	000648	1.9484382	-167.196	-0.007887993	0.006838	50057.1	-3.99616E-05	
Initial polar angle /degrees of																	
initial planet orbit about star 1	0		AU	AU	AU	ms^-1	ms^-1	ms^-1	0.0	00072	1.9484355	-185.107	-0.007888001	0.007598	50057	-4.44018E-05	
		theta	r	x1	y1	vx1	vy1	v2	0.0	000792	1.9484326	-203.017	-0.007888009	0.008357	50056.9	-4.8842E-05	
timestep /years	0.000072	0	0.7	0.525	0	0	86088.15985	86088.15985	0.0	000864	1.9484293	-220.928	-0.007888018	0.009117	50056.78	-5.32822E-05	_
									0.0	000936	1.9484258	-238.838	-0.007888028	0.009877	50056.66	-5.77224E-05	
									0.0	001008	1.9484221	-256.749	-0.007888038	0.010637	50056.52	-6.21625E-05	
theta0 (stars) /radians	0			2					0.0	00108	1.9484181	-274.659	-0.007888049	0.011396	50056.37	-6.66027E-05	_
theta0 (planet) /radians	0	_							0.0	001152	1.9484137	-292.57	-0.007888061	0.012156	50056.22	-7.10429E-05	
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delta_t /s	2270.592			1.5					0.0	001296	1.9484043	-328.391	-0.007888086	0.013676	50055.87	-7.99233E-05	-
mass 0 (planet) /kg	5.9/E+24	_							0.0	001368	1.9483992	-346.302	-0.0078881	0.014435	50055.69	-8.43634E-05	
mass 1 (star1) /kg	1.1934E+31	_							0.0	00144	1.9483938	-364.212	-0.007888115	0.015195	50055.49	-8.88036E-05	-
mass 2 (star2) /Kg	5.976E+5U								0.0	001512	1.9463881	-382.123	-0.00788813	0.015955	50055.29	-9.52458E-05	-
Mass of Sup /kg	1.0205+20	-							0.0	001584	1.9483822	-400.034	-0.007888146	0.017474	50055.07	-9./084E-05	-
Mass of Farth /kg	5.07F+24	-							ar1 0.0	001030	1.940370	-417.944	-0.007000103	0.01/4/4	50054.84	-0.000102124	-
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									0.0	002376	1.948299	-597.055	-0.00788837	0.025071	50052.02	-0.000146526	-
					x/AU				0.0	002448	1.9482898	-614.966	-0.007888394	0.025831	50051.68	-0.000150966	
									0.0	00252	1.9482803	-632.877	-0.007888419	0.026591	50051.33	-0.000155406	
									0.0	002592	1.9482706	-650.789	-0.007888445	0.02735	50050.98	-0.000159846	
									0.0	002664	1.9482606	-668.7	-0.007888472	0.02811	50050.61	-0.000164286	
										000706	4.0400500	COC C12	0.007000.000	0.00007	50050.00	0.000460706	



Initial polar angle /degrees of initial planet orbit about star 1

timestep /years

0

0.0001

i.e. the motion of a planet within the mutual orbits of a binary star system

Example 4: Accurate plotting of graphs and geometry and their inter-relationships



Note: Learning to curve sketch a mathematical function *by hand* is one of the best methods for becoming adept at the subject and appreciating its interconnectedness. I *only* use IT for graph plotting when accuracy, or rapid variation, is required.

Example 4: Accurate plotting of graphs and geometry and their inter-relationships





Example 5: Confirming that you are correct!

Accurate drawing of *normals* to scale is often tricky. A tool can help to confirm whether your sketch is correct, and build confidence and precision into a lesson!



Example 5: Confirming that you are correct!

It is important that students can obtain satisfaction that an extended calculation is correct and also, *why* the answer is what it is. It is *not* always expedient to go through every question from scratch. An informative *automated* display of a solution can solve this problem.



Example 6: Using a tool to generate the 'right' numbers to make your point



Half Marathon Statistics	
NAME: Dr. French	HOUSE:

Table 1 is a frequency table corresponding to athletes who participated in the inaugural Southampton Half Marathon (13.1 miles or 21.1km) in April 2015.

DATE:

Question 1

Question 2

x < 1.3

 $1.3 \le x < 1.4$

 $1.4 \le x < 1.5$

 $1.5 \le x < 1.6$

 $1.6 \le x < 1.7$

 $1.7 \le x < 1.8$

 $1.8 \le x < 1.9$

 $1.9 \le x < 2.0$

 $2.0 \le x < 2.1$

 $2.1 \le x < 2.2$

 $2.2 \le x < 2.4$

 $2.4 \le x < 2.6$

 $2.6 \le x < 2.8$

 $2.8 \le x < 3.0$

Ouestion 3

4258

 $x \ge 3.0$

frequency density =

Time range /hours



A tool which generates statistical data can facilitate the creation of *interesting* and realistic resources. This would normally be very time consuming!



$$\overline{E} \propto \frac{1}{4258} \left(\begin{array}{c} 0.65 \times 10 + 1.35 \times 33 + 1.45 \times 10 + 1.55 \times 234 + 1.65 \times 384 + \dots \\ + 1.75 \times 533 + 1.47 \times 537 + 1.95 \times 521 + 2.05 \times 462 + \dots \\ + 2.15 \times 401 + 2.3 \times 543 + 2.15 \times 210 + 2.9 \times 127 + \dots \\ + 2.9 \times 47 + 3.25 \times 26 \end{array} \right)$$

$$= 8495.185 = 11.995 \text{ Mays}$$
Page 1 of 3



28

Example 7: (Physics) Live visual analysis of experimental data



File: E:\AndyFrench\Documents\AF\Programming\MATLAB\Apps with GUIs\SoundAnalyser\SoundAnalyser\Ringing bell.mp3 loaded in 1.9804 seconds.

Example 8: When you really need an animated demo!

Pentagon

How to construct a Pentagon using just a compass and a straightedge







Programming

Thursdays. 1600-1745. Mill or Science School IT suites Keyboard shortcuts:Q - QuitT - Load text fileI - Load image fileM - Load music fileP - PlayR - Reverse

B - Restart music

>

XXXX GUI template

>

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Restart

Load a text file and display in a list box	
A comedy of errors.txt	Load a text file Open 14049
The Comedy of Errors by William Shakespeare	
ACTI	
SCENELA hall in DUKE SOLINUS'S palace	
Load an image file, drag and zoom	
Load an image file Open	
Sybil.jpg	
Choose a colour scheme	
Normal full colour	
Load and play a .wav or .mp3 file	
Sound snip.mp3	Load an MP3 or .wav file Open

File: E:\AndyFrench\Documents\AF\Programming\MATLAB\Apps with GUIs\GUI template\XXXX\Sound snip.mp3 loaded in 1.7405 seconds.

Record sound for five seconds and then play it back

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229		
230	<pre>%List box which displays the loaded text</pre>	
231	function LISTtext Callback(hObject, eventdata, handles)	
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240	%%% IMAGE %%%	_
241		
242	<pre>% Load image file (e.g .jpg, .png etc) pushbutton</pre>	
243	<pre>- function PUSHloadimage_Callback(hObject, eventdata, handles)</pre>	_
244 -	global d	
245		_
246	<pre>%Bring up a windows file browser</pre>	
247 -	<pre>[filename, pathname] = uigetfile({'*.jpg;*.png;*.bmp'}, 'Choose image file to load');</pre>	
248		
249	Sonly proceed if file selected	
250 -	if filename ~=0	
251		
252	<pre>spelete any previous image delete(d image image handle);</pre>	
253 -	derete(d.image.image_nandre);	
255	SStore the filename in the GUI data	= -
256 -	d.image.filename = [pathname.filename]:	
257	crimage.fifename [pasimame,fifename],	
258	*Update the edit box with the filename only	
259 -	<pre>set(d.h.EDITloadimage, 'string', filename);</pre>	
260		
261	<pre>%Load the image</pre>	
262 -	d.image.image = flipdim(imread(d.image.filename), 1);	
263		_
264	&Apply desired basic image procesing	
265 -	RGB_output = basic_image_processing(d.image.image , d.image.colour_scheme);	





What is the difference between these images? The one on the left *also contains the entire works of Shakespeare.*

Find out how to do this, and much more, at **ProgSoc.** Mill, Thursdays 1600-1800.









Welcome to Spherium. Dragging the mouse in the main axes will result in a 3D rotation. Use the + and - buttons to zoom in and out, and the >, < etc to translate the figure. Hydrogenic orbital spheria take the form H XN e.g. H P1. Note the blue square must be pressed to update Spherium following 3D rotation.

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Geometry

Mountaineering

Areas of basic shapes Arcs, sectors, radians Argand diagram complex loci Cartesian equation of a circle Circle theorems Coordinate transformations Lines and angles Loci & constructions Lorentz transform Polygons Symmetry Transformations using matrices: Intro Transformations using matrices: Rotations Transformations using matrices: Not about the origin Transformations using matrices: Invariant lines Vectors Vector equations of lines and planes Volumes of basic solids

Music

Philosophy

Geometric transformations using matrices IGCSE trigonometry proofs First 86 Pythagorean triples Special triangles Proof of circle theorems Basic geometry notes Chimborazo

Maths is Fun: Geometric Construction animations



Circle theorems There are five main circle theorems, which relate to triangles or quadrilaterals drawn inside the circumference of a circle.



since they both meet at the origin of the circle, and therefore two edges of each triangle are circle radii.

Special Relativity is a theory of dynamics proposed by Albert Einstein in 1905. The key mathematical element is the use of the Lorentz Transform. This extends the equations of Galilean Relativity, which relate the Cartesian x, y, z coordinates of an object to coordinates of the same object as viewed in a frame of reference moving at velocity V in the positive x direction relative to the x,y,z system. Let S denote the x,y,z coordinate system and S' denote the x',y',z' coordinates of the moving frame.

The Lorentz transform incorporates the strange (but seeming true!) fact that the speed of light is the same for both S and S' frames. In other words, if a torch is shone from frame S, the speed of the light observed by S' would be the same speed as in S, and not the speed of light minus V.

The consequence of this effect is profound. It results in length contraction, time dilation and time synchronisation changes between the S and S' frames.

Galilean relativity



Consider the following candidates for the Lorentz transform of the spatial coordinates between the S and S' frames:

> $x = \gamma (x' + Vt')$ $x' = \gamma (x - Vt)$ y = y'z = z'

 γ is a function of V. In order to be consistent with Galilean relativity, it must be *unity* when $V \ll c$

Hence:

$$\begin{aligned} x &= \gamma \left(x' + Vt' \right) & x' &= \gamma \left(x - Vt \right) \\ \frac{x}{\gamma} &= x' + Vt' & \frac{x'}{\gamma} &= x - Vt \\ t' &= \frac{x}{\gamma V} - \frac{x'}{V} & t &= \frac{x}{V} - \frac{x'}{\gamma V} \\ t' &= \frac{x}{\gamma V} - \frac{\gamma \left(x - Vt \right)}{V} & t &= \frac{\gamma \left(x' + Vt' \right)}{V} - \frac{x'}{\gamma V} \\ \therefore & t' &= \gamma \left(t - \frac{x}{V} \left(1 - \frac{1}{\gamma^2} \right) \right) & \therefore & t &= \gamma \left(t' + \frac{x'}{V} \left(1 - \frac{1}{\gamma^2} \right) \right) \end{aligned}$$

Galilean relativity appears to work just fine in normal scenarios on Earth, i.e. when $V \ll c$ where the speed of light $c = 2.998 \times 10^8 \, {\rm ms}^{-1}$ The effects of Special relativity are only significant when V is close to c.

> Now consider a spherical light pulse emitted when x' = x. Since it radiates out at speed c in both S and S' from their (respective) origins, we can compare the radii r,r' of the pulse as observed from S and S'

$$r'^{2} = c^{2}t'^{2} = x'^{2} + y'^{2} + z'^{2}$$
$$r^{2} = c^{2}t^{2} = x^{2} + y^{2} + z^{2}$$

н

Since
$$y = y'$$
, $z = z'$ this means $c^2 t'^2 - x'^2 = c^2 t^2 - x^2$

Now when
$$x' = 0$$
, $x = Vt$
Hence $c^{2}t'^{2} = c^{2}t^{2} - V^{2}t^{2}$
 $\Rightarrow t' = t\sqrt{1 - \frac{V^{2}}{c^{2}}}$
 $\therefore 1 - \frac{1}{\gamma^{2}} = 1 - 1 - \frac{V^{2}}{c^{2}} = \frac{V^{2}}{c^{2}}$
 $\therefore \frac{1}{V} \left(1 - \frac{1}{\gamma^{2}}\right) = \frac{V}{c^{2}}$
The Lorentz Transform is now revealed!
 $x = \gamma(x'+Vt')$
 $y = y'$
 $z = z'$
 $t = \gamma\left(t' + \frac{Vx}{c^{2}}\right)$
 $x' = \gamma\left(t - \frac{Vx}{c^{2}}\right)$
So lengths
contract and time
dilates and shifts
when V becomes
close to c

Mathematics topic handout: Mechanics - Special Relativity Dr Andrew French. www.eclecticon.info PAGE 1

Preparation

I make extensive use of an electronic filing system for all my resources

I prepare weekly and copy the required files to the College network drive and my tablet



This is very useful for my teaching notes, and for helping students during workshop-style lessons





Assessment & Admin

An Excel spreadsheet for continuous, quantitative assessment A simple colour coding system is useful when briefing parents

	Y	N	N	N	N	N	Y	N	N	N	N	N N	N	N	1	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	N	Y	Y
W10 MP Topics exam	W10 MP Algebra exam	М115Н TT#1	W115H TT#2	W12SH П#1	W14C0M TT#1 - Proof stuff	W15COM TT#1	Circle Theorems test	W15C0M TT#2	W15COM TT#3	W16C0M TT#1	W16COM TT#2	W16C0M TT#3	Intermediates Jan 2015	W17C0M TT#1	W17C0M TT#2	W17C0M TT#3	W18C0M TT#1	W18COM TT#2	W18COM TT#3	W19COM TT#1	W19C0M TT#2	W20C0M TT#1	W20COM TT#2	W20COM TT#3	W21COM TT#1	W22C0M TT#1	W22C0M TT#2	W22COM TT#3	W23C0M TT#1	W23COM TT#2	W24C0M TT#1	W24C0M TT#2	W 25 COM TT#1	W25COM TT#2	W25C0M TT#3	W 26 COM TT# 1	W26C0M TT#2	W26C0M TT#3	Intermediates May 2015	W28CT TT#2	MP Summer Exam 1	MP Summer Exam 1
<mark>15%</mark>	65%	3	3	2.5	2		45%	2	2	2	2.5	33	.0	3	3	3	2.5	3	2	3	3	2.5	1	3	3		2.5	2.0	2.0	2.5	4	4	4	3	3	4	2.5		3	3	53%	55%
58%	43%	3.5	3		2.5	3	100%	1.5	3	3	2.5	33	.0		3	3.5	3.5	2.5	2	2.5	2.5	2.5	2	2.5	5	3	3	3.0	2.0		3.5	3	3.5	2	2.5	3	4		3		50%	65%
55%	35%	2	2.5	2	3	1.5	100%	1	2	2	1	2 2	.0	2.5	2	3	1.5	2.5	1	2	2	1	1	2	1	2	2.5	1.0	1.5	1.0	1	1	1	2	2.5	2	1.5	2	1	2	83%	58%
<mark>18%</mark>	85%	3	4	2.5	2.5	3	83%	3	3	4	4	33	.0	4	3	2.5	3	4	2.5	Play	3	3.5	2.5	3		3	2	2.0	2.0	2.0	3	4	4	3.5	3	3	3	3	3	3	68%	35%
50%	65%	2	3	3.5	3	4	100%	2.5	3	3	4	33	.0	3	3	3	4	3	3	2.5	2	3	4	3	3.5	3	2.5	3.0	2.0	2.0	2.5	3	4	2.5	4.5	4.5	2.5	3	3	3	55%	45%
50%	83%	Ш	Ш	Ш	2.5	2	100%	2.5	3	3	2	22	.0	3	3	3.5	3.5	2	2	2	3	ill	ill	ill	2	4	4	ill	ill	4.0	2	2.5	3	2.5		2.5	3	4	2	3.5	65%	75%
58%	68%	2.5	3	3	3	2	100%	2	1	3	3	23	.0	3	2	2.5	2.5	2.5	1.5	4	2	1	2.5	2.5	3.5	2	2.5	3.0	2.0	3.0	2.5	2.5	2.5	2.5	2	3	3	2	2	3	68%	58%
70%	58%	2	2	2	2		100%	1	2	4		2 2	.0	4	1.5	3	2	3.5	1.5	3	1	1	1	2.5	2	1.5		2.0	2.0	1.5	2	2	2	2	3	3	1	2	2	2	63%	63%
53%	70%	2.5	3	3.5	2.5	3	100%	2	2	3	2.5	33	.0	1.5	2.5	2	3.5	4		3	3	2	2	2.5	2	3		3.5	3.0	5.0	Ш	3	3	2	2.5	2	2.5	2	2	3	75%	78%
55%	55%	2.5	2.5	2	2	2.5	100%	2.5	2	2	1	2 2	.0	2.5	2.5	2.5	2	2	2	3	3	3	2.5	2.5	2	2.5	2	2.5	2.0	2.0	3	3	3	2	2	3	2.5	2.5	2	3	58%	63%
50%	43%	2.5	4.5		2		83%	2	3			23	.0	3.5	4	3	3.5	2.5	4	2.5	3	4	3	3	4	3.5		1.5	3.5	2.0	4.5	2.5	2.5	2.5	3	4	3	2	4	3	68%	53%
38%	50%	3	4	3	2.5	3	94%	2.5	2	2	3	2 2	.0	3	3	4	5	4	3	4	4	3	3	3	3.5	3.5	3.5	5.5	3.5		4	3	3	3.5	4	4.5	3.5	4	3	4	65%	43%
55%	93%	2.5	2.5	1.5	2	2	100%	2.5	2	3	2	2 2	.0	2	2	2	3	3	2	3	2.5	2.5	3	2.5	2	2	1.5	2.5	2.0	2.0	2	3.5	2.5	3	3	3	2.5	2.5	3	2	75%	53%
53%	55%	1.5	2.5	3	3	4	78%	2	Ш		2.5	33	.0	3.5	2.5	2.5	2.5	2.5	2	3	2	3	ill	ill	2	ill	ill	3.0	2.0	2.0	3	2.5	away	away	2.5	2.5	3	2.5	3	2	70%	68%
53%	28%	2.5	2.5	2	2		89%	1	3	3	2	3	.0	2.5	2.5	2	2	2.5	1.5	2.5	2.5	1.5	1	2	2.5	3	3	2.5	2.0	2.0	2	3	2.5	2.5	5	4.5	4	2.5	3	3	65%	60%
50%	48%	1.5	2	3	3	3	94%	1	2	3	1.5	2 2	.0	2	2.5	2.5	2	2	1.5	III	Ш	1	1.5	2		3	2		2.0	3.0	3	3	2	2	2	2.5	2	1.5	2	2.5	68%	65%
55%	55%	2	2.5	2	4	2.5	94%	1.5	3	2	2	2 2	.0	3	2.5	3.5	2.5	3	2	III	Ш	2	2	3	Ш	2.5	3.5	2.0	2.0	away	2	2	4	3	2	2.5	2	3	2	2.5	50%	68%
73%	85%	1.5	2	1.5	2	1	100%	1	2	2	1	2 1	.0	1.5	1	2	2	1.5	1	2.5	2.5	1	1	2	1	1.5	2	1.0	1.5	2.0	1	1	1.5	2	2	2	1.5	2	1	1.5	65%	50%
55%	63%	2.5	2.5	3.5	2	3.5	78%	1	3	2	3	33	.0	2.5	2	3.5	2.5	3	3	111	Ш	3	2.5	3	2	3	3	1.5	2.0	3.0	2.5	2.5	3	3.5	3	4	3	3	3	3	40%	50%
55%	38%	2	3	3	3	3.5	100%	2	2	3	lost!	2 2	.0	3	2.5	2.5	3	3	2.5	3	3	3	3	3.5	4	3	3.5	2.0	2.0	1.5	2.5	3	2.5	3.5	3	3	3	3	3	3.5	55%	45%
55%	40%	Ш		3	3	3	100%	2.5	2	4	3.5	4	.0		2.5	4	3	4	2	4	3		2	2					3.0	4.0	4	3.5	4.5	2.5	4.5	4.5	4.5	5	4	3	65%	55%
53%	68%	1.5	2	2.5	3	2.5	100%	1.5	3	2	2	2 2	.0	3	2	2	2.5	2.5	2	3.5	2	2	2	3	3	3	3	2.5	2.0	2.0	3	2.5	3	2.5	3	2.5	2	3	3		58%	55%
57%	59%						93%																																		63%	57%

Microsoft Word documents for School Reports

Neither Chalk and Talk or Death by

PowerPoint

A case study of how to sensitively augment traditional Mathematics and Physics teaching with computer programming and online educational resources

Teacher led, largely 'traditional' lessons can be very effective!

IT can often be a distraction. The cost often outweighs the benefit Model good Mathematics by doing it live, with a pen

I don't like Smartboards! Keep it simple. Use a projector *if needed*, but otherwise write on a huge whiteboard!

IT is perhaps *most useful for teachers, not students* in terms of resource creation, storage and dissemination, and assessment and administration

Bespoke apps / computer models can offer an enriched Mathematical experience that cannot be easily achieved without IT

> Online resources can help augment learning, but inspiration starts in the classroom, *from* the teacher



Sybil's

Stop the madness for constant group work. Just stop it. And I want to be clear about what I'm saying, because I deeply believe our offices should be encouraging casual, chatty cafe-style types of interactions you know, the kind where people come together and serendipitously have an exchange of ideas. That is great. It's great for introverts and it's great for extroverts.

But we need much more privacy and much more freedom and much more autonomy at WOrk. school, same thing. We need to be teaching kids to work together, for sure, but we also need to be **teaching them how to work on their own.** This is especially important for extroverted children too. They need to work on their own because that is where deep thought comes from in part.



TED talk "The Power of Introverts"

Susan Cain Author of *Quiet*



